



Critical Tests of Theory of the Early Universe Using the Cosmic Microwave Background

Hiichiro Komatsu (Max-Planck-Institut für Astrophysik)
Colloquium, IAP, November 9, 2012

The Breakthrough

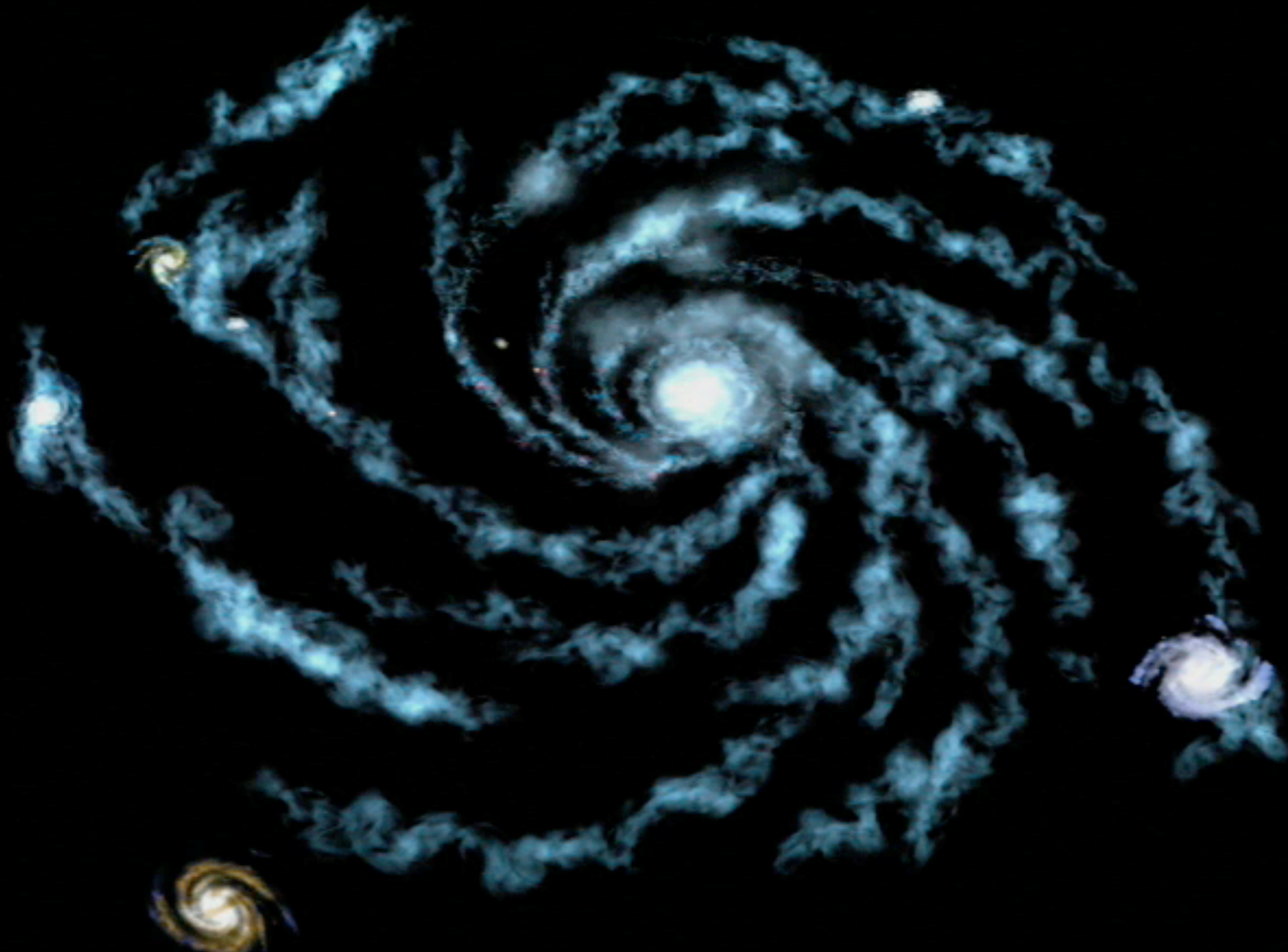
- Now we can **observe** the physical condition of the Universe when it was very young.

Cosmic Microwave Background (CMB)

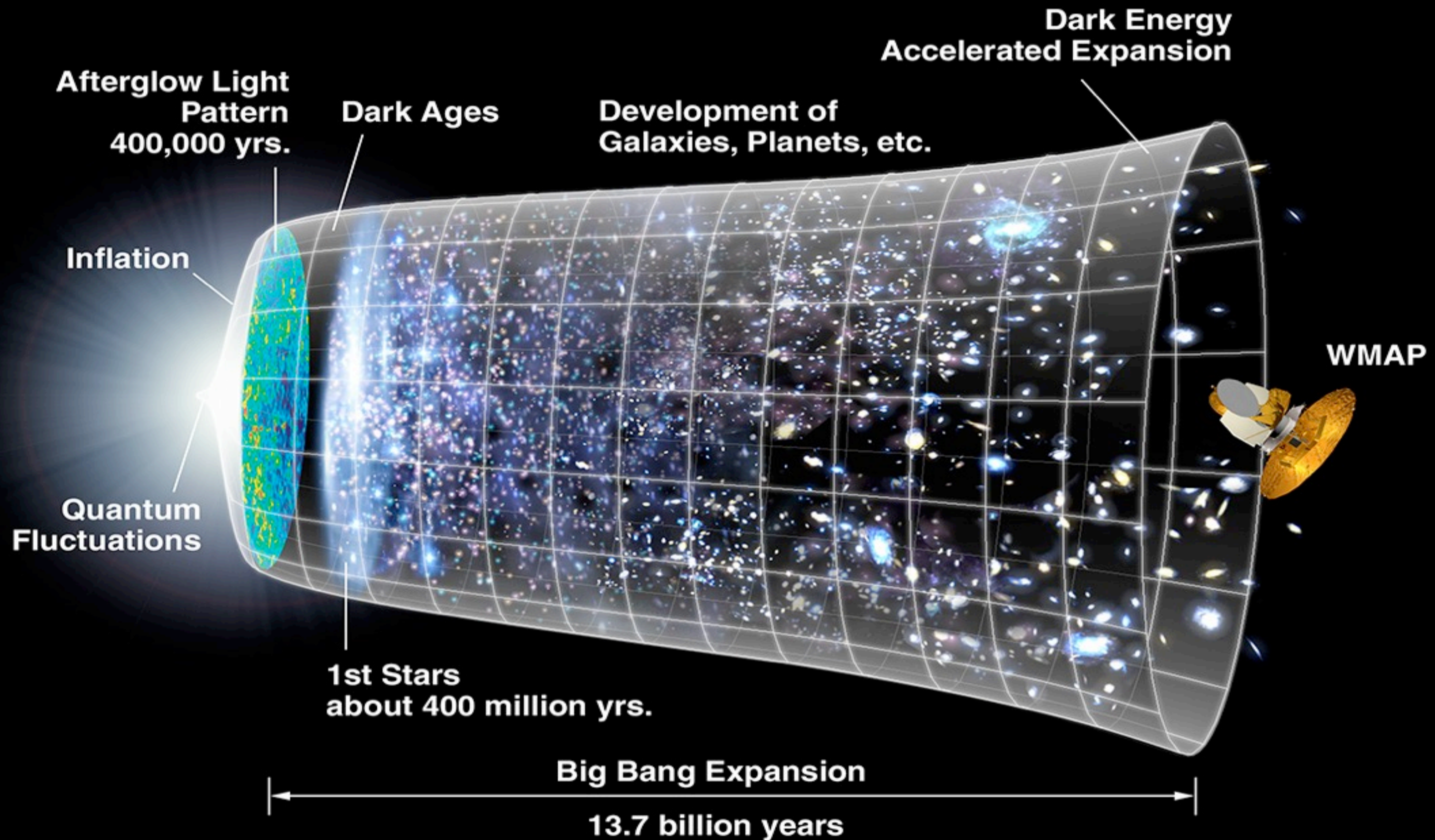
- Fossil light of the Big Bang!



From "Cosmic Voyage"



CMB: The Farthest and Oldest Light That We Can Ever Hope To Observe Directly



- When the Universe was 3000K (~380,000 years after the Big Bang), electrons and protons were combined to form neutral hydrogen.

used to be WMAP at Lagrange 2 (L2) Point

June 2001:
WMAP launched!

February 2003:
The first-year data release

March 2006:
The three-year data release

March 2008:
The five-year data release

January 2010:
The seven-year data release

*September 8, 2010:
WMAP left L2*



- L2 is 1.6 million kilometers from Earth
- WMAP leaves Earth, Moon, and Sun behind it to avoid radiation from them

used to be WMAP at Lagrange 2 (L2) Point

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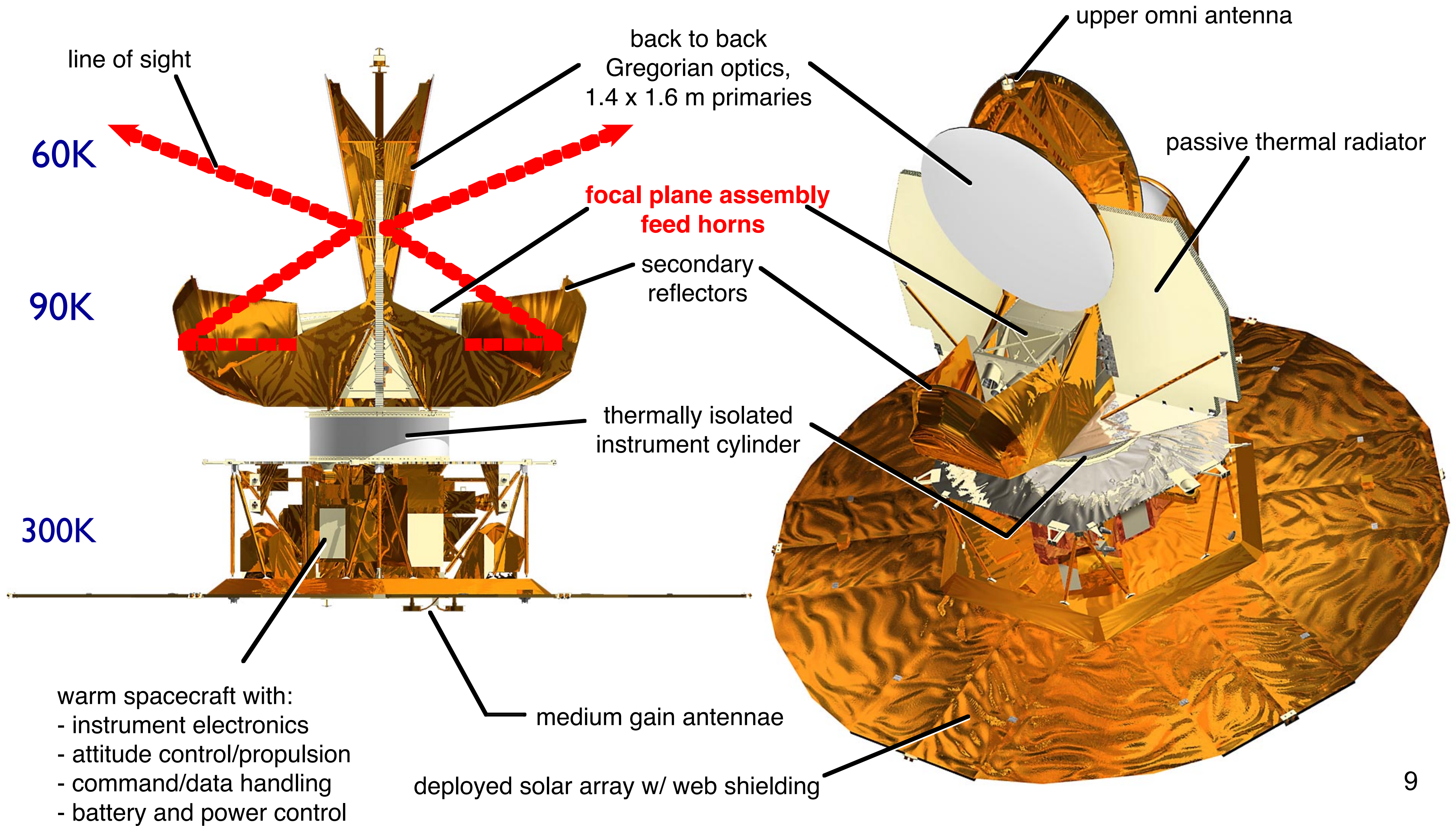
*September 8, 2010:
WMAP left L2*

**We are currently working
on the final data release!
(nine-year data release)**

- L2 is 1.6 million kilometers from Earth
- WMAP leaves Earth, Moon, and Sun behind it to avoid radiation from them

