# Effective Field Theory of Multi-Field Inflation a la Weinberg

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JCAP, arXiv:1203.2266



keywords: Effective Field Theory, Inflation

single field	C. Cheung et al. JHEP, arXiv:0709.0293	S. Weinberg PRD, arXiv:0804.4291
	The EFT of Inflation	EFT for Inflation
multi-field	L. Senatore and M. Zaldarriaga JHEP, arXiv:1009.2093 The EFT of Multifield Inflation	JCAP, arXiv:1203.2266 EFT of Multi-Field Inflation a la Weinberg

# **Effective Field Theory**

an *effective* theory:

is true for a certain domain of energy.

- two cases:
  - as a part of a true theory for whole energy scales
    - using EFT to simplify calculations!
  - in lack of a complete theory for the energy scales of interests
    - using EFT since there is no other choice!
    - constructing EFT, by considering e.g. symmetry properties of the model.

# **EFT for Multi-Field Inflation**

the most general form of Lagrangian up to the 4th order derivatives:

after simplifications:

 $\begin{cases} b_{1}^{IJ}(\vec{\varphi})\Box\varphi_{I}\Box\varphi_{J} + b_{2}^{IJK}(\vec{\varphi})\nabla_{\mu}\varphi_{I}\nabla^{\mu}\varphi_{J}\Box\varphi_{K} + b_{3}^{IJKL}(\vec{\varphi})\nabla_{\mu}\varphi_{I}\nabla^{\mu}\varphi_{J}\nabla_{\nu}\varphi_{K}\nabla^{\nu}\varphi_{L} + b_{4}^{IJ}(\vec{\varphi})\nabla_{\mu}\varphi_{I}\nabla^{\mu}\varphi_{J} \qquad (A1) \\ + b_{5}(\vec{\varphi}) + b_{6}^{IJ}(\vec{\varphi})(\nabla^{\mu}\varphi_{I})(\Box\nabla_{\mu}\varphi_{J}) + b_{7}^{I}(\vec{\varphi})(\nabla_{\mu}\nabla_{\nu}\varphi_{I})^{2} + b_{8}^{IJK}(\vec{\varphi})(\nabla_{\mu}\varphi_{I})(\nabla_{\nu}\varphi_{J})(\nabla^{\mu}\nabla^{\nu}\varphi_{K}) + b_{9}^{I}(\vec{\varphi})\nabla^{\mu}\Box\nabla_{\mu}\varphi_{I} \\ + b_{10}^{I}(\vec{\varphi})\Box^{2}\varphi_{I} + b_{11}^{I}(\vec{\varphi})(\nabla^{\mu})\nabla_{\mu}\Box\varphi_{I} + c_{1}^{IJ}(\vec{\varphi})R\nabla_{\mu}\varphi_{I}\nabla^{\mu}\varphi_{J} + c_{2}^{IJ}(\vec{\varphi})R^{\mu\nu}\nabla_{\mu}\varphi_{I}\nabla_{\nu}\varphi_{J} + c_{3}^{I}(\vec{\varphi})R\Box\varphi_{I} \\ + c_{4}^{I}(\vec{\varphi})(\nabla^{\mu}R)(\nabla_{\mu}\varphi_{I}) + c_{5}(\vec{\varphi})\Box R + c_{6}^{I}(\vec{\varphi})R_{\mu\nu}\nabla^{\mu}\nabla^{\nu}\varphi_{I} + a_{1}(\vec{\varphi})R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} + a_{2}(\vec{\varphi})R_{\mu\nu}R^{\mu\nu} + a_{3}(\vec{\varphi})R^{2} + a_{4}(\vec{\varphi})R \end{cases}$ 

