# Regular DGP in the core of a hypermonopole

Antonio De Felice Louvain University IAP – 16 February 2009 with Christophe Ringeval



# Introduction

• Do we live in D > 4 dimensions?

## Introduction

#### • Do we live in D > 4 dimensions?

Deepest idea of '900 – ST motivated

## Introduction

- Do we live in D > 4 dimensions?
- Deepest idea of '900 ST motivated
- Maybe traces at LHC

• How can gravity be 4D?

- How can gravity be 4D?
- ST: KK way, compact small ED

- How can gravity be 4D?
- ST: KK way, compact small ED
- RS: non-compact, finite volume ED

- How can gravity be 4D?
- ST: KK way, compact small ED
- RS: non-compact, finite volume ED
- DGP: non-compact, infinite volume

## DGP model:

$$S = \frac{M_P^2}{2} \int d^4 V R + \frac{M_*^{2+n}}{2} \int d^{n+4} V R$$

#### DGP model:

$$S = \frac{M_P^2}{2} \int d^4 V R + \frac{M_*^{2+n}}{2} \int d^{n+4} V R$$

Basis of success and problems [Dvali, Gabadadze, Porrati PLB (2000)][Gregory,

Kaloper, Myers, Padilla JHEP (2007)]

#### DGP model:

$$S = \frac{M_P^2}{2} \int d^4 V R + \frac{M_*^{2+n}}{2} \int d^{n+4} V R$$

# Basis of success and problems [Dvali, Gabadadze, Porrati PLB (2000)][Gregory,

Kaloper, Myers, Padilla JHEP (2007)]

#### • Chosen profile for $M_st(X)$ [Kolanovic, Porrati, Rombouts PRD (2003)][Shaposnikov,

Tinyakov, Zuleta PRD (2004)]

## • Varying mass $\rightarrow$ ScT-like theory

- Varying mass ScT-like theory
- Cascading DGP: ghost appears but cured via CC

- Varying mass ScT-like theory
- Cascading DGP: ghost appears but cured via CC
- Is there a classical FT which has DGP mechanism?

#### • Brane at the core of a string in 6D [Ringeval, Rombouts PRD (2005)]

- Brane at the core of a string in 6D [Ringeval, Rombouts PRD (2005)]
- 4D gravity trapping not achieved

- Brane at the core of a string in 6D [Ringeval, Rombouts PRD (2005)]
- 4D gravity trapping not achieved
- Unless violate energy conditions

- Brane at the core of a string in 6D [Ringeval, Rombouts PRD (2005)]
- 4D gravity trapping not achieved
- Unless violate energy conditions
- No-go theorem: Gravity was 6D

Introducing more dimensions

- Introducing more dimensions
- Potential gets a term, n angular ED

- Introducing more dimensions
- Potential gets a term, n angular ED

$$V \propto \frac{n(n-1)k}{r^2} + \dots$$

- Introducing more dimensions
- Potential gets a term, n angular ED

$$V \propto \frac{n(n-1)k}{r^2} + \dots$$

• Needed n > 1 and k > 0

- Introducing more dimensions
- Potential gets a term, n angular ED

$$V \propto \frac{n(n-1)k}{r^2} + \dots$$

- Needed n > 1 and k > 0
- Positively curved, n=2



#### Choose 3 ED

#### Hints

#### Choose 3 ED

### • Topology $R \times S^2$ , asympt. $R^3$

### Hints

#### Choose 3 ED

- Topology  $R imes S^2$ , asympt.  $R^3$
- Topological defect in the origin

### Hints

- Choose 3 ED
- Topology  $R \times S^2$ , asympt.  $R^3$
- Topological defect in the origin
- 't Hooft-Polyakov Monopole

Topological defects: Classical Field Theory

- Topological defects: Classical Field Theory
- DGP can be realized in 7D 'tP-monopole

- Topological defects: Classical Field Theory
- DGP can be realized in 7D 'tP-monopole
- 3 ED, associated SO(3) 'tP field configuration