WMAP Spacecraft

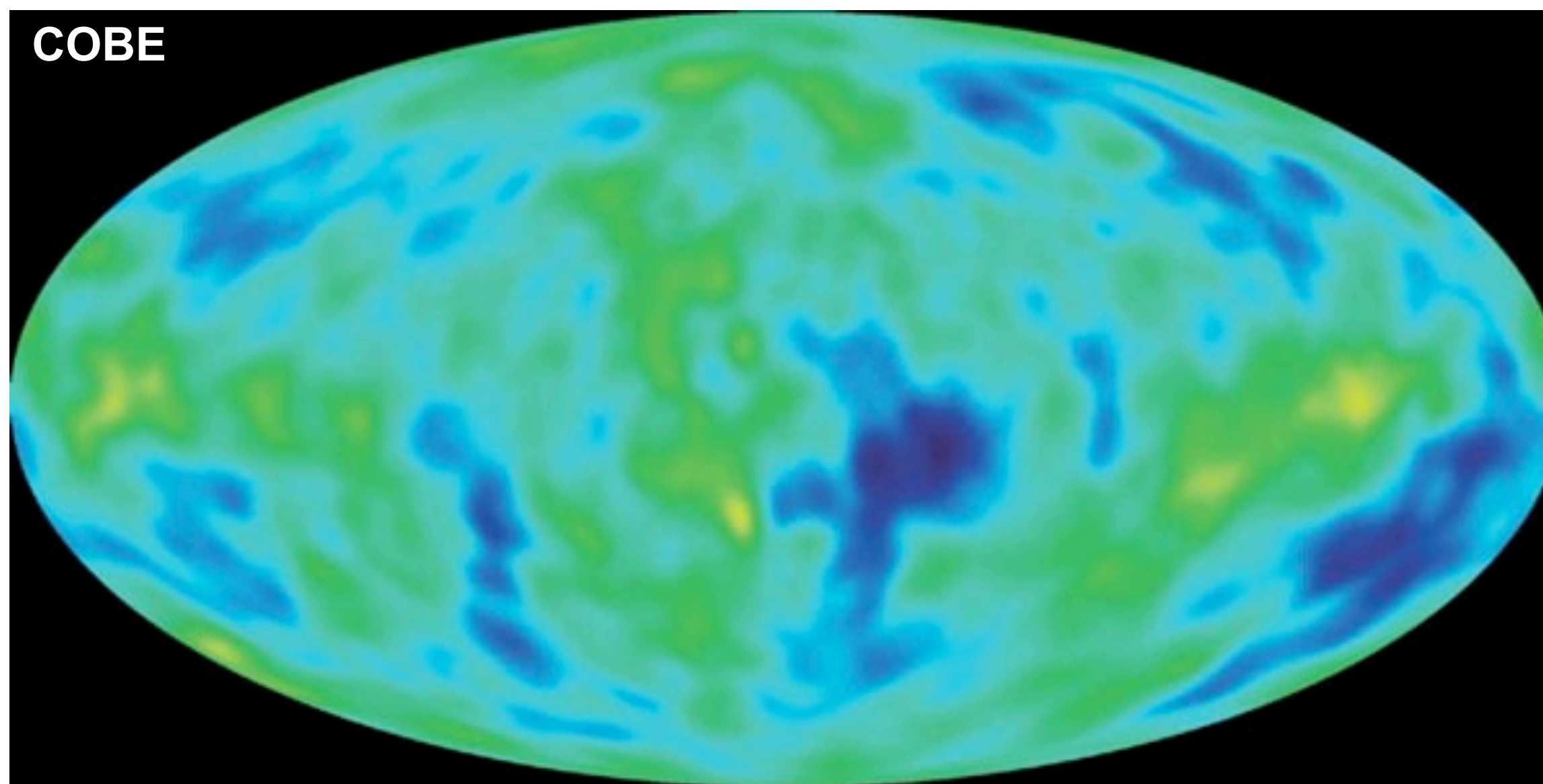
Radiative Cooling: No Cryogenic System



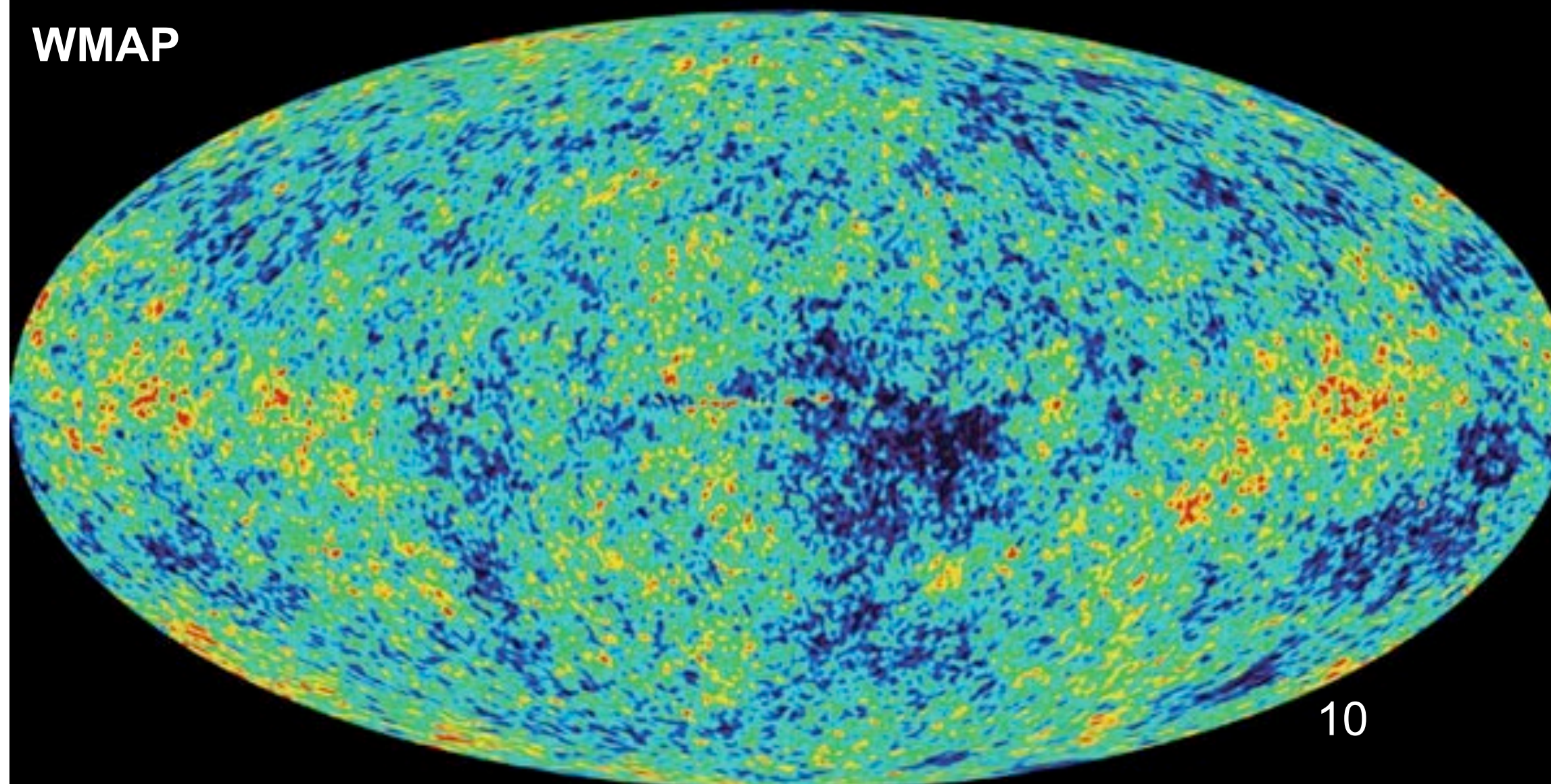
COBE to WMAP (x35 better resolution)



COBE
1989



WMAP
2001



WMAP 7-Year Science Team

- C.L. Bennett
- G. Hinshaw
- N. Jarosik
- S.S. Meyer
- L. Page
- D.N. Spergel
- E.L. Wright
- M.R. Greason
- M. Halpern
- R.S. Hill
- A. Kogut
- M. Limon
- N. Odegard
- G.S. Tucker
- J. L. Weiland
- E. Wollack
- J. Dunkley
- B. Gold
- E. Komatsu
- D. Larson
- M.R.olta
- K.M. Smith
- C. Barnes
- R. Bean
- O. Dore
- H.V. Peiris
- L. Verde

WMAP 7-Year Papers

- **Jarosik et al.**, “*Sky Maps, Systematic Errors, and Basic Results*” [Astrophysical Journal Supplement Series \(ApJS\), 192, 14 \(2011\)](#)
- **Gold et al.**, “*Galactic Foreground Emission*” [ApJS, 192, 15 \(2011\)](#)
- **Weiland et al.**, “*Planets and Celestial Calibration Sources*” [ApJS, 192, 19 \(2011\)](#)
- **Bennett et al.**, “*Are There CMB Anomalies?*” [ApJS, 192, 17 \(2011\)](#)
- **Larson et al.**, “*Power Spectra and WMAP-Derived Parameters*” [ApJS, 192, 16 \(2011\)](#)
- **Komatsu et al.**, “*Cosmological Interpretation*” [ApJS, 192, 18 \(2011\)](#)

Cosmic Pie Chart: 7-year

● Standard Model

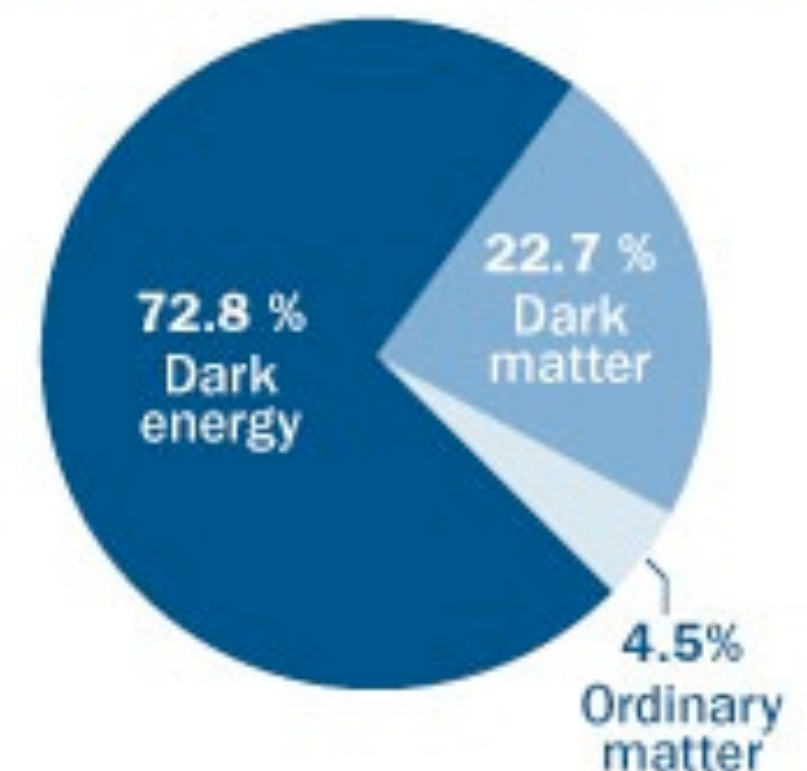
- H&He = **4.58%** ($\pm 0.16\%$)
- Dark Matter = **22.9%** ($\pm 1.5\%$)
- Dark Energy = **72.5%** ($\pm 1.6\%$)
- $H_0 = 70.2 \pm 1.4$ km/s/Mpc
- Age of the Universe = 13.76 billion years (± 0.11 billion years)