#### before simplifications!

# perturbations:

perturbations in single field model (Weinberg's paper)

$$\begin{split} \mathcal{L} &= \sqrt{g} \bigg[ -\frac{M^2}{2} g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - M_P^2 U(\varphi) + f(\varphi) \bigg( g^{\mu\nu} \varphi_{,\mu} \varphi_{,\nu} \bigg)^2 \bigg] \\ &= \bar{\mathcal{L}} - \frac{1}{2} a^3 \bigg( M^2 + 4f(\bar{\varphi}) \dot{\bar{\varphi}}^2 \bigg) \times \bigg( -\dot{\delta\varphi}^2 + a^{-2} (\vec{\nabla} \delta\varphi)^2 \bigg) \\ &+ 4a^3 f(\bar{\varphi}) \dot{\bar{\varphi}}^2 \bigg( \dot{\delta\varphi}^2 + \dot{\delta\varphi}^3 / \dot{\bar{\varphi}} - a^{-2} \dot{\delta\varphi} (\vec{\nabla} \delta\varphi)^2 / \dot{\bar{\varphi}} + \frac{1}{4} \dot{\delta\varphi}^4 / \dot{\bar{\varphi}}^2 - \frac{1}{2} a^{-2} \dot{\delta\varphi}^2 (\vec{\nabla} \delta\varphi)^2 / \dot{\bar{\varphi}}^2 + \frac{1}{4} a^{-4} (\vec{\nabla} \delta\varphi)^4 / \dot{\bar{\varphi}}^2 \bigg) \end{split}$$

-- it is up to 4<sup>th</sup> order of perturbations automatically.

-- speed of sound  $\neq$  1

-- large non-Gaussianity?

-- speed of sound is constrained by validity of EFT!

# perturbations:

#### multi-field case:

$$a^{-3}\Delta\mathcal{L}^{(2)} = b_{3}^{IJKL}(\bar{\varphi}) \bigg\{ \dot{\bar{\varphi}}_{I} \dot{\bar{\varphi}}_{J} \delta\dot{\varphi}_{K} \delta\dot{\varphi}_{L} + \dot{\bar{\varphi}}_{I} \dot{\bar{\varphi}}_{K} \delta\dot{\varphi}_{J} \delta\dot{\varphi}_{L} + \dot{\bar{\varphi}}_{I} \dot{\bar{\varphi}}_{L} \delta\dot{\varphi}_{K} \delta\dot{\varphi}_{J} + \dot{\bar{\varphi}}_{K} \dot{\bar{\varphi}}_{J} \delta\dot{\varphi}_{L} + \dot{\bar{\varphi}}_{K} \dot{\bar{\varphi}}_{L} \delta\dot{\varphi}_{I} \delta\dot{\varphi}_{J} \bigg\} - a^{-2} \dot{\bar{\varphi}}_{I} \dot{\bar{\varphi}}_{J} \partial_{i} \delta\varphi_{L} - a^{-2} \dot{\bar{\varphi}}_{K} \dot{\bar{\varphi}}_{L} \partial_{i} \delta\varphi_{I} \partial^{i} \delta\varphi_{J} \bigg\} - \frac{M^{2}}{2} \delta^{IJ} \bigg[ - \delta\dot{\varphi}_{I} \delta\dot{\varphi}_{J} + a^{-2} \partial_{i} \delta\varphi_{I} \partial^{i} \delta\varphi_{J} \bigg],$$

$$\begin{aligned} a^{-3}\Delta\mathcal{L}^{(3)} &= b_{3}^{IJKL}(\bar{\varphi}) \bigg\{ \dot{\bar{\varphi}}_{I}\dot{\delta\varphi}_{J}\dot{\delta\varphi}_{K}\dot{\delta\varphi}_{L} + \dot{\bar{\varphi}}_{J}\dot{\delta\varphi}_{K}\dot{\delta\varphi}_{L} + \dot{\bar{\varphi}}_{K}\dot{\delta\varphi}_{L}\dot{\delta\varphi}_{I}\dot{\delta\varphi}_{J} + \dot{\bar{\varphi}}_{L}\dot{\delta\varphi}_{K}\dot{\delta\varphi}_{I}\dot{\delta\varphi}_{J} \\ &- a^{-2} \bigg( \dot{\bar{\varphi}}_{I}\dot{\delta\varphi}_{J}\partial_{i}\delta\varphi_{K}\partial^{i}\delta\varphi_{L} + \dot{\bar{\varphi}}_{J}\dot{\delta\varphi}_{I}\partial_{i}\delta\varphi_{K}\partial^{i}\delta\varphi_{L} + \dot{\bar{\varphi}}_{K}\dot{\delta\varphi}_{L}\partial_{i}\delta\varphi_{I}\partial^{i}\delta\varphi_{J} + \dot{\bar{\varphi}}_{L}\dot{\delta\varphi}_{K}\partial_{i}\delta\varphi_{I}\partial^{i}\delta\varphi_{J} \bigg\}, \end{aligned}$$

