"No-hair" and uniqueness results for analogue black holes

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Outline



2 Uniqueness results for black-hole flows



Introduction

Hawking radiation

[Hawking 75, Unruh 76]

- semiclassical approximation for quantum fields in a black hole space-time \to vacuum for an infalling observer \neq vacuum for an observer at infinity
 - \rightarrow production of pairs of particles
- thermal radiation at $T = T_H = \kappa/(2\pi)$

Is HR observable?

- solar-mass black hole $\Rightarrow T_H \approx 6 \times 10^{-8} \text{K}$
- $\bullet\,$ cosmic microwave background temperature $\approx 2.7 {\rm K}$

 \Rightarrow unlikely to be observable

Analogue Gravity

Main idea [Unruh 81]

- sound waves in a non-relativistic fluid at rest
 - \Rightarrow massless excitations $\partial_t^2 \phi \Delta \phi = 0$

non-uniform flow

- $\rightarrow \Box \phi = 0$ in a curved space-time metric (analogue metric)
- \rightarrow analogue Killing horizon \Leftrightarrow transonic flow

Why analogue gravity?

- observe the analogue version of the Hawking effect
- explore possible phenomenologies from quantum gravity effects (e.g., dispersion) [Jacobson 91, Unruh 95]
- understand condensed matter phenomena through the analogy [Finazzi et al. 15]

Non-linear effects

- $\bullet\,$ nonlinear/backreaction terms different from those of QFT+GR $\rightarrow\,$ the correspondence breaks down
- however,
 - \bullet possible similarities \rightarrow how far can the analogy extend?

 \bullet interesting phenomena from a condensed matter/cold atoms point of view

 most studied example: "black-hole laser effect" [Corley and Jacobson 99, Leonhardt and Philbin 08]
 experimental realisation: [Steinhauer 14]

Black-hole laser

- $\bullet\,$ linear order: dynamical instability $\rightarrow\,$ exponential growth of unstable modes
- non-linear behavior: variety of possible outcomes [Michel and Parentani 13, Michel and Parentani 15, de Nova et al. 15]
 - saturation on a stable (non-lasing) stationary solution
 - periodic emission of superposed soliton trains
 - seemingly aperiodic emission of solitons
- behavior of black-hole and white-hole flows separately currently unclear

Uniqueness results for black-hole flows

BH flows in BEC: setup

model and assumptions

- (1+1)D Bose-Einstein (quasi-)condensate (BEC)
- dilute and weakly interacting
- ${\scriptstyle \bullet}\,$ correlation length \gg other length scales

repulsive two-body interactions

 \Rightarrow Gross-Pitaevskii equation (GPE):

$$i\partial_t \psi = -\frac{1}{2}\partial_x^2 \psi + V(x)\psi + a^2(x) \left|\psi\right|^2 \psi$$

V(x): (stationary) external potential $a^2(x) > 0$: two-body interactions

BH flows in BEC: setup

Toy-model for analytical calculations

•
$$V(x) = \theta(-x)V_{-} + \theta(x)V_{+}$$

•
$$a(x) = \theta(-x)a_- + \theta(x)a_+$$

•
$$(a_+^2 - a_-^2)(V_+ - V_-) < 0$$

 \Rightarrow solutions with a homogeneous density

$$\rho_0 = \frac{V_+ - V_-}{a_-^2 - a_+^2}$$

Stationary equation

general stationary configuration:

$$\psi(x,t) = e^{-i\omega t} \sqrt{\rho(x)} e^{i\int^x u(y)dy}$$

- $\omega:$ angular frequency
- ρ : atomic density
- u: local velocity

conservation of the number of atoms $\Rightarrow \rho(x) u(x) = J$

resulting equation on ρ :

$$\frac{\partial_x^2 \sqrt{\rho}}{\sqrt{\rho}} = 2\left(V - \omega\right) + 2a^2 \rho + \frac{J^2}{\rho^2}$$

Solutions in homogeneous backgrounds (1)

Homogeneous a and $V \Rightarrow$ integration over x

$$(\partial_x \rho)^2 = 4a^2 (\rho(x) - \rho_1) (\rho(x) - \rho_2) (\rho(x) - \rho_3),$$

 $(\rho_1,\rho_2,\rho_3)\in\mathbb{C}^3$ such that

$$\begin{cases} \rho_1 + \rho_2 + \rho_3 = 2(\omega - V)/a^2 \\ \rho_1 \rho_2 \rho_3 = J^2/a^2 \end{cases}$$

bounded solutions require

$$J^2 \le J_{\max}^2 = \frac{8(\omega - V)^2}{27a^4}$$

and $(\rho_1, \rho_2, \rho_3) \in \mathbb{R}^3$

Solutions in homogeneous backgrounds (2)