Universal Stats

Age of the universe today
13.75 billion years

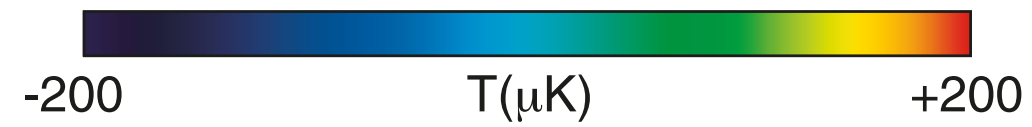
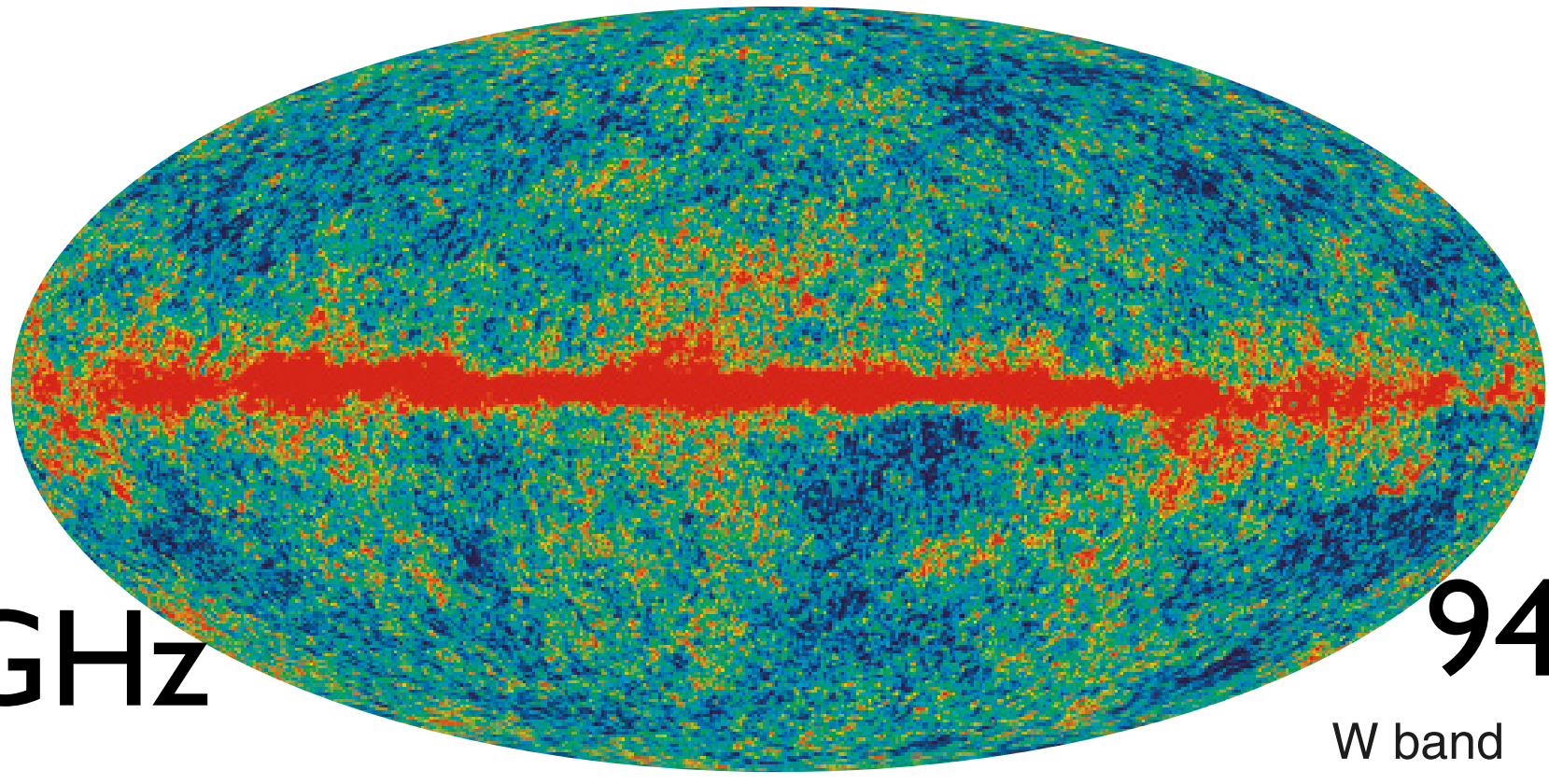
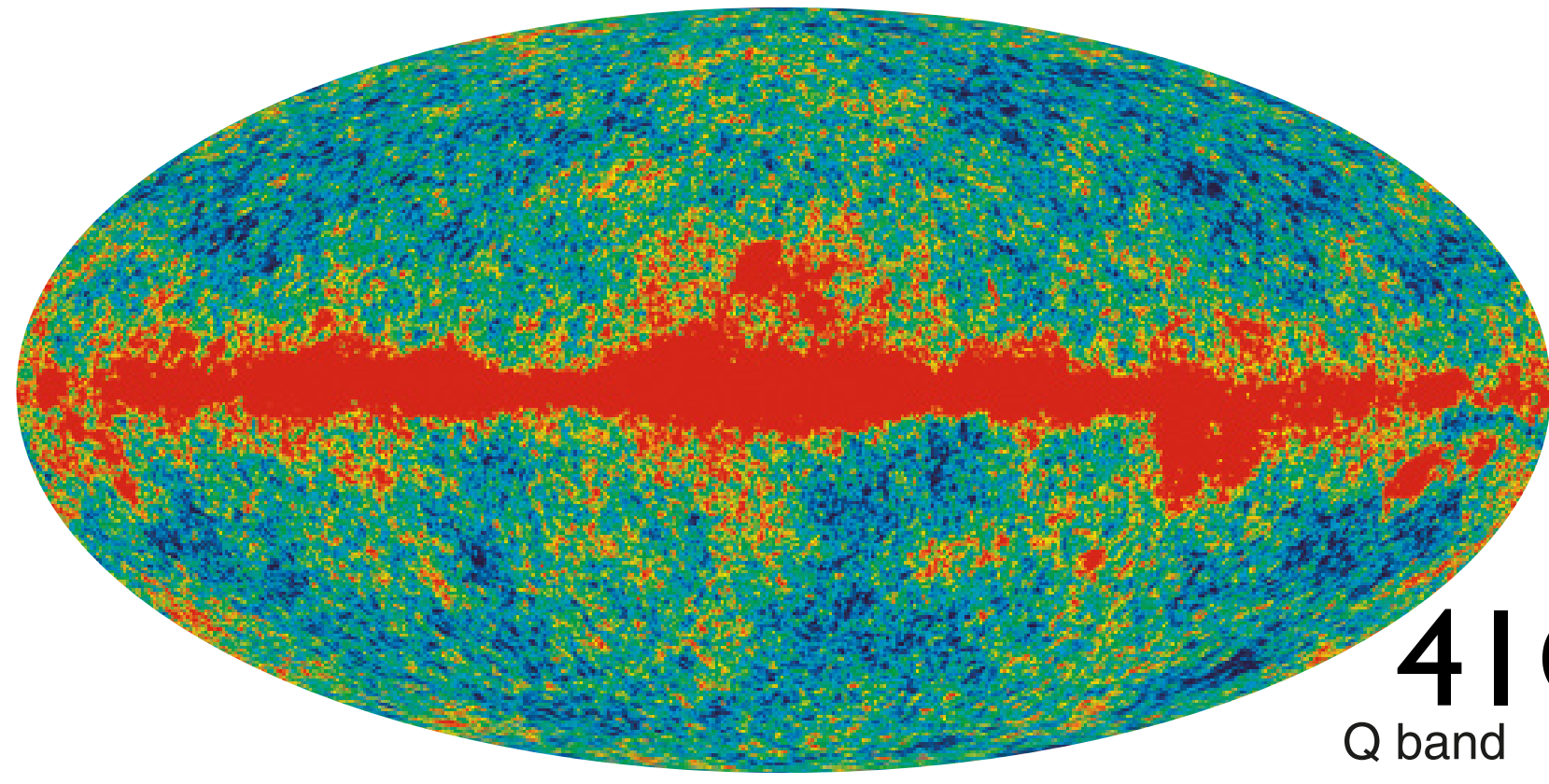
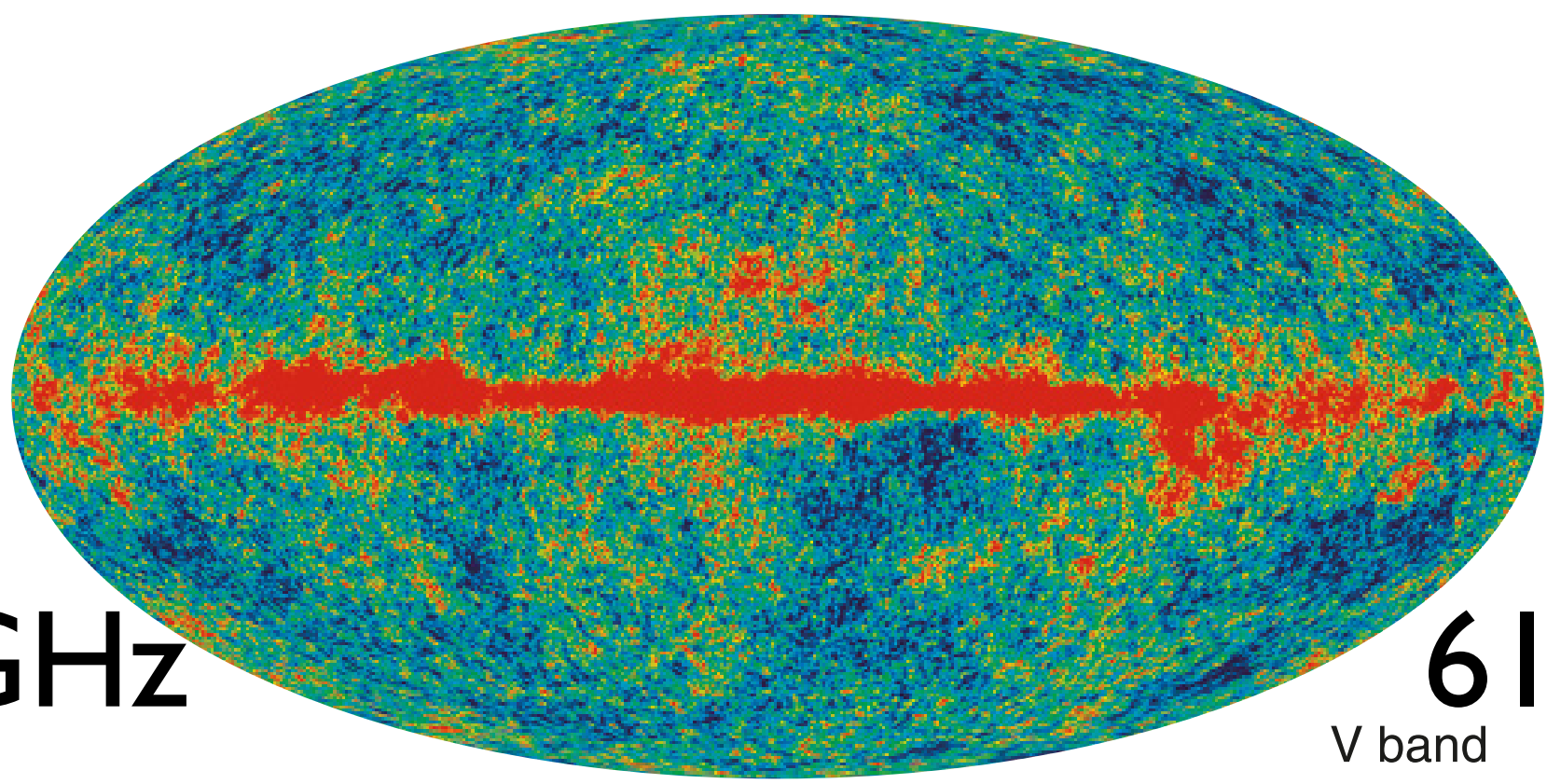
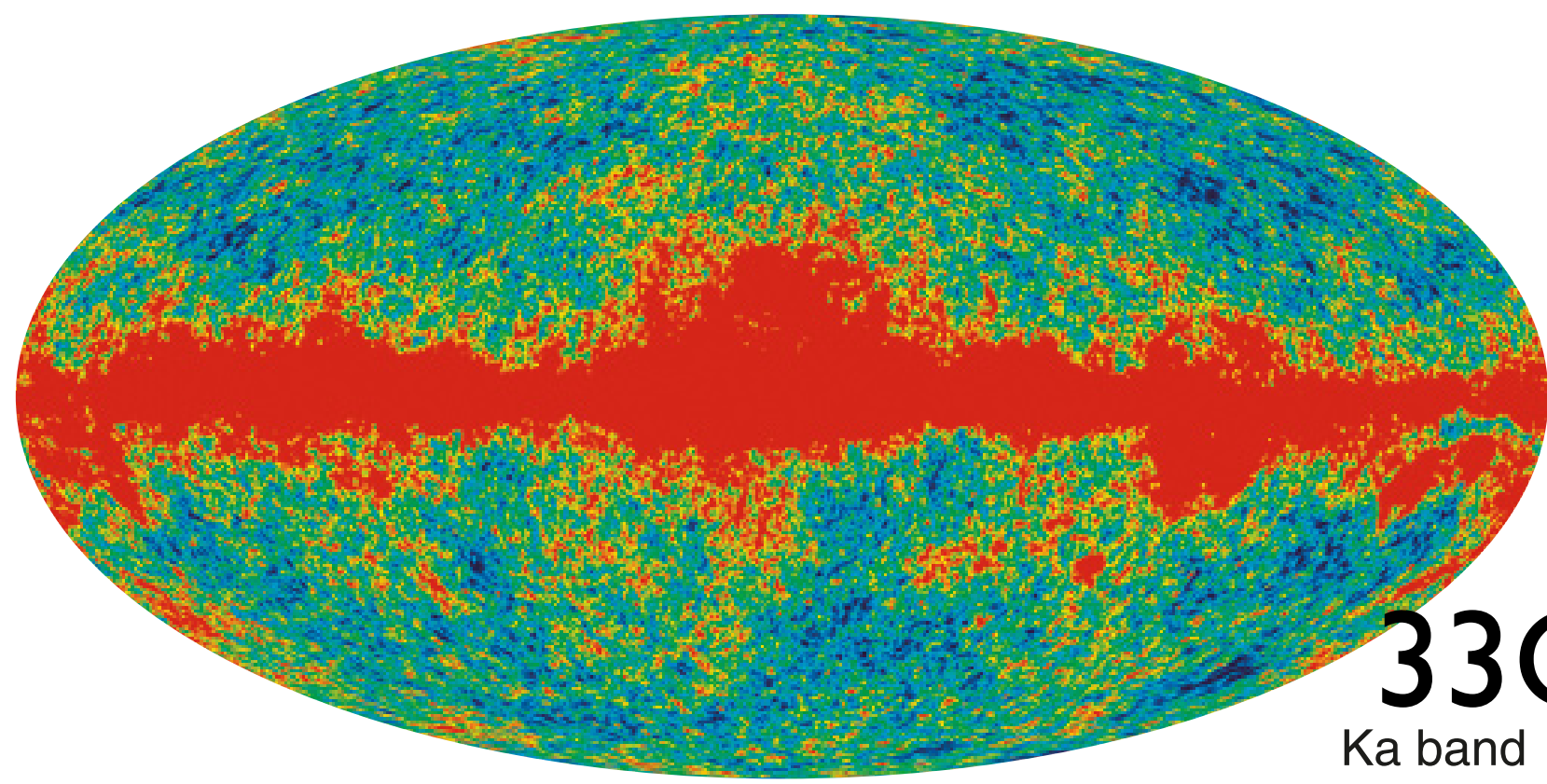
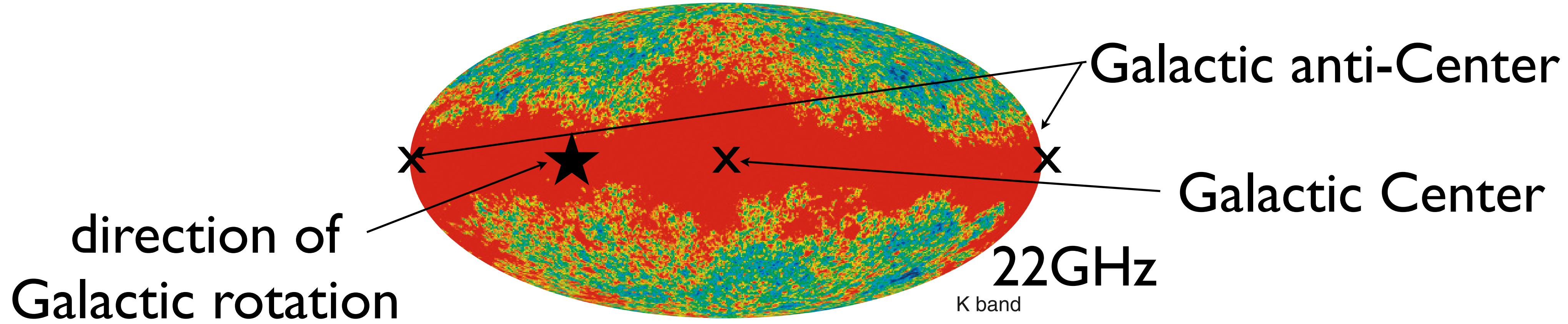
Age of the cosmos at
time of reionization
457 million years

Universe composition

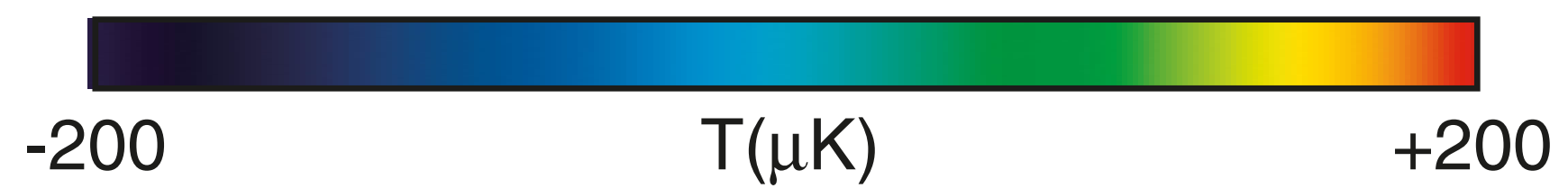
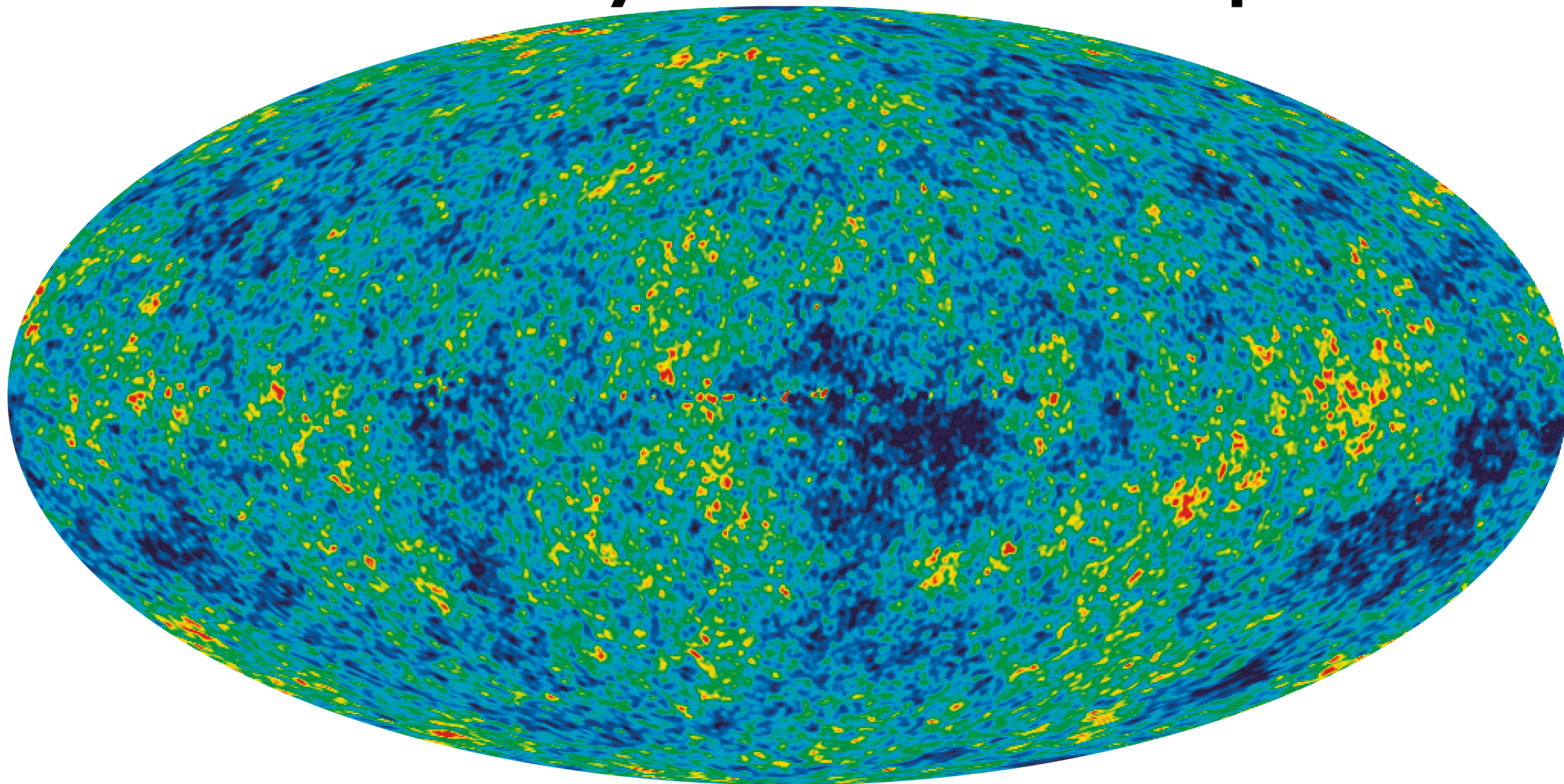