$$a^{-3}\Delta\mathcal{L}^{(4)} = b_3^{IJKL}(\bar{\varphi}) \bigg\{ \dot{\delta\varphi_I} \dot{\delta\varphi_J} \dot{\delta\varphi_K} \dot{\delta\varphi_L} - a^{-2} \bigg( \dot{\delta\varphi_I} \dot{\delta\varphi_J} \partial_i \delta\varphi_K \partial^i \delta\varphi_L + \dot{\delta\varphi_K} \dot{\delta\varphi_L} \partial_i \delta\varphi_I \partial^i \delta\varphi_J \bigg) \\ + a^{-4} \partial_i \delta\varphi_I \partial^i \delta\varphi_J \partial_j \delta\varphi_K \partial^j \delta\varphi_L \bigg\}.$$

perturbations!

## perturbations:

#### two-field case:

$$a^{-3}\Delta\mathcal{L}^{(2)} = \dot{\delta\varphi}^{2} \Big[ \frac{M_{1}^{2}}{2} + 6g_{1}\dot{\varphi}^{2} + 3g_{3}\dot{\varphi}\dot{\chi} + (g_{5} + g_{6})\dot{\chi}^{2} \Big] + \dot{\delta\chi}^{2} \Big[ \frac{M_{2}^{2}}{2} + 6g_{2}\dot{\chi}^{2} + 3g_{4}\dot{\varphi}\dot{\chi} + (g_{5} + g_{6})\dot{\varphi}^{2} \Big] \\ + \dot{\delta\varphi}\dot{\delta\chi} \Big[ 3g_{3}\dot{\varphi}^{2} + 3g_{4}\dot{\chi}^{2} + 4(g_{5} + g_{6})\dot{\varphi}\dot{\chi} \Big] \\ - a^{-2} \Big( \partial_{i}\delta\varphi\partial^{i}\delta\varphi \Big[ \frac{M_{1}}{2} + 2g_{1}\dot{\varphi}^{2} + g_{3}\dot{\varphi}\dot{\chi} + g_{5}\dot{\chi}^{2} \Big] + \partial_{i}\delta\chi\partial^{i}\delta\chi \Big[ \frac{M_{2}}{2} + 2g_{2}\dot{\chi}^{2} + g_{4}\dot{\varphi}\dot{\chi} + g_{5}\dot{\varphi}^{2} \Big] \\ + \partial_{i}\delta\varphi\partial^{i}\delta\chi \Big[ g_{3}\dot{\varphi}^{2} + g_{4}\dot{\chi}^{2} + 2g_{6}\dot{\varphi}\dot{\chi} \Big] \Big),$$

.... and cubic and quartic terms!

... transition to adiabatic and entropy curvature perturbations!

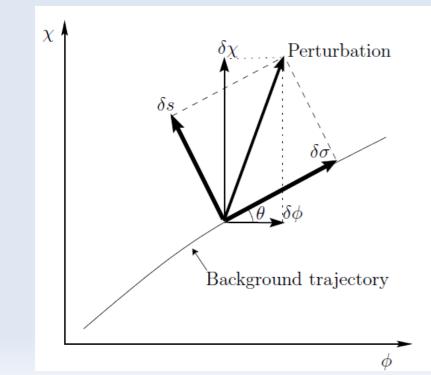
### perturbations: adiabatic & entropy modes

#### adiabatic mode:

entropy mode:

$$\delta \sigma \equiv \vec{T}.\vec{\delta}, \qquad \qquad \delta s \equiv \vec{N}.\vec{\delta}$$

$$\vec{\delta} \equiv (\delta \varphi, \delta \chi), \qquad \qquad \vec{T} = (\cos \theta, \sin \theta) \equiv (\dot{\varphi}/\dot{\sigma}, \dot{\chi}/\dot{\sigma}), \qquad \qquad \vec{N} \equiv (\sin \theta, -\cos \theta)$$



Gordon et al. arXiv:astro-ph/0009131

### perturbations: adiabatic & entropy modes

 $(\delta s + \theta \delta \sigma)$ 

note that just these two combinations appear in this formalism!