- $\rho_1 \leq \rho_2 \leq \rho_3 \rightarrow \text{solution}$ oscillates between ρ_1 and ρ_2
- $\rho_2 \approx \rho_1 \Rightarrow$ small-amplitude oscillations over a supersonic flow
- $\rho_2 \approx \rho_3 \Rightarrow$ soliton train over a subsonic flow

Characterization of BH flows

Conditions on BH flows

- continuity of ρ and ρ'
- asymptotically homogeneous
- transonic flow

• Mach number increases along the direction of the flow

6 degrees of freedom – (1+2) asymptotic conditions – 2 matching condition \Rightarrow 1-parameter families of solutions, relation $\omega(J)$

In fact, only one solution at fixed J

BH flows: two types of solutions

Homogeneous solution

- $\rho_{3,-} = \rho_{2,-} = \rho_{2,+} = \rho_{1,+}$
- $\rho(x) = \rho_0$
- transonic iff $a_{-}\rho_{0}^{3/2} > |J| > a_{+}\rho_{0}^{3/2}$



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Waterfall solution

$$\rho_{3,-} = \rho_{2,-}, \\ \rho_{2,+} = \rho_{1,+} = \rho_{1,-},$$

 half-soliton matched with a homogeneous solution

• requires
$$|J| > a_- \rho_0^{3/2}$$



"No-hair" and uniqueness results for analogue black holes

Smooth potentials: numerical results

Potentials of the form

$$f(x) = \begin{cases} f_{-} & x < -\sigma_f/2\\ \frac{f_{+}+f_{-}}{2} + \frac{f_{+}-f_{-}}{\sigma_f}x & -\frac{\sigma_f}{2} \le x \le \frac{\sigma_f}{2}\\ f_{+} & x > \sigma_f/2 \end{cases},$$

 $f\in\left\{a^2,V\right\}$ "hairy" BH flows found for $\sigma_{a^2}>\sigma_V,$ but seem to be unstable



Introduction	Linear perturbations
Uniqueness results for black-hole flows	Whitham's modulation theory in a nutshell
Stability analysis	Application to our model
Conclusion and outlook	Numerical results

Stability analysis

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The Bogoliubov-deGennes equation

Perturbations in $\psi(x,t) = \psi_0(x) (1 + \phi(x,t))$ \rightarrow BdG equation:

$$i\partial_t \phi = -\frac{1}{2}\partial_x^2 \phi - \frac{\partial_x \psi_0}{\psi_0}\partial_x \phi + a^2(x) \left|\psi_0^2\right| (\phi + \phi^*)$$

homogeneous background solution \Rightarrow

$$i(\partial_t + v_0\partial_x)\phi = -\frac{1}{2}\partial_x^2\phi + a^2(x)\rho_0(\phi + \phi^*)$$

basis of solutions of the form $\phi(x,t)=Ue^{-i\omega t+ikx}+V^*e^{i\omega t-ikx}$ \rightarrow dispersion relation:

$$(\omega - v_0 k)^2 = a^2 \rho_0 k^2 + \frac{k^4}{4}$$

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Decay of linear perturbations: Idea of the calculation (1)

- find a basis of *in* modes $e^{-i\omega t}\phi^{(j)}_{\omega}(x)$
- expand the initial perturbation on this basis:

$$\phi(x,0) = \int \sum_{j} a_{\omega}^{(j)} \phi_{\omega}^{(j)}(x) d\omega$$

• evolve the solution in time:

$$\phi(x,t) = \int e^{-i\omega t} \sum_{j} a_{\omega}^{(j)} \phi_{\omega}^{(j)}(x) d\omega$$

• expand each in mode into plane waves:

$$\phi(x,t) = \int e^{-i\omega t} \sum_{j} a_{\omega}^{(j)} \left(\sum_{l} b_{\omega}^{(j,l)} e^{ik_{\omega}^{l}x} \right) d\omega$$

Decay of linear perturbations: Idea of the calculation (2)

- choice of the space of "acceptable" initial perturbations: no divergence in $a_{\omega}^{(i)}$ nor its derivatives
- late times: integral dominated by two contributions
 - ightarrow points where $\left| b_{\omega}^{(j,l)} \right|
 ightarrow \infty$