- Topological defects: Classical Field Theory
- DGP can be realized in 7D 'tP-monopole
- 3 ED, associated SO(3) 'tP field configuration
- 3 ED asympt flat: min.dim. positely curved

 $S = \frac{1}{2\kappa^2} \int d^7 V e^{\psi} [R - \partial \psi^2 - U]$ 

$$S = \frac{1}{2\kappa^2} \int d^7 V e^{\psi} [R - \partial \psi^2 - U] + \int d^7 V [-\frac{1}{2}D\vec{\Phi}^2 - \frac{1}{4}\vec{H}^2 - \frac{1}{8}\lambda(\vec{\Phi}^2 - v^2)^2]$$

$$S = \frac{1}{2\kappa^2} \int d^7 V e^{\psi} [R - \partial \psi^2 - U] + \int d^7 V [-\frac{1}{2}D\vec{\Phi}^2 - \frac{1}{4}\vec{H}^2 - \frac{1}{8}\lambda(\vec{\Phi}^2 - v^2)^2]$$

• EF,  $U=m_d^2\psi^2\exp(rac{2}{5}\psi)$ , SO(3) Higgs  $ec{\Phi}$ 

$$S = \frac{1}{2\kappa^2} \int d^7 V e^{\psi} [R - \partial \psi^2 - U] + \int d^7 V [-\frac{1}{2}D\vec{\Phi}^2 - \frac{1}{4}\vec{H}^2 - \frac{1}{8}\lambda(\vec{\Phi}^2 - v^2)^2]$$

• EF,  $U = m_d^2 \psi^2 \exp(\frac{2}{5}\psi)$ , SO(3) Higgs  $\vec{\Phi}$ •  $D_A \vec{\Phi} = \partial_A \vec{\Phi} - q \vec{C}_A \times \vec{\Phi}$ 

$$S = \frac{1}{2\kappa^2} \int d^7 V e^{\psi} [R - \partial \psi^2 - U] + \int d^7 V [-\frac{1}{2}D\vec{\Phi}^2 - \frac{1}{4}\vec{H}^2 - \frac{1}{8}\lambda(\vec{\Phi}^2 - v^2)^2]$$

• EF,  $U=m_d^2\psi^2\exp(rac{2}{5}\psi)$ , SO(3) Higgs  $ec{\Phi}$ 

•  $D_A \vec{\Phi} = \partial_A \vec{\Phi} - q \vec{C}_A \times \vec{\Phi}$ 

•  $\vec{H}_{AB} = \partial_A \vec{C}_B - \partial_B \vec{C}_A - q \vec{C}_A \times \vec{C}_B$
# Directions

Choose ansatz

# **Directions**

### Choose ansatz

Find background solution

# Directions

- Choose ansatz
- Find background solution
- Study perturbations

#### • Static ansatz

 $ds^2 = e^{\sigma(r)} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + dr^2 + \omega(r)^2 d\Omega^2$ 

#### Static ansatz

$$ds^2 = e^{\sigma(r)} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + dr^2 + \omega(r)^2 d\Omega^2$$

### • $r, \theta, \phi$ spherical coordinates

#### Static ansatz

$$ds^2 = e^{\sigma(r)} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + dr^2 + \omega(r)^2 d\Omega^2$$

- $r, \theta, \phi$  spherical coordinates
- Mapping from SO(3) to 3 ED's:  $\vec{\Phi} = vf(r)\vec{u}_r$

#### Static ansatz

$$ds^2 = e^{\sigma(r)}\eta_{\mu
u}dx^\mu dx^
u + dr^2 + \omega(r)^2 d\Omega^2$$

- $r, \theta, \phi$  spherical coordinates
- Mapping from SO(3) to 3 ED's:  $\vec{\Phi} = vf(r)\vec{u}_r$
- Gauge fields

$$\vec{C}_{\theta} = \frac{1 - Q(r)}{q} \vec{u}_{\phi} \quad \vec{C}_{\phi} = -\frac{1 - Q(r)}{q} \sin \theta \vec{u}_{\theta}$$

Time independent energy distribution

- Time independent energy distribution
- Building the monopole

- Time independent energy distribution
- Building the monopole
- $f \rightarrow 1, Q \rightarrow 0$  as  $r \rightarrow \infty$