# shape of non-Gaussianity

due to previous slide: for example:

$$(\vec{T}.\vec{\delta})^3 = \delta \dot{\sigma}^3 - 3 \dot{\theta} \delta \dot{\sigma}^2 \delta s + 3 \dot{\theta}^2 \dot{\delta \sigma} \delta s^2 - \dot{\theta}^3 \delta s^3$$

equilateral NG in adiabatic mode

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local NG in entropy mode

--- in this formalism

the "Cosine" between different kinds of NG is fixed!

# amplitude of NG

$$\frac{\mathcal{L}^{(3)}}{\mathcal{L}^{(2)}} = \frac{\{H^3, H^2\dot{\theta}, H\dot{\theta}^2, \dot{\theta}^3\}}{\{H^2, H\dot{\theta}, \dot{\theta}^2\}} \times \left(\frac{f(g_i)}{M^2}\dot{\sigma}^2\right) \times \frac{\delta\sigma}{\dot{\sigma}} = \frac{\{H^3, H^2\dot{\theta}, H\dot{\theta}^2, \dot{\theta}^3\}}{H \times \{H^2, H\dot{\theta}, \dot{\theta}^2\}} \times \left(\frac{f(g_i)}{M^2}\dot{\sigma}^2\right) \times \zeta$$

$$\dot{\theta} >> H$$

$$f_{NL} \sim \frac{\mathcal{L}^{(3)}}{\mathcal{L}^{(2)}} \zeta^{-1} \sim \frac{\dot{\theta}}{H} \times \left(\frac{f(g_i)}{M^2} \dot{\sigma}^2\right)$$

-- validity condition of EFT i.e.  $\frac{f(g_i)}{M^2} \dot{\sigma}^2 < 1$  constrains the amplitude of NG!

-- except if the curvature of classical (background) path be large!

-- or: if by a mechanism (e.g. Vainshtein) one can modify the validity condition of EFT!

-- Senatore & Zaldarriaga model is based on Cheung et al.'s work!

-- in Cheung's work, EFT is constructed on perturbations' level!

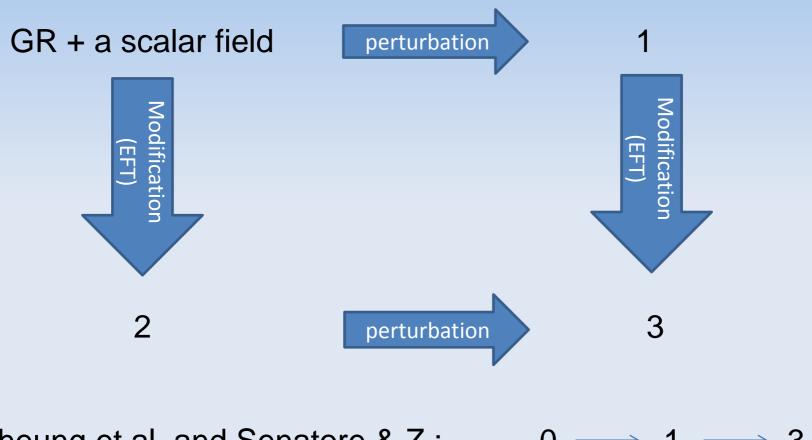
-- since their model is single field, the perturbation is associated to adiabatic mode!

-- so in Senatore & Z., the entropy modes are added into a base with already known adiabatic mode!

but

-- in our case we started with zeroth order term of perturbations.

-- then defined the perturbations without any distinguishability between adiabatic and entropy modes!