*“ScienceNews” article on
the WMAP 7-year results*

How did we obtain these numbers?

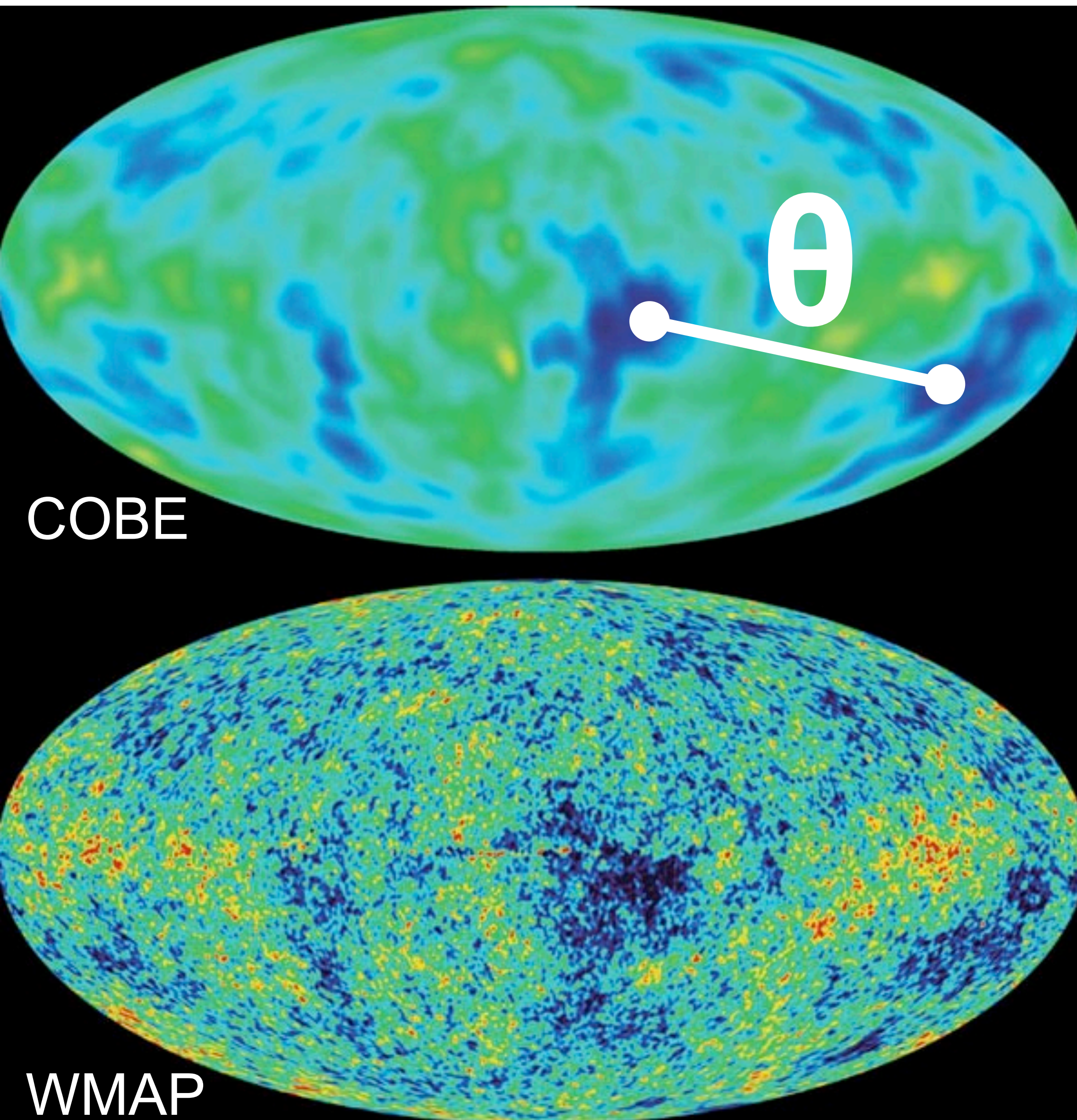


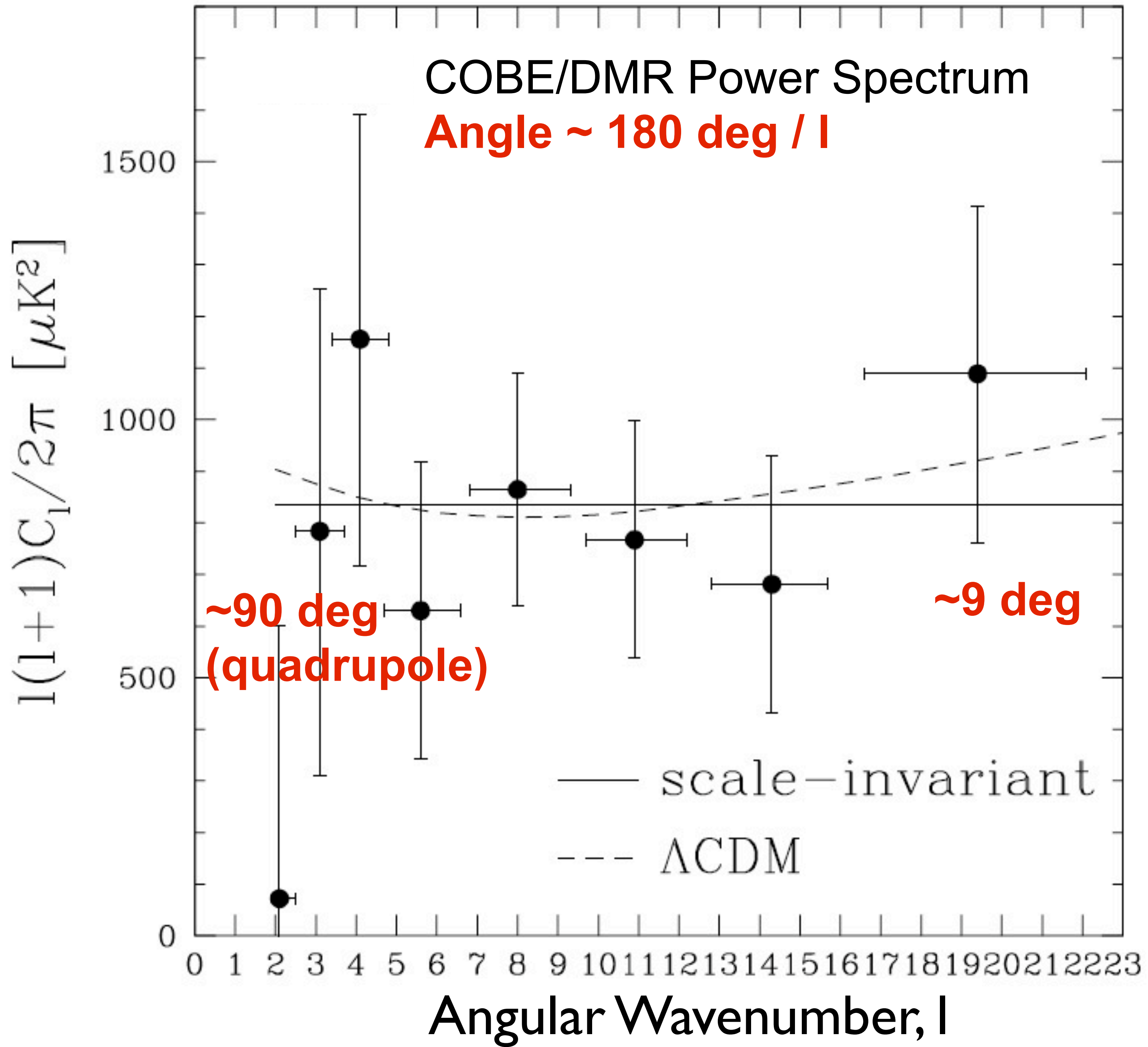
Galaxy-cleaned Map



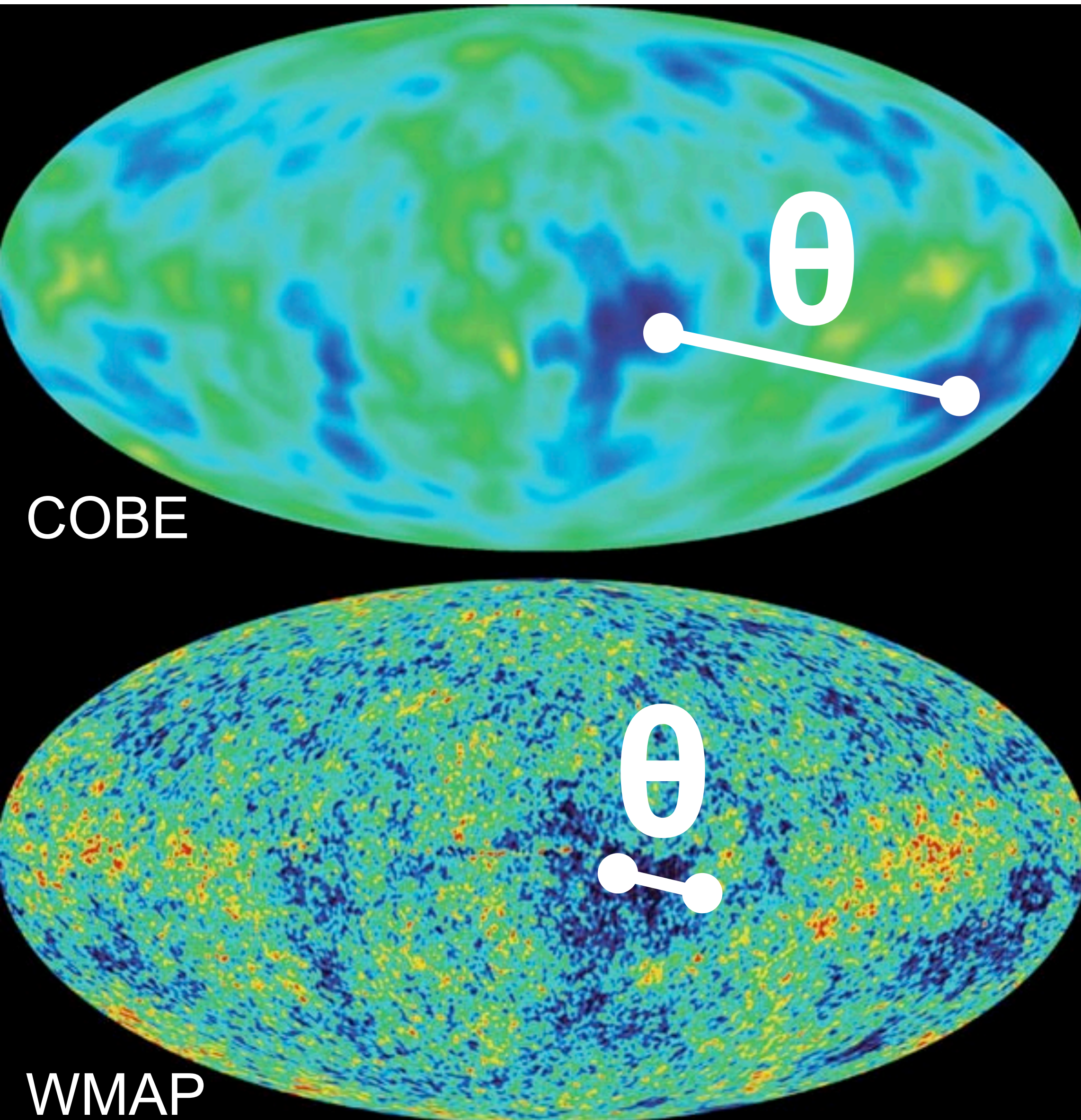
Analysis: 2-point Correlation

- $C(\theta) = (1/4\pi) \sum (2l+1) C_l P_l(\cos\theta)$
- How are temperatures on two points on the sky, separated by θ , are correlated?
- **“Power Spectrum,”** C_l
 - How much fluctuation power do we have at a given angular scale?
 - $l \sim 180 \text{ degrees} / \theta$