- Time independent energy distribution
- Building the monopole

• 
$$f \rightarrow 1, Q \rightarrow 0$$
 as  $r \rightarrow \infty$ 

• f(0) = 0, Q(0) = 1

•  $\sigma=0,\omega
ightarrow r,\psi
ightarrow 0$  as  $r
ightarrow\infty$ 

• 
$$\sigma=0,\omega o r,\psi o 0$$
 as  $r o\infty$   
•  $\sigma'(0)=\psi'(0)=0,\,\omega\sim r$  as  $r o 0$ 

• 
$$\sigma = 0, \omega \to r, \psi \to 0$$
 as  $r \to \infty$   
•  $\sigma'(0) = \psi'(0) = 0, \ \omega \sim r$  as  $r \to 0$ 

Define

$$\alpha \equiv \kappa^2 v^2 \quad \epsilon \equiv \frac{q^2 v^2}{\lambda v^2} \quad \beta \equiv \frac{m_d^2}{\lambda v^2}$$

• 
$$\sigma = 0, \omega \to r, \psi \to 0$$
 as  $r \to \infty$   
•  $\sigma'(0) = \psi'(0) = 0, \ \omega \sim r$  as  $r \to 0$ 

Define

$$\alpha \equiv \kappa^2 v^2 \quad \epsilon \equiv \frac{q^2 v^2}{\lambda v^2} \quad \beta \equiv \frac{m_d^2}{\lambda v^2}$$

• 3 param. family

• The metric is Minkowski

- The metric is Minkowski
- Solutions outside a black-hole

- The metric is Minkowski
- Solutions outside a black-hole
- Regular everywhere: TopDef

- The metric is Minkowski
- Solutions outside a black-hole
- Regular everywhere: TopDef
- No HOR, no SING

### 10 1st-order non-linear ODE's

### 10 1st-order non-linear ODE's

Relaxation methods: failure

- 10 1st-order non-linear ODE's
- Relaxation methods: failure
- Bad initial guess?

- 10 1st-order non-linear ODE's
- Relaxation methods: failure
- Bad initial guess?
- Linearization?

# **Shooting method**

It works, but slow

# **Shooting method**

- It works, but slow
- Need to shoot from 0 and infinity

## **Shooting method**

- It works, but slow
- Need to shoot from 0 and infinity
- Compactify infinity into a segment,  $0 < \rho < 1$ ,

$$r = \frac{\rho}{1 - \rho}$$

 $\bullet$  Shoot from 0 to middle point,  $\rho=1/2$ 

- Shoot from 0 to middle point, ho=1/2
- Shoot from 1 to middle point

- Shoot from 0 to middle point, ho=1/2
- Shoot from 1 to middle point
- 10 conditions to impose at  $\rho = 1/2$

- Shoot from 0 to middle point, ho=1/2
- Shoot from 1 to middle point
- 10 conditions to impose at  $\rho=1/2$
- Continuity of fields and deriv

Need 10 shooting parameters

- Need 10 shooting parameters
- Invert 10x10 matrix

- Need 10 shooting parameters
- Invert 10x10 matrix
- Need to Taylor exp. at 0 and 1

- Need 10 shooting parameters
- Invert 10x10 matrix
- Need to Taylor exp. at 0 and 1
- Find indep. shooting parameters

## New methods

### Recent method in 2 points BVP

### **New methods**

### Recent method in 2 points BVP

• Cash and Mazzia, algorithm realized in Fortran
## **New methods**

- Recent method in 2 points BVP
- Cash and Mazzia, algorithm realized in Fortran
- Translated into C++

## **New methods**

- Recent method in 2 points BVP
- Cash and Mazzia, algorithm realized in Fortran
- Translated into C++
- Complex algorithm

Pro: Much Faster,

- Pro: Much Faster,
- Stable,

- Pro: Much Faster,
- Stable,
- Same solutions of shooting,

- Pro: Much Faster,
- Stable,
- Same solutions of shooting,
- Explore param. space.