Cheung et al. and Senatore & Z.: $0 \longrightarrow 1 \longrightarrow 3$ Weinberg and this talk: $0 \longrightarrow 2 \longrightarrow 3$ 

 so in Senatore & Zaldarriaga, the shift symmetry results in a Lagrangian similar to

$$a^{-3}\Delta\mathcal{L}^{(2)} = \dot{\delta\varphi}^{2} \Big[ \frac{M_{1}^{2}}{2} + 6g_{1}\dot{\varphi}^{2} + 3g_{3}\dot{\varphi}\dot{\chi} + (g_{5} + g_{6})\dot{\chi}^{2} \Big] + \dot{\delta\chi}^{2} \Big[ \frac{M_{2}^{2}}{2} + 6g_{2}\dot{\chi}^{2} + 3g_{4}\dot{\varphi}\dot{\chi} + (g_{5} + g_{6})\dot{\varphi}^{2} \Big] + \dot{\delta\varphi}\dot{\delta\chi} \Big[ 3g_{3}\dot{\varphi}^{2} + 3g_{4}\dot{\chi}^{2} + 4(g_{5} + g_{6})\dot{\varphi}\dot{\chi} \Big] - a^{-2} \Big( \partial_{i}\delta\varphi\partial^{i}\delta\varphi \Big[ \frac{M_{1}}{2} + 2g_{1}\dot{\varphi}^{2} + g_{3}\dot{\varphi}\dot{\chi} + g_{5}\dot{\chi}^{2} \Big] + \partial_{i}\delta\chi\partial^{i}\delta\chi \Big[ \frac{M_{2}}{2} + 2g_{2}\dot{\chi}^{2} + g_{4}\dot{\varphi}\dot{\chi} + g_{5}\dot{\varphi}^{2} \Big] + \partial_{i}\delta\varphi\partial^{i}\delta\chi \Big[ g_{3}\dot{\varphi}^{2} + g_{4}\dot{\chi}^{2} + 2g_{6}\dot{\varphi}\dot{\chi} \Big] \Big),$$

 $\delta \varphi \rightarrow \delta \varphi + c_1 \text{ and } \delta \chi \rightarrow \delta \chi + c_2$ 

i.e. there are just *derivatives* of adiabatic and entropy perturbations!

but in our model the case is different!

shift sym	metry: $\delta_{\varphi}$	$\varphi \to \delta \varphi + c_1$ and $\phi$	$\delta\chi \to \delta\chi + c_2$	
due to				
	$\delta \sigma \equiv \vec{T}.\vec{\delta},$	$\delta s\equiv \vec{N}.\vec{\delta}$		
$\vec{\delta} \equiv \left(\delta\varphi, \delta\chi\right),$	$\vec{T} = (\cos \theta,$	$\sin \theta \equiv (\dot{\varphi}/\dot{\sigma}, \dot{\chi}/\dot{\sigma}),$	$\vec{N} \equiv (s$	$\sin\theta, -\cos\theta)$

results in

$$\delta \sigma \to \delta \sigma + (c_1 \cos \theta + c_2 \sin \theta)$$
  
 $\delta s \to \delta s + (c_1 \sin \theta - c_2 \cos \theta)$ 

which causes a new symmetry for adiabatic and entropy modes:

$$\begin{aligned} \dot{\delta\sigma} - \dot{\theta}\deltas \to \dot{\delta\sigma} - \dot{\theta}\deltas & \longleftarrow \vec{T}.\vec{\delta} \\ \dot{\deltas} + \dot{\theta}\delta\sigma \to \dot{\deltas} + \dot{\theta}\delta\sigma & \longleftarrow \vec{N}.\vec{\delta} \end{aligned}$$

# conclusions

-- this model does not predict a large non-Gaussianity except:

- -- for a highly curved classical path in phase-space!
- -- or if a shielding mechanism allows large first correction term in EFT.

-- different shapes of non-Gaussianity are correlated!

-- in contrast to Senatore & Zaldarriaga, we suggest EFT for multifiled inflation should be constructed as

$$\Delta \mathcal{L} \propto \sum c_{n_0, n_1, \dots, n_N} \left( \vec{T} \cdot \vec{\delta} \right)^{n_0} \left( \vec{N}_1 \cdot \vec{\delta} \right)^{n_1} \left( \vec{N}_2 \cdot \vec{\delta} \right)^{n_2} \dots \left( \vec{N}_N \cdot \vec{\delta} \right)^{n_N}$$

# thank you!