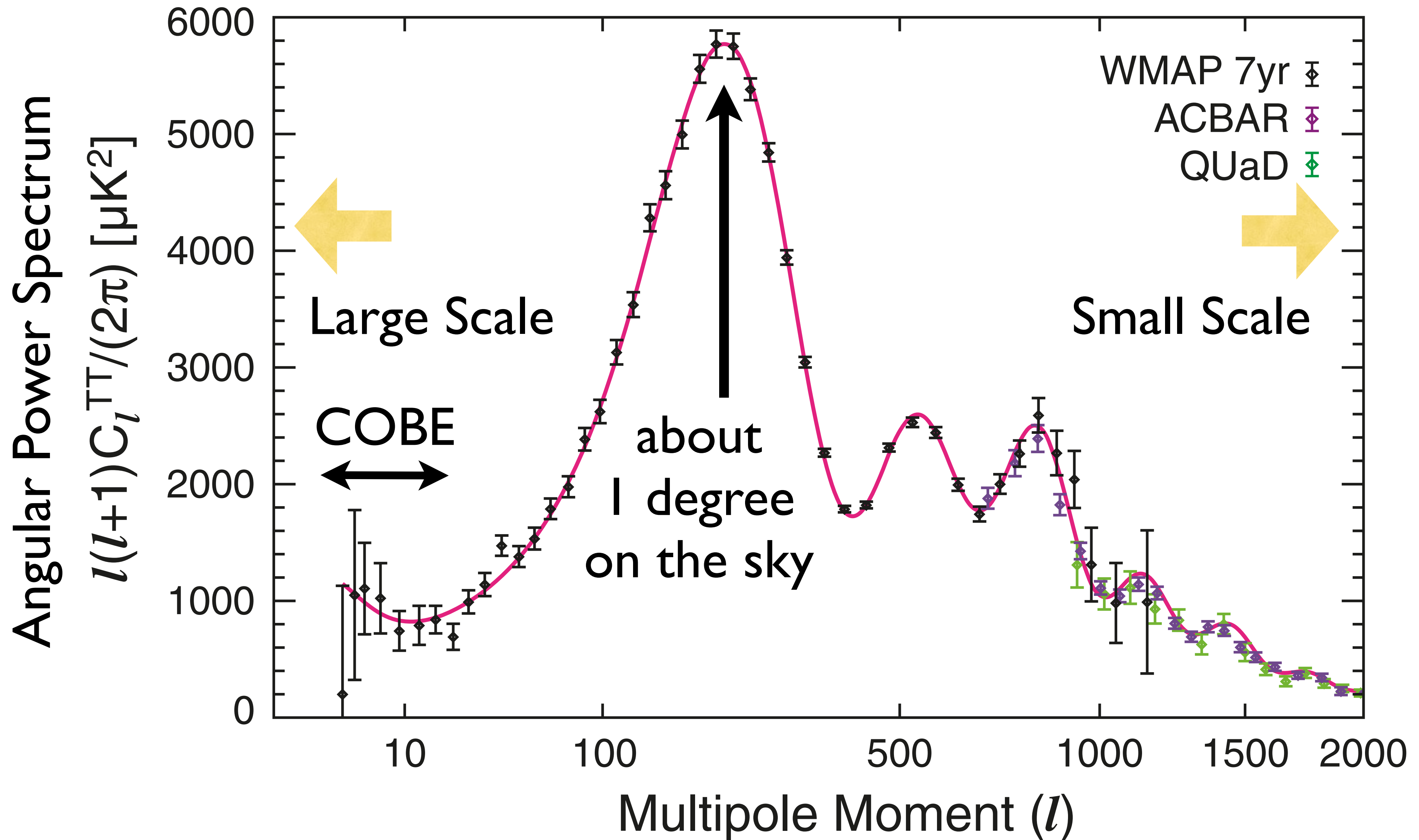


COBE To WMAP

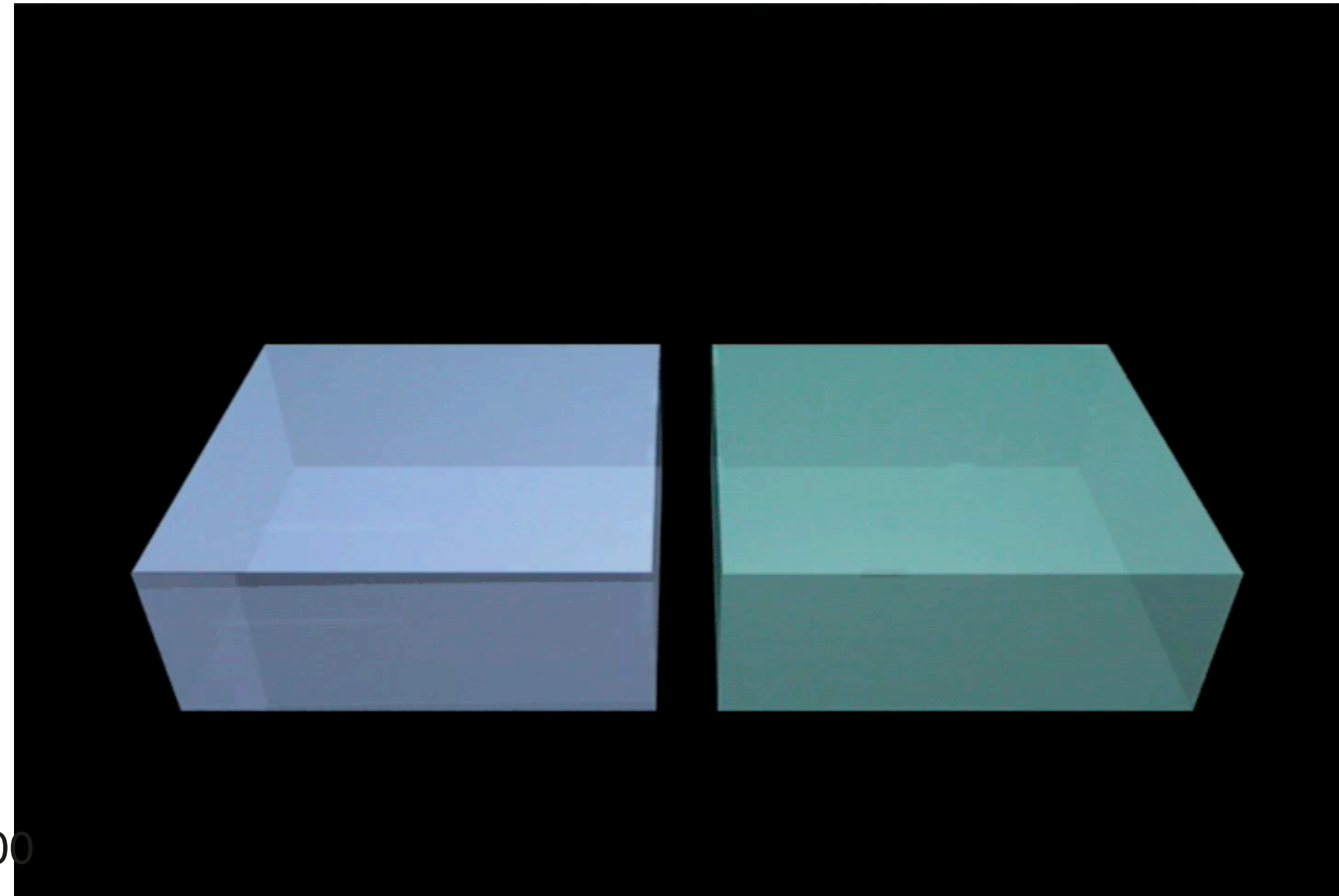
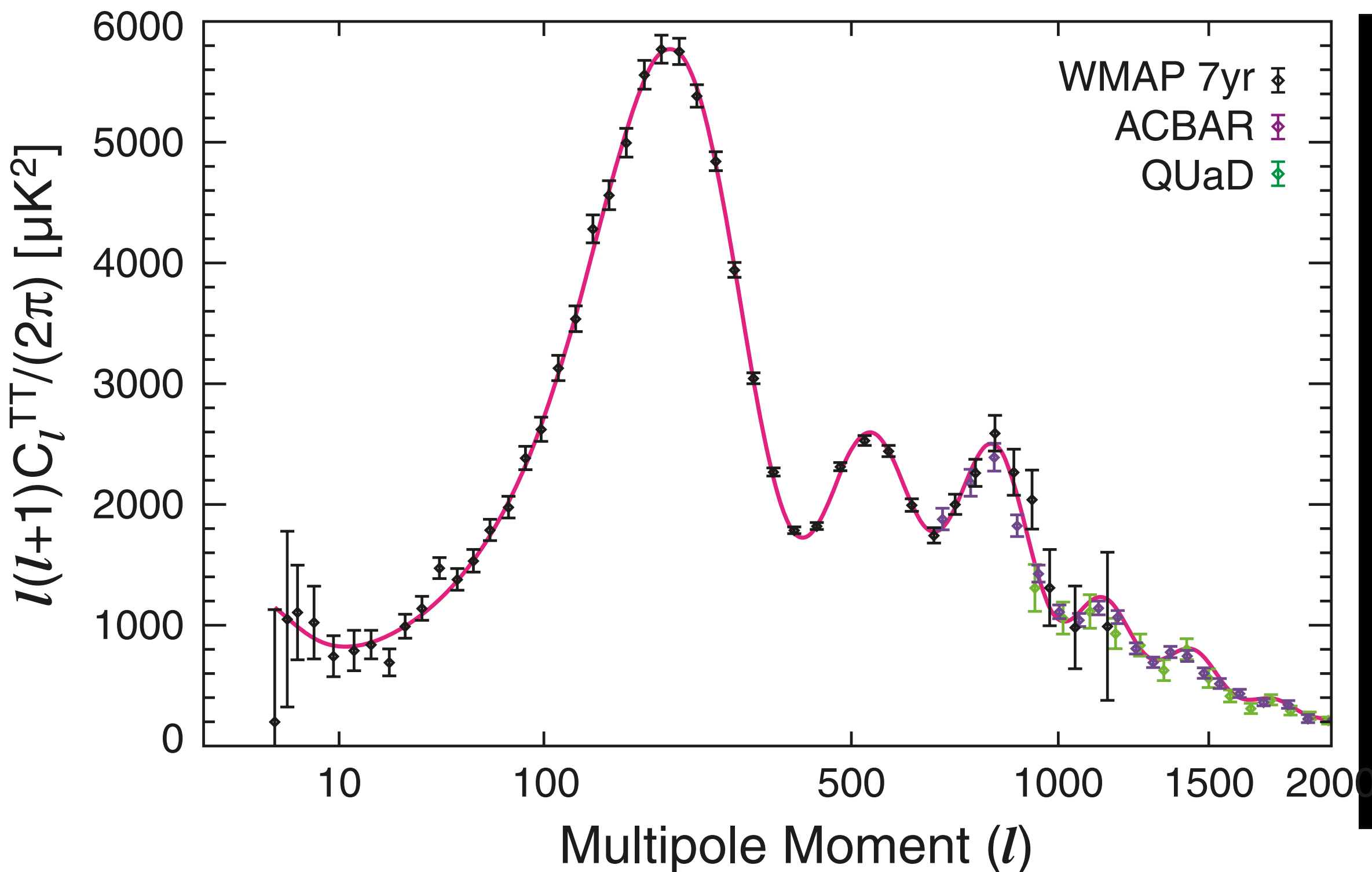


- COBE is unable to resolve the structures below ~ 7 degrees
- WMAP's resolving power is 35 times better than COBE.
- What did WMAP see?

WMAP Power Spectrum

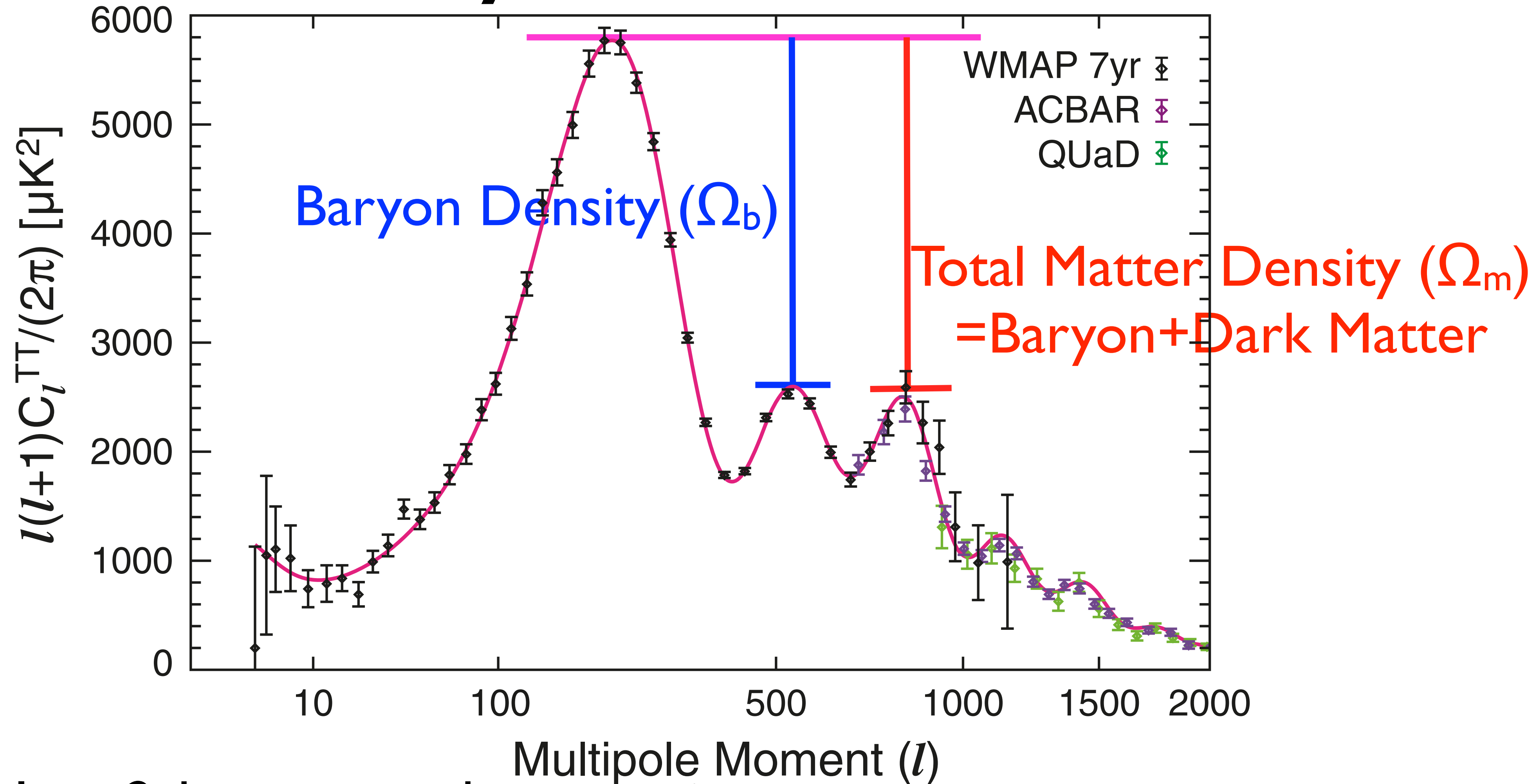


The Cosmic Sound Wave



- “*The Universe as a Miso soup*”
- *Main Ingredients: protons, helium nuclei, electrons, photons*
- We measure the composition of the Universe by analyzing the wave form of the cosmic sound waves.