- Pro: Much Faster,
- Stable,
- Same solutions of shooting,
- Explore param. space.
- Cons =  $\{\}$



Example  $\alpha = 2$ ,  $\epsilon = 0.5$ ,  $\beta = 1$ . Typical topological configuration for the fields. Minkowski at boundary. TWPBVP: Cash and Mazzia algorithm, JCAM (2005).



Region where  $\omega$  no longer grows, cylindrically shaped ED, resonant gravitons at this scale

• Lowering  $\epsilon$ , spread out

- Lowering  $\epsilon$ , spread out
- $\alpha$  very high: problems, BH?

- Lowering  $\epsilon$ , spread out
- $\alpha$  very high: problems, BH?
- Change  $\beta$ , diff. shape

- Lowering  $\epsilon$ , spread out
- $\alpha$  very high: problems, BH?
- Change  $\beta$ , diff. shape
- Dilaton, no strong influence

Higgs, gauge fields: typical

- Higgs, gauge fields: typical
- Exponential decay at infinity

- Higgs, gauge fields: typical
- Exponential decay at infinity
- Other fields, inverse pow at infty values

- Higgs, gauge fields: typical
- Exponential decay at infinity
- Other fields, inverse pow at infty values
- $\bullet$  No numerical instability found, except too high  $\alpha$

## Directions

Background solution

## **Directions**

- Background solution
- Study perturbations

## Directions

- Background solution
- Study perturbations
- Look for (meta-)stable modes

## **Tensor fluctuations**

• Introduce 4D TT pert:  $\eta_{\mu\nu} \rightarrow \eta_{\mu\nu} + h_{\mu\nu}$ 

## **Tensor fluctuations**

- Introduce 4D TT pert:  $\eta_{\mu\nu} \rightarrow \eta_{\mu\nu} + h_{\mu\nu}$
- New radial variable, and  $A = \exp(\sigma)$ ,

$$z\equiv m_h\int e^{-\sigma/2}dr$$

#### **Tensor fluctuations**

- Introduce 4D TT pert:  $\eta_{\mu\nu} \rightarrow \eta_{\mu\nu} + h_{\mu\nu}$
- New radial variable, and  $A = \exp(\sigma)$ ,

$$z\equiv m_h\int e^{-\sigma/2}dr$$

Linearize Einstein equations

$$h_{\mu\nu}'' + F'h_{\mu\nu}' + m_h^{-2}\Box h_{\mu\nu} - \frac{A}{\omega^2}L^2h_{\mu\nu} = 0$$

$$h_{\mu\nu}'' + F'h_{\mu\nu}' + m_h^{-2}\Box h_{\mu\nu} - \frac{A}{\omega^2}L^2h_{\mu\nu} = 0$$

• 
$$F' = \frac{3}{2}A'/A + 2\omega'/\omega + \psi'$$

$$h''_{\mu\nu} + F'h'_{\mu\nu} + m_h^{-2}\Box h_{\mu\nu} - \frac{A}{\omega^2}L^2h_{\mu\nu} = 0$$

• 
$$F' = \frac{3}{2}A'/A + 2\omega'/\omega + \psi'$$

• 4D Fourier + Spherical Harmonics

$$h = \sum_{m,l} Y_{lm}( heta,\phi) \int dar{M} \hat{h}_{lm}(ar{M},z) \Delta_{ar{M}}(x)$$

$$h''_{\mu\nu} + F'h'_{\mu\nu} + m_h^{-2}\Box h_{\mu\nu} - \frac{A}{\omega^2}L^2h_{\mu\nu} = 0$$

• 
$$F' = \frac{3}{2}A'/A + 2\omega'/\omega + \psi'$$

• 4D Fourier + Spherical Harmonics

$$h = \sum_{m,l} Y_{lm}( heta,\phi) \int dar{M} \hat{h}_{lm}(ar{M},z) \Delta_{ar{M}}(x)$$