 \rightarrow saddle points where $x \frac{dk_{\omega}^{(l)}}{d\omega} = t$

- singularities in $b_{\omega}^{(j,l)}$: only for $\omega \to 0$; come in pairs which cancel each other \to no contribution to leading order
- saddle points at $\omega=\pm\omega_{\rm max}\to{\rm Gaussian}$ integration \to decay in $O(t^{-3/2})$
- (possibly logarithmic factors from the Gaussian integration)

Whitham's modulation theory: Main ideas

[A. Kamchatnov, Nonlinear Periodic Waves and Their Modulations: An Introductory Course, 2000]

- ${\scriptstyle \bullet}$ goal: find approximate solutions of a (1+1)d PDE which either
 - \bullet vary slowly with space and time

• show "fast", quasi-periodic oscillations modulated over a "slow" scale

- \bullet separation of scales \rightarrow averaging procedure
- \bullet average the conservation laws over the "fast" scale \to equations of evolution for locally conserved quantities
- the homogeneous GPE is integrable in the inverse scattering sense \rightarrow infinite number of conserved quantities

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Integrability in the AKNS sense

Integrability in the AKNS sense \Leftrightarrow compatibility condition of the linear system

$$\begin{cases} \partial_x \psi_\lambda(x,t) = U(x,t;\lambda)\psi_\lambda(x,t) \\ \partial_t \psi_\lambda(x,t) = V(x,t;\lambda)\psi_\lambda(x,t) \end{cases}$$

(+technical conditions)

- $\psi_{\lambda}(x,t)$: two-component vector
- $U(x,t;\lambda)$, $V(x,t;\lambda)$: 2 by 2 matrices
- λ : complex parameter ("spectral parameter")

Explicitly,

$$\forall \lambda \in \mathbb{C}, [U(x,t;\lambda), V(x,t;\lambda)] = \partial_x V(x,t;\lambda) - \partial_t U(x,t;\lambda)$$

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AKNS and conserved quantities

consider two solutions ϕ and φ

define
$$g \equiv \phi_1 \varphi_1$$
 and $P \equiv -(\phi_1 \varphi_2 - \phi_2 \varphi_1)^2 / 4$
 $\rightarrow P$ depends only on λ

Conservation law:

$$\forall \lambda, \partial_t \left(\frac{U_{11}}{g} \right) - \partial_x \left(\frac{V_{11}}{g} \right) = 0$$

expansion in powers of $\lambda \rightarrow$ infinite number of conserved quantities

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From exact conservation to slow modulation (1)

- two ways to apply the above for modulated solutions:
 - \bullet do the full analysis exactly \rightarrow difficult to extract precise results

 \bullet use that the solution varies little on small scales \rightarrow define an effective local g

- \bullet the "local" and "global" g are related by a factor \sqrt{P} (defined locally)
- we consider solutions for which P is a polynomial \rightarrow the (exact) conservation laws give evolution equations on its roots λ_i Riemann invariants

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From exact conservation to slow modulation (2)

• Whitham evolution equations:

$$\partial_t \lambda_i + v_i \left(\{\lambda_j\}\right) \partial_x \lambda_i = 0$$
$$v_i \left(\{\lambda_j\}\right) = \frac{\oint \frac{V_{11}(x,t;\mu)}{\sqrt{P(x,t;\mu)}g(x,t;\lambda_i)} d\mu}{\oint \frac{U_{11}(x,t;\mu)}{\sqrt{P(x,t;\mu)}g(x,t;\lambda_i)} d\mu}$$

- \Rightarrow the problem reduces to finding characteristics
- suitable ansatz \Rightarrow known relation between $\{\lambda_i\}$ and ψ

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The AKNS scheme for the homogeneous GPE: main steps

• find U and V for the GPE

$$U = \begin{pmatrix} -i\lambda & ia\psi\\ -ia\psi^* & i\lambda \end{pmatrix}, V = \begin{pmatrix} -i\lambda^2 - i\frac{a^2}{2} \left|\psi^2\right| & ia\lambda\psi - \frac{a}{2}\partial_x\psi\\ -ia\lambda\psi^* & i\lambda^2 + i\frac{a^2}{2} \left|\psi^2\right| \end{pmatrix}$$

• ansatz for g and P:

$$g(x,t;\lambda) = ia\psi(x,t) \left(\lambda - \mu(x,t)\right)$$
$$P(\lambda) = \lambda^4 - s_1\lambda^3 + s_2\lambda^2 - s_3\lambda + s_4$$

justification a posteriori by recovering the solutions we are looking for

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The AKNS scheme for the homogeneous GPE: main steps