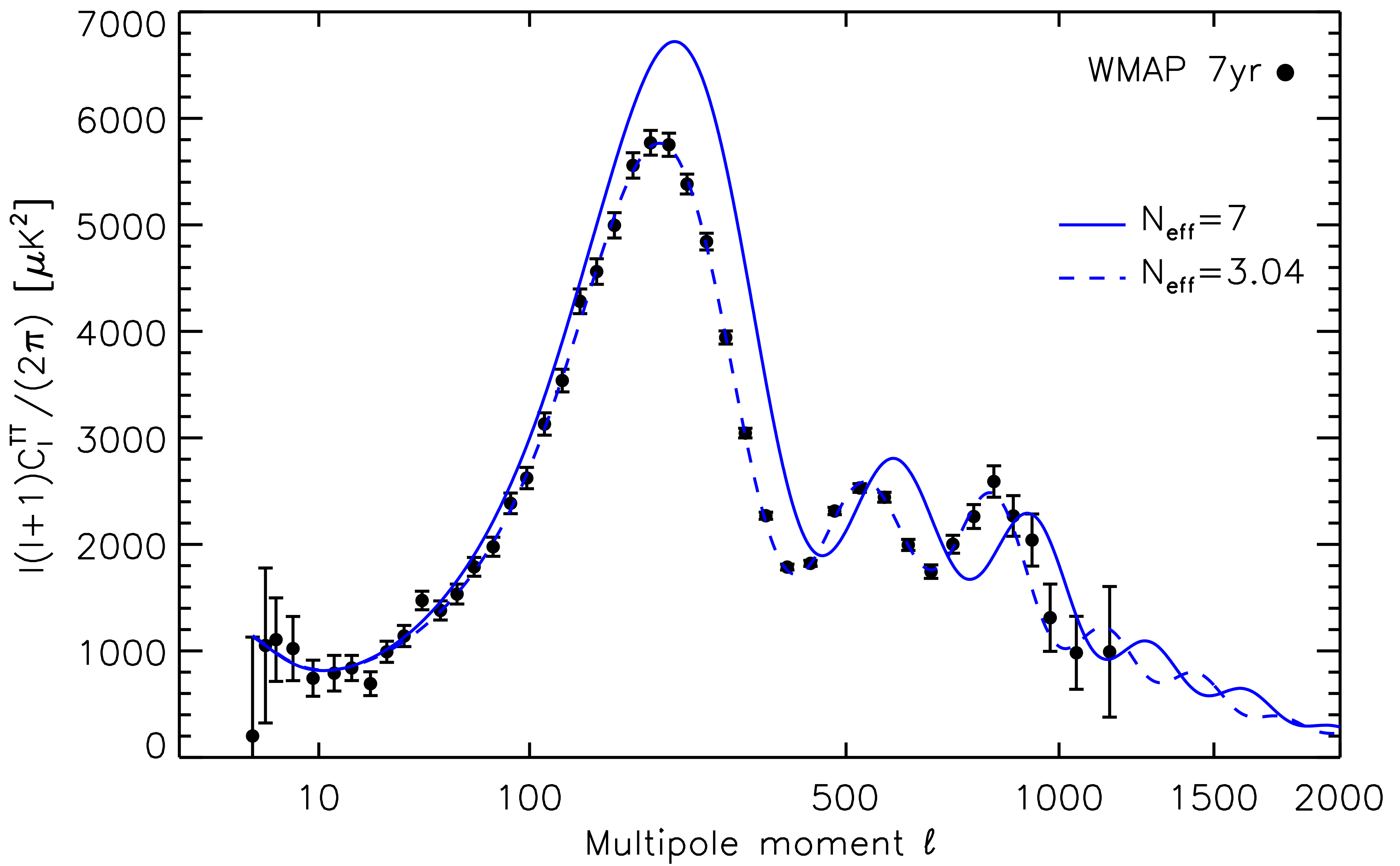
CMB to Baryon & Dark Matter

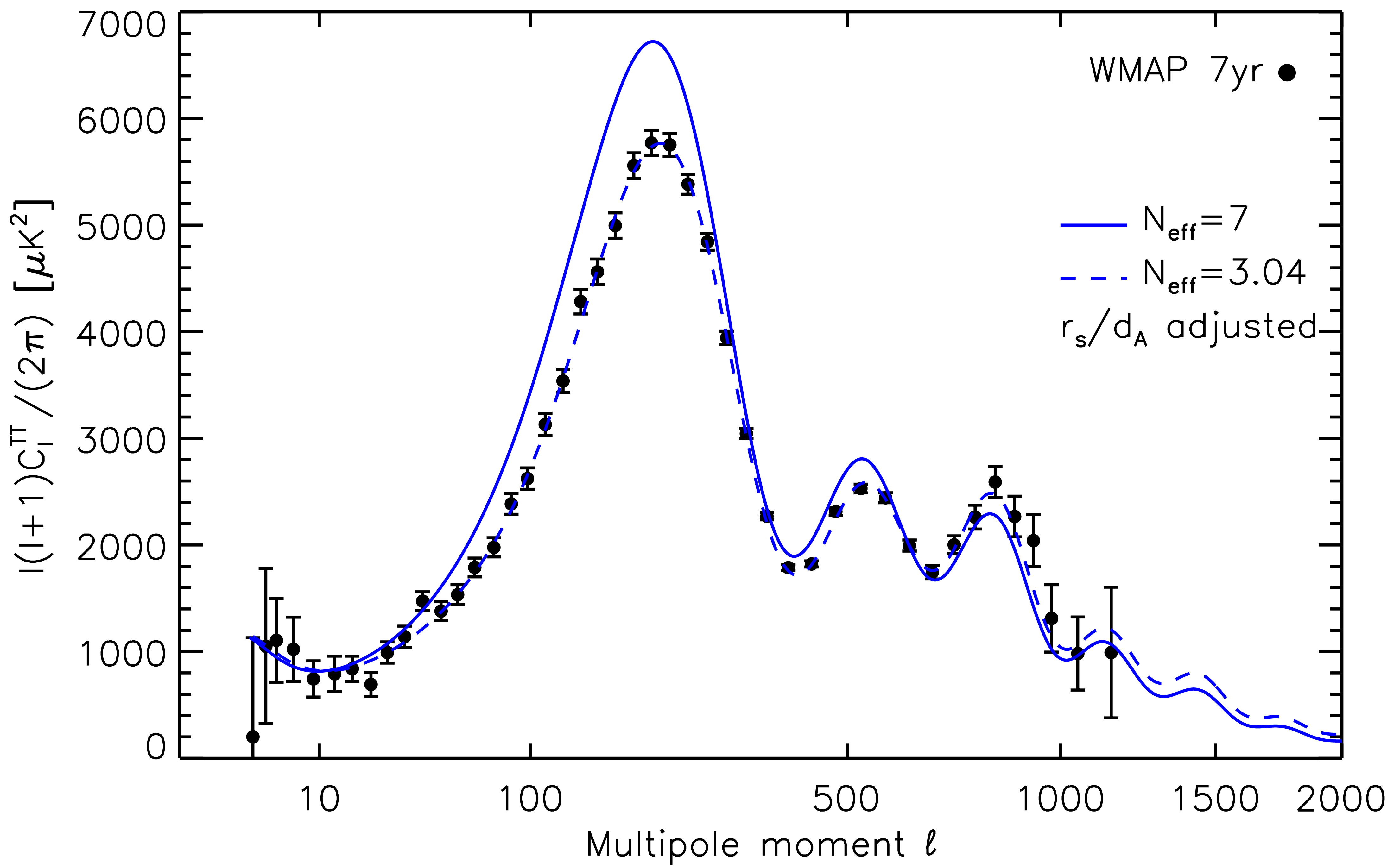


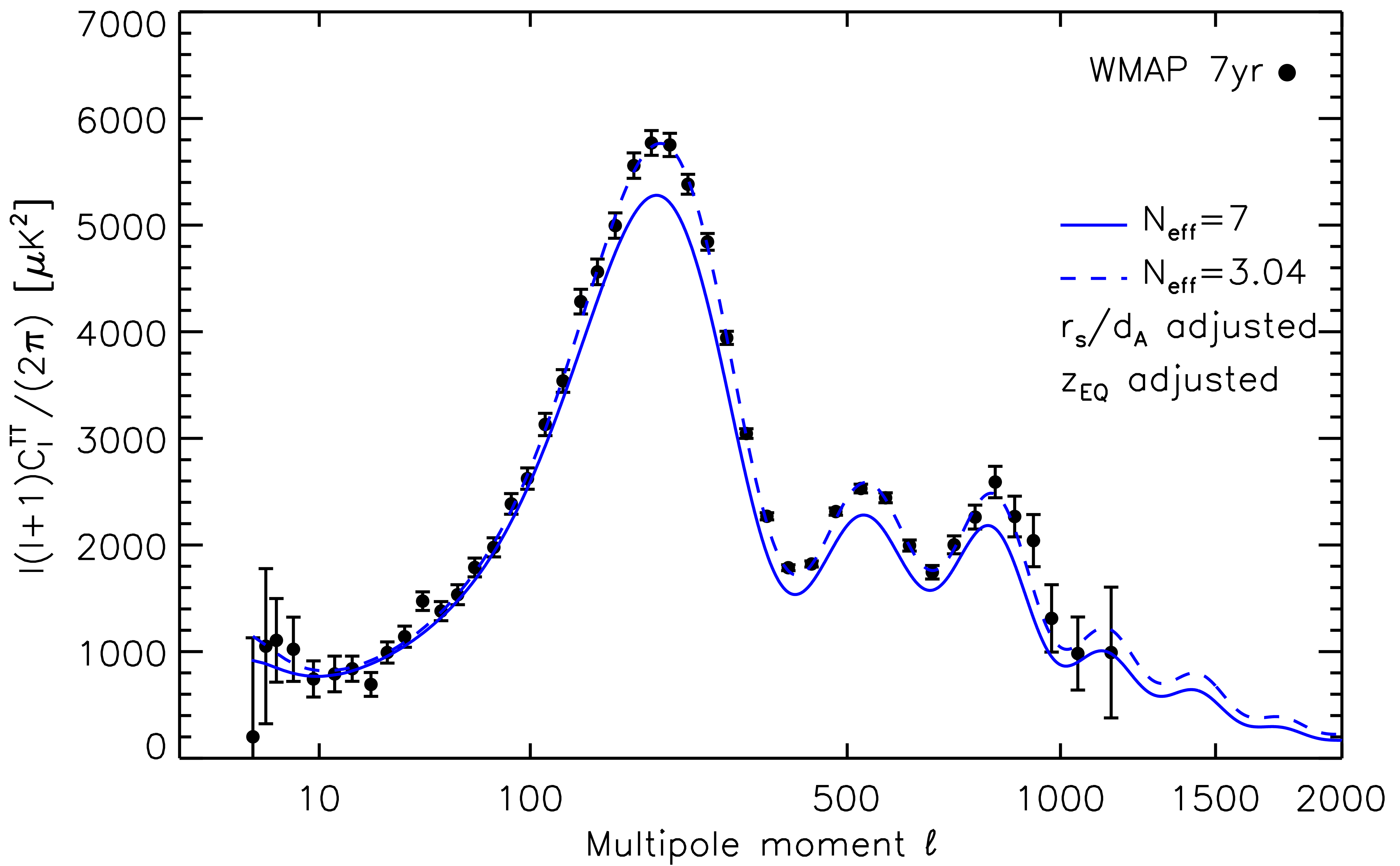
- l -to- 2 : baryon-to-photon ratio
- l -to- 3 : matter-to-radiation ratio (z_{EQ} : equality redshift)

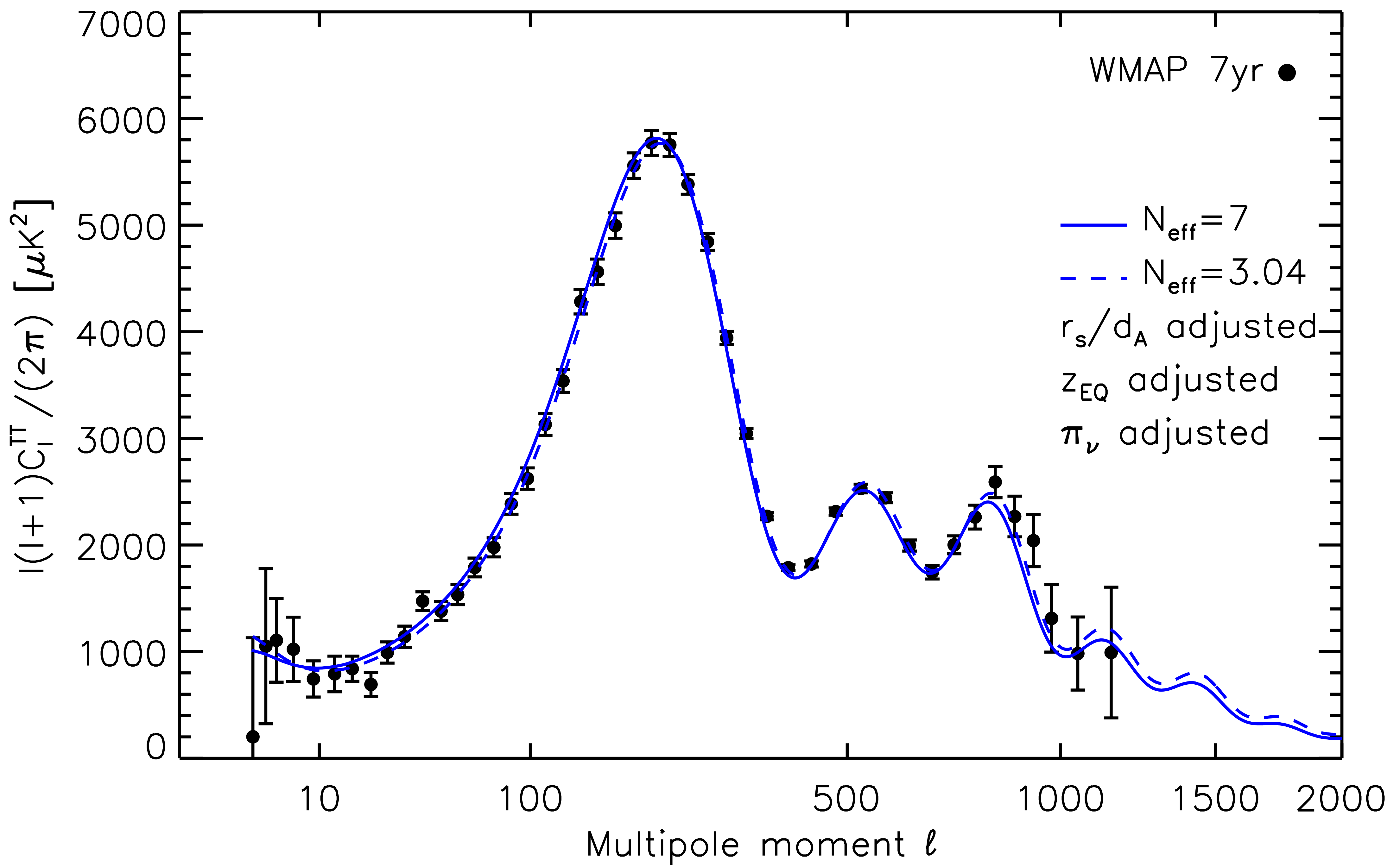
3rd-peak “Spectroscopy”