•  $\Delta$  eigenfun of  $m_h^{-2}\Box$  with  $\bar{M}^2$  eigenval

$$h'' + F'h' - \frac{A}{\omega^2}l(l+1)h + \bar{M}^2h = 0$$

$$h'' + F'h' - \frac{A}{\omega^2}l(l+1)h + \bar{M}^2h = 0$$

• New perturbation field,  $\xi(0) = 0$ ,  $\xi = e^{\psi/2 + 3\sigma/4} \omega h$ 

$$h'' + F'h' - \frac{A}{\omega^2}l(l+1)h + \bar{M}^2h = 0$$

• New perturbation field,  $\xi(0) = 0$ ,  $\xi = e^{\psi/2 + 3\sigma/4} \omega h$ 

#### • No term in $\xi'$

$$-\xi'' + \left[ W^2 + W' + \frac{A}{\omega^2} l(l+1) \right] \xi = \bar{M}^2 \xi$$

$$h'' + F'h' - \frac{A}{\omega^2}l(l+1)h + \bar{M}^2h = 0$$

• New perturbation field,  $\xi(0) = 0$ ,  $\xi = e^{\psi/2 + 3\sigma/4} \omega h$ 

• No term in  $\xi'$ 

$$-\xi'' + \left[ W^2 + W' + \frac{A}{\omega^2} l(l+1) \right] \xi = \bar{M}^2 \xi$$

•  $W = \frac{3}{4}\sigma' + \omega'/\omega + \frac{1}{2}\psi'$ 

## • 1D Schroedinger eq., for z > 0

• 1D Schroedinger eq., for z > 0

Rewritten as

 $[TT^{\dagger} + l(l+1)A\omega^{-2}]\xi = \bar{M}^2\xi$ 

• 1D Schroedinger eq., for z > 0

Rewritten as

$$[TT^{\dagger} + l(l+1)A\omega^{-2}]\xi = \bar{M}^2\xi$$

•  $\overline{M} \ge 0$ , and operator T = d/dz + W

• 1D Schroedinger eq., for z > 0

Rewritten as

$$[TT^{\dagger} + l(l+1)A\omega^{-2}]\xi = \bar{M}^2\xi$$

•  $\overline{M} \ge 0$ , and operator T = d/dz + W

• Zero mode,  $\bar{M}=l=0$ ,  $\xi'=W\xi$ 

• 1D Schroedinger eq., for z > 0

Rewritten as

$$[TT^{\dagger} + l(l+1)A\omega^{-2}]\xi = \bar{M}^2\xi$$

- $\overline{M} \ge 0$ , and operator T = d/dz + W
- Zero mode,  $\bar{M}=l=0$ ,  $\xi'=W\xi$
- Not normalizable
## **SUSY-QM**

## • QM with central potential $V_{2,l=0} \equiv W^2 + W'$

## **SUSY-QM**

#### • QM with central potential $V_{2,l=0} \equiv W^2 + W'$



Depends on  $\alpha, \beta, \epsilon$ . For  $V_2$ : ground state  $\xi_0$  not normalizable. For  $V_1$ ,  $1/\xi_0$ not regular in 0. SUSY is broken and spectrum  $M^2 > 0$ .

## Solution of radial SE

•  $u_{M,l}(z)$  solution of Schr.Eq.

## Solution of radial SE

## • $u_{M,l}(z)$ solution of Schr.Eq.

Orthonormal basis

## **Solution of radial SE**

- $u_{M,l}(z)$  solution of Schr.Eq.
- Orthonormal basis
- Completeness relation

$$\int_0^\infty u^*_{M,l}(z_1) u_{M,l}(z_2) dM = \delta(z_1-z_2)$$

## • Retarded Green function for TT source $S_{\mu u}$ [Garriga, Tanaka PRL

(2000)]

## • Retarded Green function for TT source $S_{\mu\nu}$ [Garriga, Tanaka PRL (2000)]

#### • Find

$$h_{\mu\nu}(X_1) = -\frac{2\kappa^2}{m_h^2\omega} e^{-\psi(z_1) - 3\sigma/4} \\ \times \int G_{\xi}(X_1; X_2) e^{-\psi(z_2) + 3\sigma/4} \omega(z_2) S_{\mu\nu} d^7 X_2$$