• find the relationships between λ_i and the physical solution

$$\rho_3 = \left(\lambda_4 + \lambda_3 - \lambda_1 - \lambda_2\right)^2 / (4a^2)$$

+ permutations

• characteristic velocities:

$$v_i = \frac{1}{2} \left(\frac{L}{\partial_{\lambda_i} L} - s_1 \right)$$

• wavelength of the periodic solutions:

$$L = \frac{a^2}{\sqrt{(\lambda_4 - \lambda_2)(\lambda_3 - \lambda_1)}} K\left(\frac{(\lambda_4 - \lambda_3)(\lambda_2 - \lambda_1)}{(\lambda_4 - \lambda_2)(\lambda_3 - \lambda_1)}\right)$$

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Scale-invariant solutions

the initial configuration introduces no scale \Rightarrow scale-invariant solutions $\lambda_i(z\equiv x/t)$

Whitham's modulation equations:

$$(v_i - z) \frac{d\lambda_i}{dz} = 0 \Rightarrow v_i = z \text{ or } \frac{d\lambda_i}{dz} = 0$$

asymptotically homogeneous \Rightarrow two Riemann invariant are equal

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Scale-invariant solutions



dispersive shock wave (DSW)



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Solutions in step-like potentials

- initial conditions preserving scale invariance:
 - homogeneous density perturbation

 \bullet piecewise-constant density perturbation with a single discontinuity at x=0

- look for solutions with $\rho = \rho_0$ for $z \in I$, where I is an interval containing $0 \rightarrow$ local convergence to a homogeneous solution.
- linear order: two outgoing solutions for x > 0; one for x < 0 \Rightarrow look for (NL) solutions with two waves downstream and one wave upstream
- each wave can be a DSW or a SW \Rightarrow 8 possibilities

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Homogeneous initial condition: two series of solutions





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Homogeneous initial condition: analytical results

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- dimensionless parameters: $\gamma \equiv a_{-}/a_{+}, \ \eta \equiv \sqrt{\rho_{i}/\rho_{0}},$ $\nu \equiv v_{i}/(\sqrt{\rho_{0}}a_{-})$
- solution with 1 DSW and 2 SW \Leftrightarrow

$$\begin{cases} \frac{\gamma-1}{\gamma+1} < \eta < 1\\ \frac{1}{\gamma} + 2(1-\eta) < \nu < 1 \end{cases}$$

 $1 < \eta < \frac{\gamma + 1}{\gamma - 1}$ $2\frac{\eta}{\gamma} - \gamma^{-2} \left(\eta + \frac{\eta}{\gamma} - 1\right)^{-1} < \nu < 3 - 2\eta$

• solution with 2 DSW and 1 SW \Leftrightarrow

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Step-like initial configuration: Domains of existence



Conclusion: A solution always exists when starting sufficiently close to a homogeneous BH flow

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Numerical results (1)

What can be learned from numerical resolution:

- validity of Whitham's theory
- changes due to smooth initial conditions and/or potentials
- time-evolution of configurations outside the domain of existence of solutions with 3 SW/DSW

Numerical integration of the GPE with homogeneous or step-like initial conditions

 \rightarrow very good agreement with the solutions of Whitham's equations at late times (as expected)

 \rightarrow earlier times: solution of Whitham's equations + perturbations in $O(t^{-\alpha})\text{, }\alpha\approx 3/2$

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Numerical results (2)





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Numerical results (3)

 $\bullet\,$ smooth initial conditions and/or potentials $\rightarrow\,$ same late-time solutions

(but formation time of the SW and DSW approximately linear in their scales of variation)

• Initial conditions outside the domain of existence of these solutions \Rightarrow variety of behaviors

 \rightarrow convergence to the homogeneous solution through the emission of overlapping non-linear waves

- $\rightarrow\,$ "hairy BH" with an extended, stationary soliton train
- \rightarrow periodic emission of solitons

Conclusions

- BH flows of the GPE and KdV equation: described by a few conserved quantities (uniqueness)
- the homogeneous (or near-homogeneous) solutions are linearly stable
- non-linear stability: local convergence to the homogeneous (or near-homogeneous) solution through the emission of 3 non-linear waves + decaying perturbations
- $\bullet\,$ white-hole flows: generically unstable $\to\,$ non-asymptotically uniform and/or nonstationary solutions

Outlook

- prove the non-linear stability of BH flows without relying on the hypotheses of the Whitham theory
- extend the analytical results to smooth potentials and more general perturbations
- analyze the stability of BH waterfall solutions
- understand the generation of soliton trains by WH flows

Thank you for your attention!