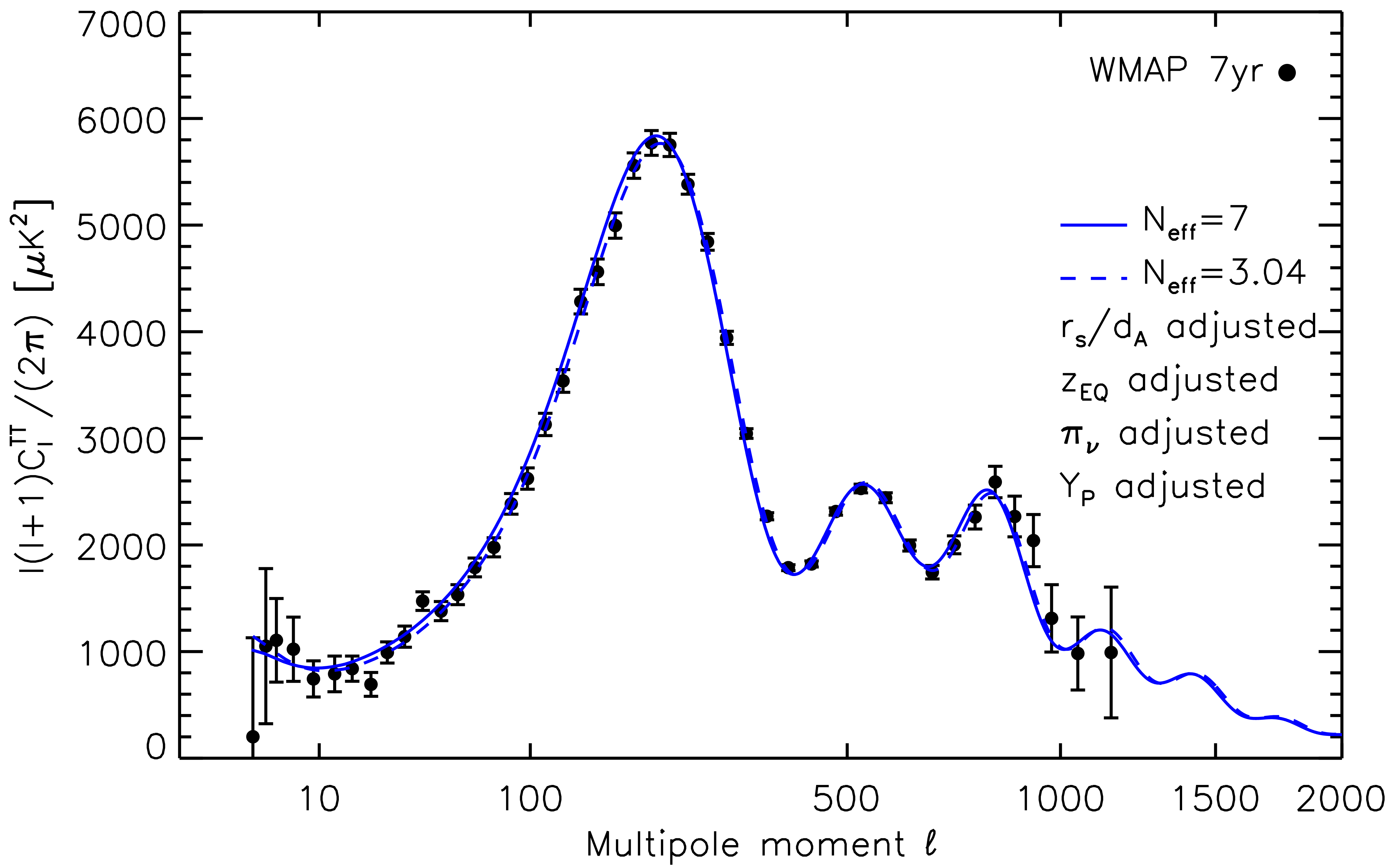
- Total Matter = Baryons (H&He) + Dark Matter
- Total Radiation = Photons + Neutrinos (+new radiation)
 - Neutrino temperature = $(4/11)^{1/3}$ Photon temperature
- So, for a given assumed value of the number of neutrino species (or the number of new radiation species, i.e., zero), we can measure the dark matter density.
- Or, we can get the dark matter density from elsewhere, and determine the number of radiation species!



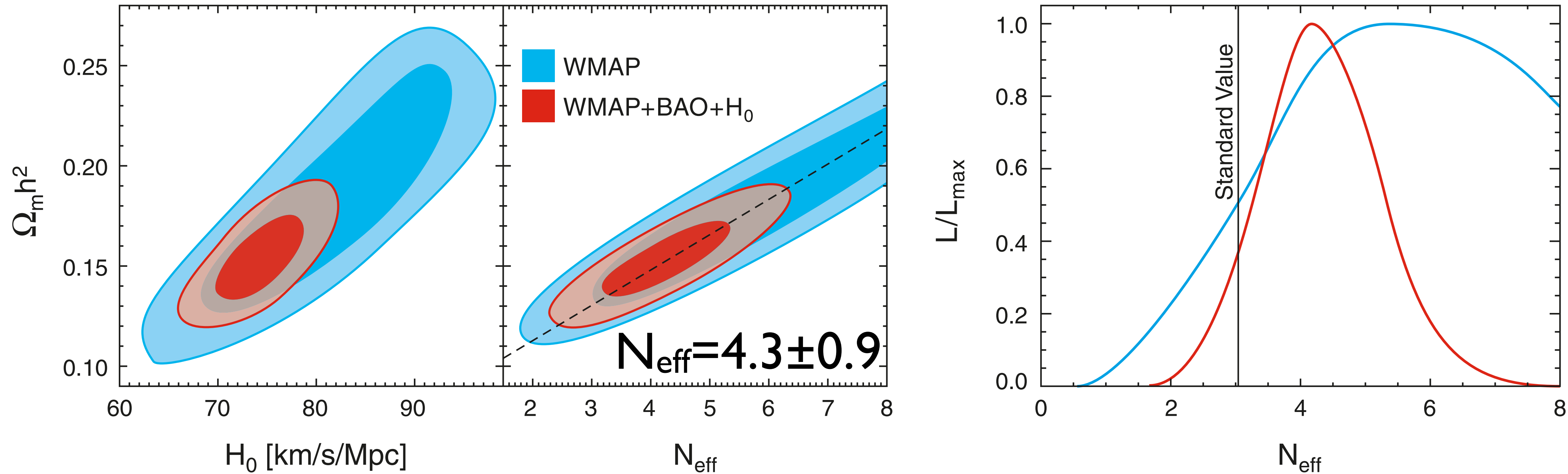






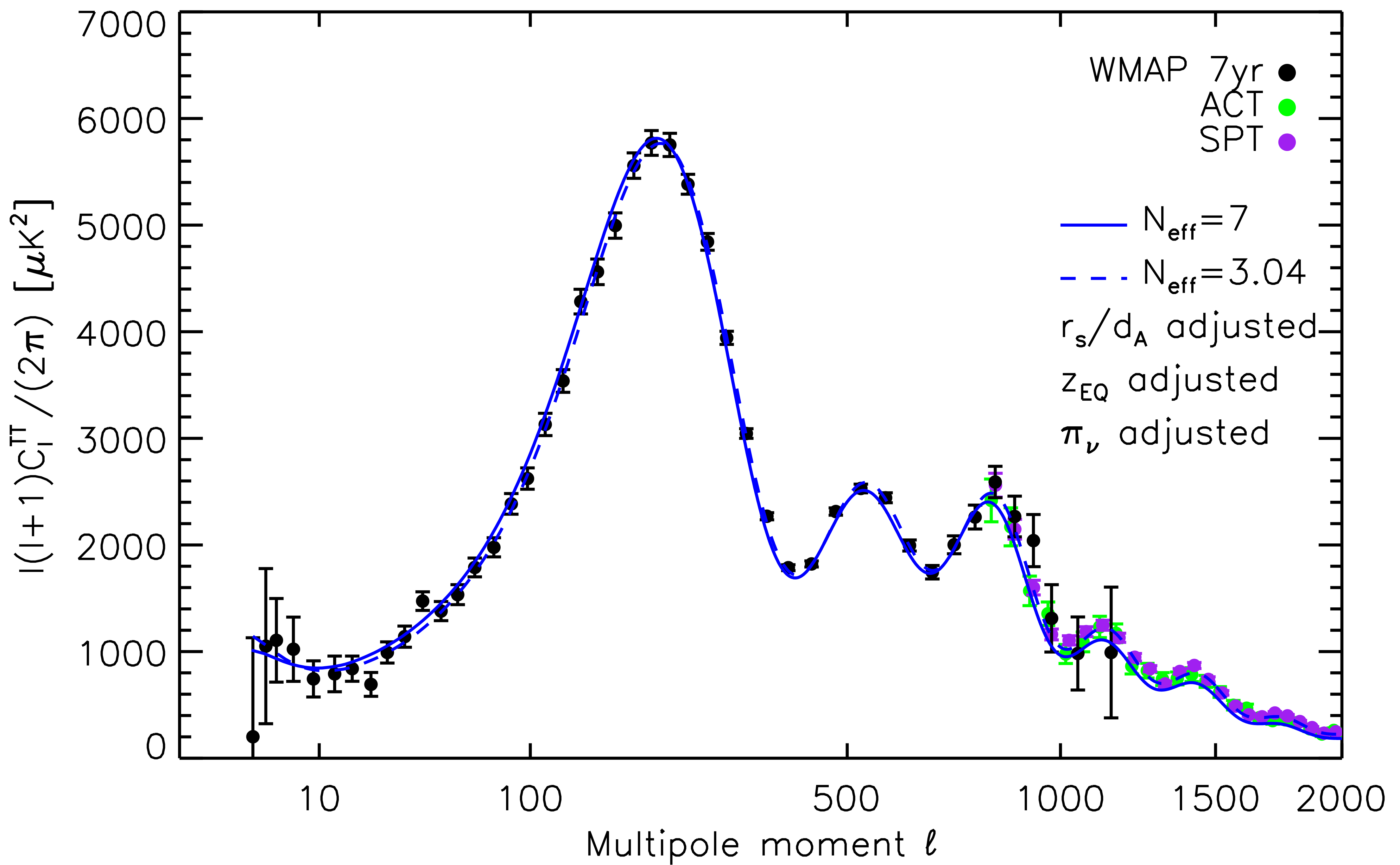


“3rd peak spectroscopy”: Number of Relativistic Species

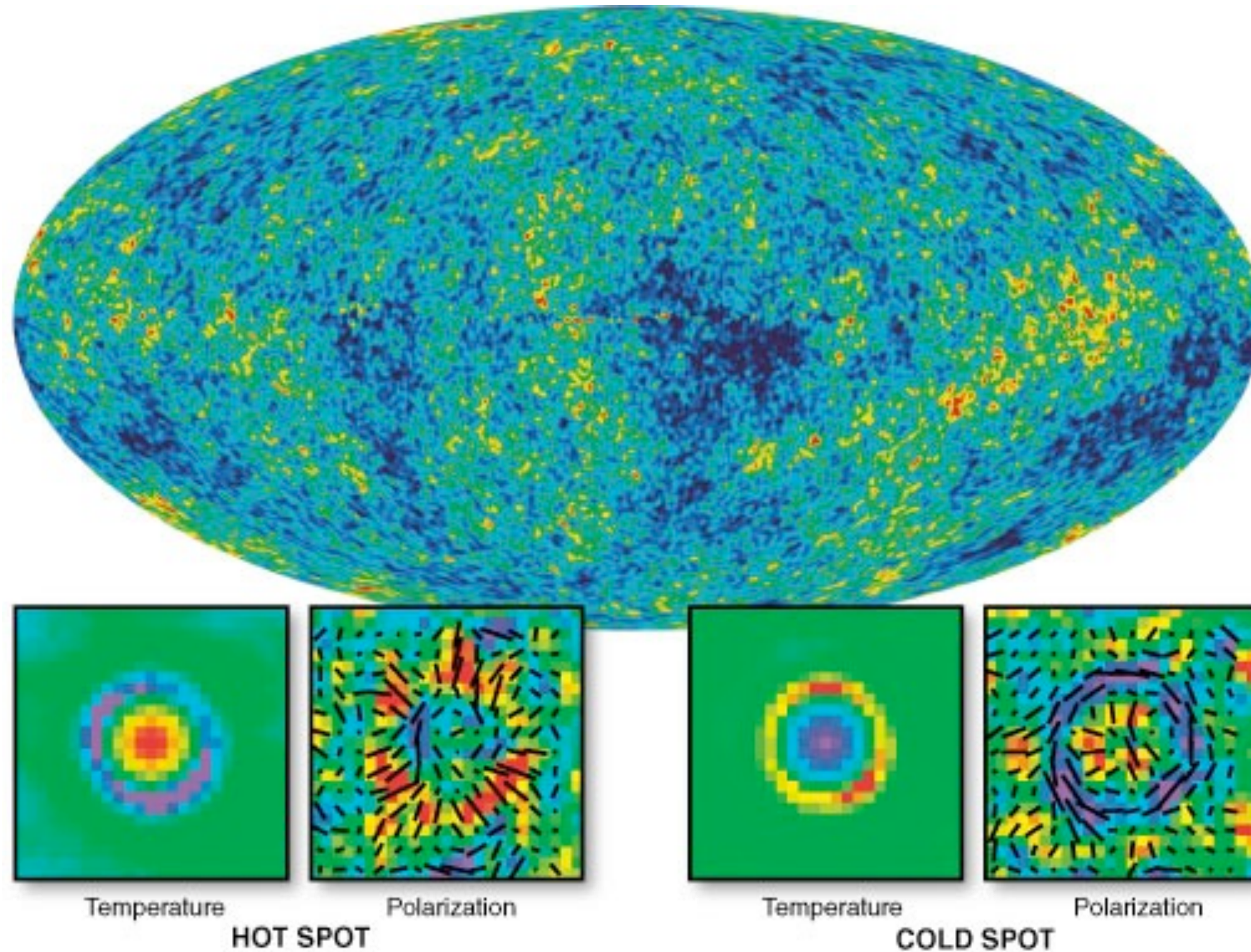


$$N_{\text{eff}} = 3.04 + 7.44 \left(\frac{\Omega_m h^2}{0.1308} \frac{3139}{1 + z_{\text{eq}}} - 1 \right)$$