#### **GF** expression

# $G_{\xi}'' + (-V_2 + \Box + e^{\sigma} \omega^{-2} L) G_{\xi} = \delta^4 (x^{\mu} - x^{\mu'})$ $\times \delta(\cos\theta - \cos\theta') \delta(\phi - \phi') \delta(z - z')$

## **GF** expression

$$G_{\xi}'' + (-V_2 + \Box + e^{\sigma}\omega^{-2}L)G_{\xi} = \delta^4(x^{\mu} - x^{\mu'})$$
$$\times \delta(\cos\theta - \cos\theta')\delta(\phi - \phi')\delta(z - z')$$

$$\begin{aligned} G_{\xi} &= -\int \frac{d^4 p}{(2\pi)^4} e^{i p_{\mu}(x_1^{\mu} - x_2^{\mu})} \sum_{l,m} Y_l^m(\theta_1, \phi_1) Y_l^{m*}(\theta_2, \phi_2) \\ & \times \int \frac{u_{M,l}(z_1) u_{M,l}^*(z_2) dM}{M^2 + \vec{p}^2 - (p^0 + i\epsilon)^2} \end{aligned}$$

#### Flat case

• Setting  $\psi = \sigma = 0, \omega = r$ 

#### Flat case

- Setting  $\psi = \sigma = 0, \omega = r$
- Source  $S_{\mu\nu} = z^{-2}\delta(z)\delta(\cos\theta)\delta(\phi)s_{\mu\nu}$

#### Flat case

- Setting  $\psi = \sigma = 0, \omega = r$
- Source  $\overline{S_{\mu\nu}} = z^{-2}\delta(z)\delta(\cos\overline{\theta})\delta(\phi)s_{\mu\nu}$
- Radial solution:  $u_{M,l}^{\flat} = \sqrt{Mz} J_{l+1/2}(Mz)$

Contribution in the origin

- Contribution in the origin
- Only l = 0 mode gives contribution

- Contribution in the origin
- Only l = 0 mode gives contribution
- Solution

$$h_{\mu\nu}^{\flat} = \lim_{z \to 0} \frac{\kappa^2}{8\pi^2 m_h^2} \int d^3 \vec{x}_2 s_{\mu\nu} \int dM \frac{|u_{M,0}^{\flat}(z)|^2}{z^2} \frac{e^{-M\Delta \vec{x}}}{|\Delta \vec{x}|}$$

## - Bessel expansion $u_{M,0}^\flat\sim\sqrt{2/\pi}Mz$

- Bessel expansion  $u^{\flat}_{M,0} \sim \sqrt{2/\pi} M z$ 

Solution simplifies to

$$h_{\mu\nu}^{\flat} = \frac{2\kappa^2}{4\pi^3 m_h^2} \int d^3 \vec{x}_2 \frac{s_{\mu\nu}(\vec{x}_2)}{|\Delta \vec{x}|^4}$$

• Bessel expansion  $u_{M,0}^{\flat} \sim \sqrt{2/\pi}Mz$ 

Solution simplifies to

$$h_{\mu\nu}^{\flat} = \frac{2\kappa^2}{4\pi^3 m_h^2} \int d^3 \vec{x}_2 \frac{s_{\mu\nu}(\vec{x}_2)}{|\Delta \vec{x}|^4}$$

- Power law dependence  $1/|\Delta \vec{x}|^{d-2}$  in d=6 spatial dimensions

- Bessel expansion  $u_{M,0}^{\flat} \sim \sqrt{2/\pi}Mz$
- Solution simplifies to

$$h_{\mu\nu}^{\flat} = \frac{2\kappa^2}{4\pi^3 m_h^2} \int d^3 \vec{x}_2 \frac{s_{\mu\nu}(\vec{x}_2)}{|\Delta \vec{x}|^4}$$

- Power law dependence  $1/|\Delta \vec{x}|^{d-2}$  in d=6 spatial dimensions
- In d = 3, 1/r dependence: 4D gravity

- Bessel expansion  $u_{M,0}^{\flat} \sim \sqrt{2/\pi}Mz$
- Solution simplifies to

$$h_{\mu\nu}^{\flat} = \frac{2\kappa^2}{4\pi^3 m_h^2} \int d^3 \vec{x}_2 \frac{s_{\mu\nu}(\vec{x}_2)}{|\Delta \vec{x}|^4}$$

- Power law dependence  $1/|\Delta \vec{x}|^{d-2}$  in d=6 spatial dimensions
- In d = 3, 1/r dependence: 4D gravity

•  $4\pi^3$ : d-2 times the surface of d-1 unit sphere.

## Monopole case

#### At infinity to Bessel functions

#### Monopole case

- At infinity to Bessel functions
- Need spectral density  $ho(M) \equiv |u_{M,0}(0)|^2/|u_{M,0}^{\flat}(0)|^2$

#### Monopole case

- At infinity to Bessel functions
- Need spectral density  $ho(M) \equiv |u_{M,0}(0)|^2/|u_{M,0}^{\flat}(0)|^2$

Therefore one finds

$$h_{\mu\nu} = \frac{2\kappa^2}{8\pi^3 m_h^2} \int d^3 \vec{x}_2 \frac{\mathcal{L}\{\rho(M)M^2\}}{|\Delta \vec{x}|} s_{\mu\nu}(\vec{x}_2)$$





#### • Constant $\rho$ : 7D gravity



- Constant  $\rho$ : 7D gravity
- $\rho$  strongly peaked: resonant metastable modes



- Constant  $\rho$ : 7D gravity
- $\rho$  strongly peaked: resonant metastable modes
- At infinity,  $A_\infty/A_0 \ll 1$



- Constant  $\rho$ : 7D gravity
- $\rho$  strongly peaked: resonant metastable modes
- At infinity,  $A_\infty/A_0 \ll 1$
- No  $M^2 < 0$  bound state

#### • Trapping a graviton, $\rho \sim 1 + C\delta(M - m_g)$

• Trapping a graviton,  $\rho \sim 1 + C\delta(M - m_g)$ •  $\mathcal{L}\{\rho M^2\} = 2/|\Delta \vec{x}|^3 + Cm_g^2 e^{-m_g|\Delta \vec{x}|}$ 

• Trapping a graviton,  $\rho \sim 1 + C\delta(M - m_g)$ •  $\mathcal{L}\{\rho M^2\} = 2/|\Delta \vec{x}|^3 + Cm_g^2 e^{-m_g|\Delta \vec{x}|}$ • 4D gravity:  $(m_g/C)^{1/3} < |\Delta \vec{x}|m_g < 1$ , and  $m_q < C$ 

Trapping a graviton, ρ ~ 1 + Cδ(M − m<sub>g</sub>)
L{ρM<sup>2</sup>} = 2/|Δx̄|<sup>3</sup> + Cm<sup>2</sup><sub>g</sub>e<sup>-m<sub>g</sub>|Δx̄|</sup>
4D gravity: (m<sub>g</sub>/C)<sup>1/3</sup> < |Δx̄|m<sub>g</sub> < 1, and m<sub>g</sub> < C</li>
7D at small/large distance

• Fixing 4D between two scales  $s_1 < \Delta x < s_2$ 

• Fixing 4D between two scales  $s_1 < \Delta x < s_2$ • DGP if  $m_g s_2 < 1$  and  $m_g s_1 > (m_g/C)^{1/3}$  • Fixing 4D between two scales  $s_1 < \Delta x < s_2$ • DGP if  $m_g s_2 < 1$  and  $m_g s_1 > (m_g/C)^{1/3}$ • Tuning?

- Fixing 4D between two scales  $s_1 < \Delta x < s_2$
- DGP if  $m_g s_2 < 1$  and  $m_g s_1 > (m_g/C)^{1/3}$
- Tuning?
- We observed fractional power
- Fixing 4D between two scales  $s_1 < \Delta x < s_2$
- DGP if  $m_g s_2 < 1$  and  $m_g s_1 > (m_g/C)^{1/3}$
- Tuning?
- We observed fractional power