← from external data
← from 3rd peak

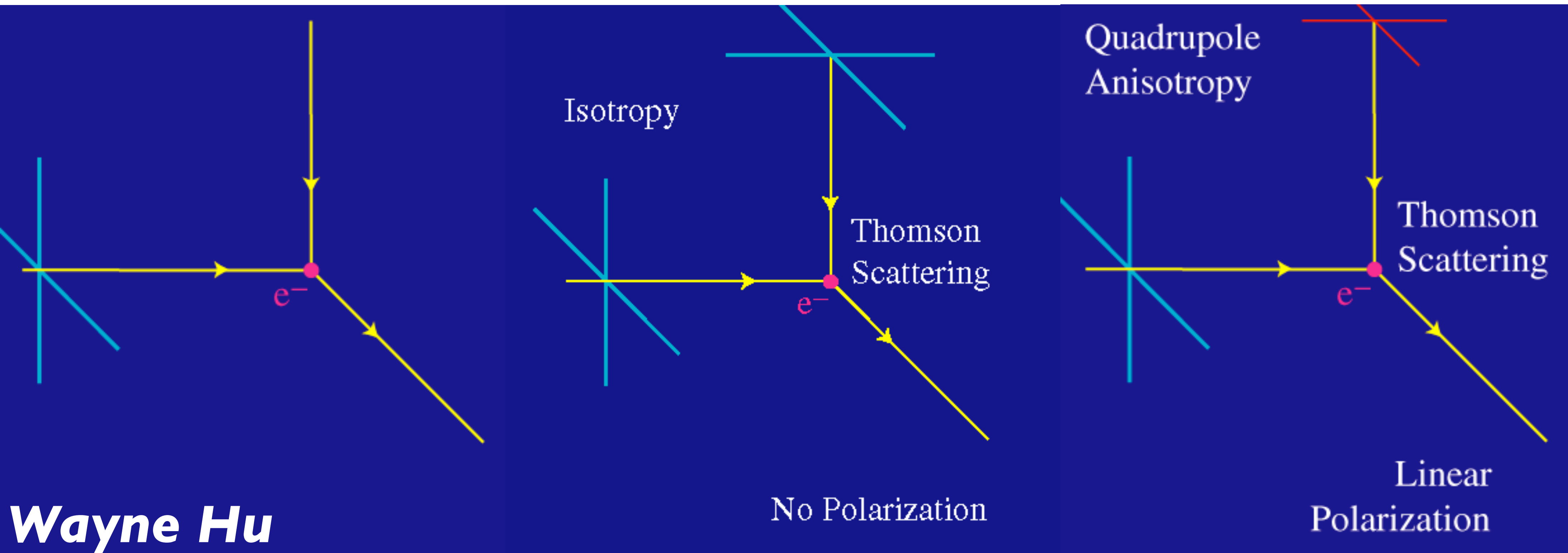


CMB Polarization



- *CMB is (very weakly) polarized!*

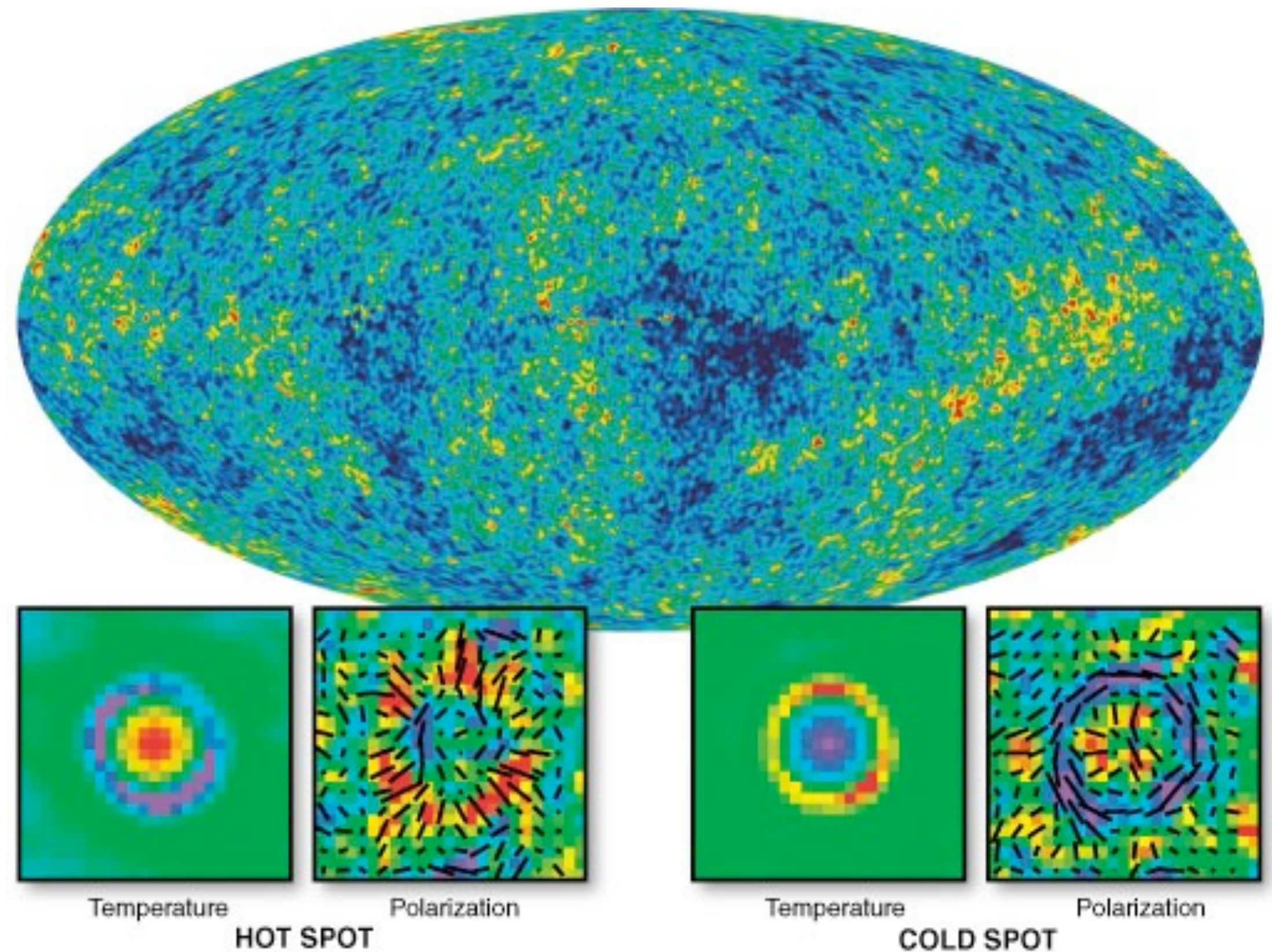
Physics of CMB Polarization



- CMB Polarization is created by a local temperature **quadrupole** anisotropy.

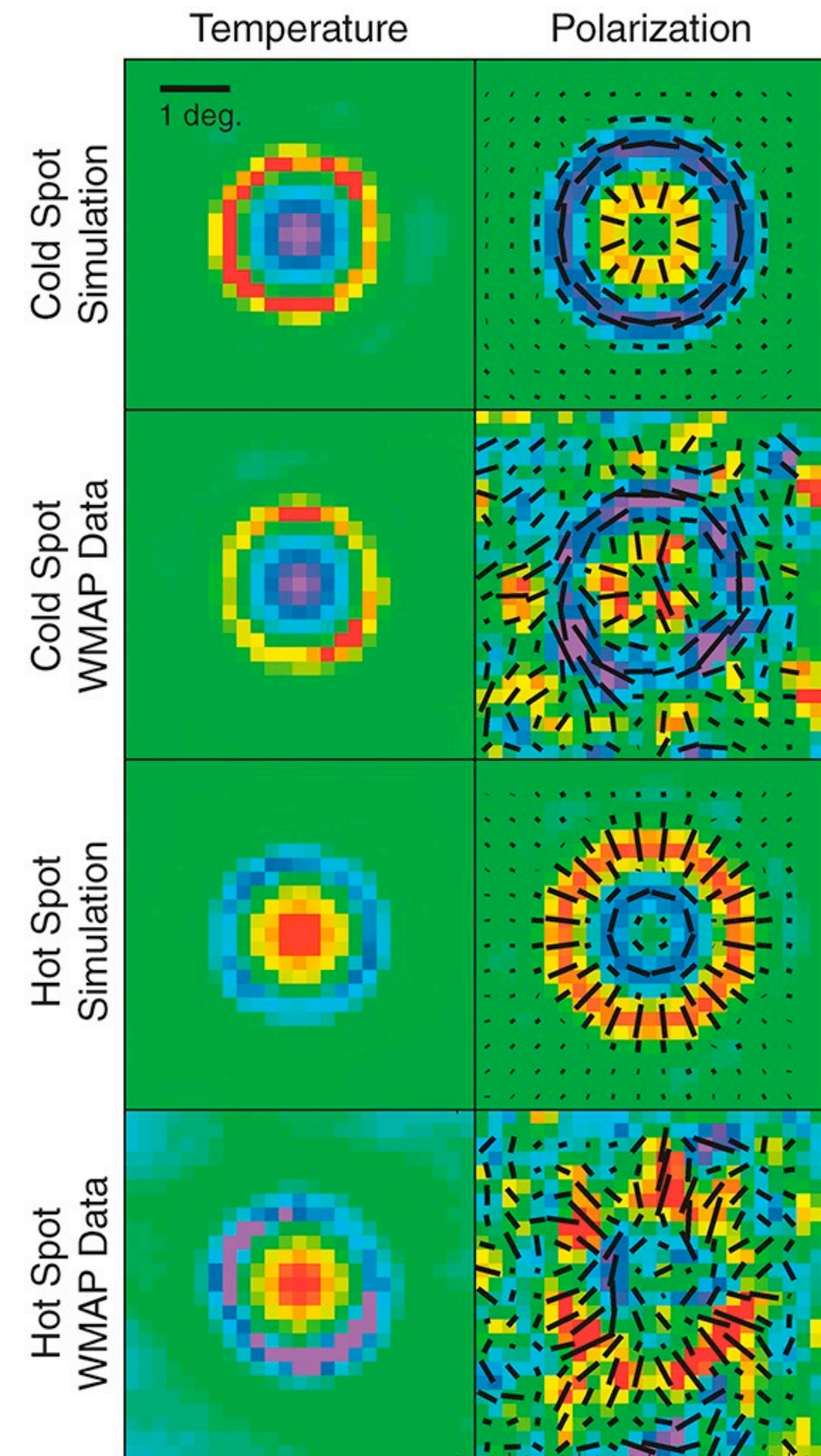
Stacking Analysis

- Stack polarization images around temperature hot and cold spots.
- Outside of the Galaxy mask (not shown), there are **12387 hot spots** and **12628 cold spots**.

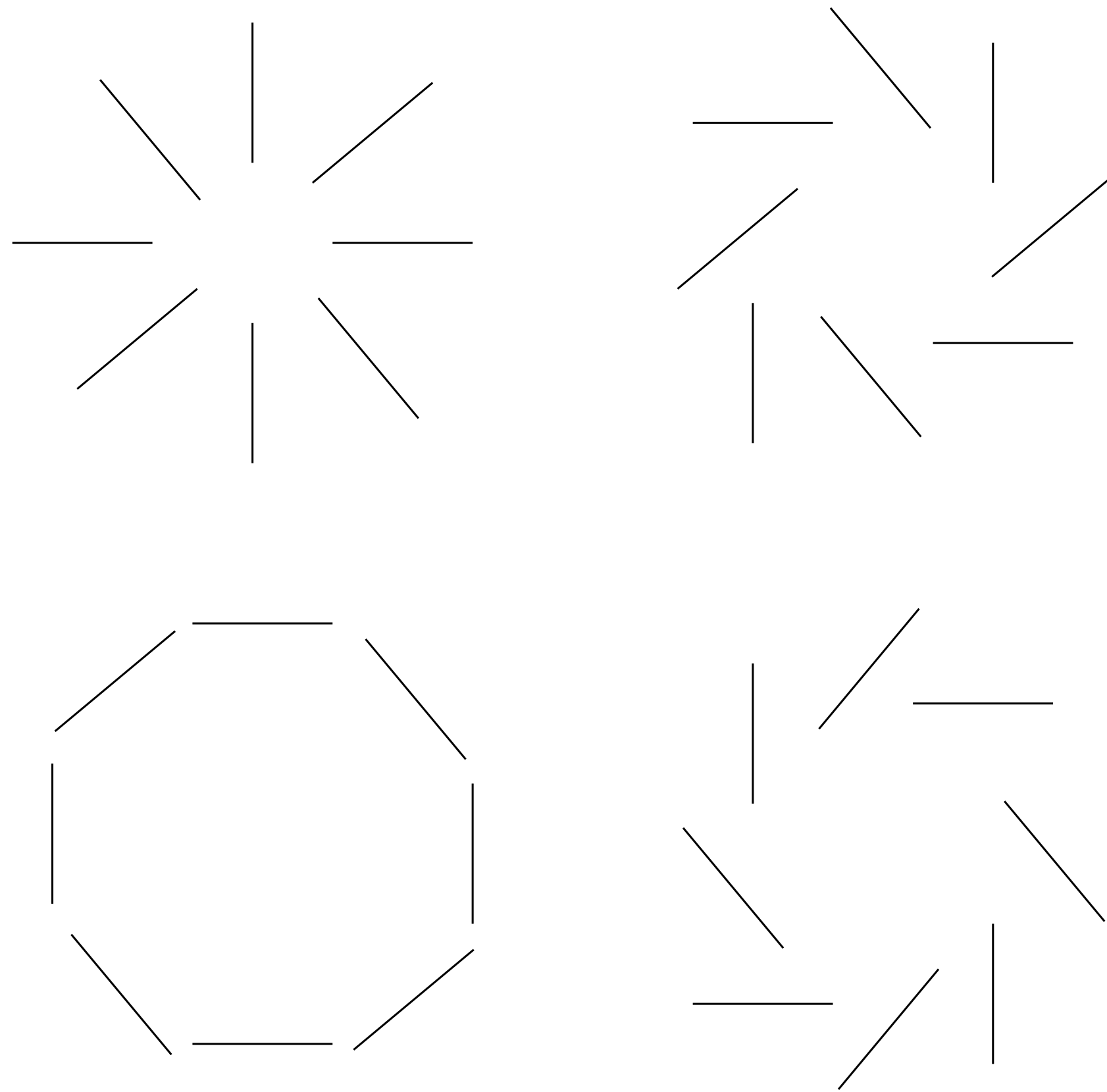


Radial and Tangential Polarization Patterns around Temp. Spots

- All hot and cold spots are stacked
- “Compression phase” at $\theta=1.2$ deg and “slow-down phase” at $\theta=0.6$ deg are predicted to be there and we observe them!
- The overall significance level: 8σ



E-mode and B-mode

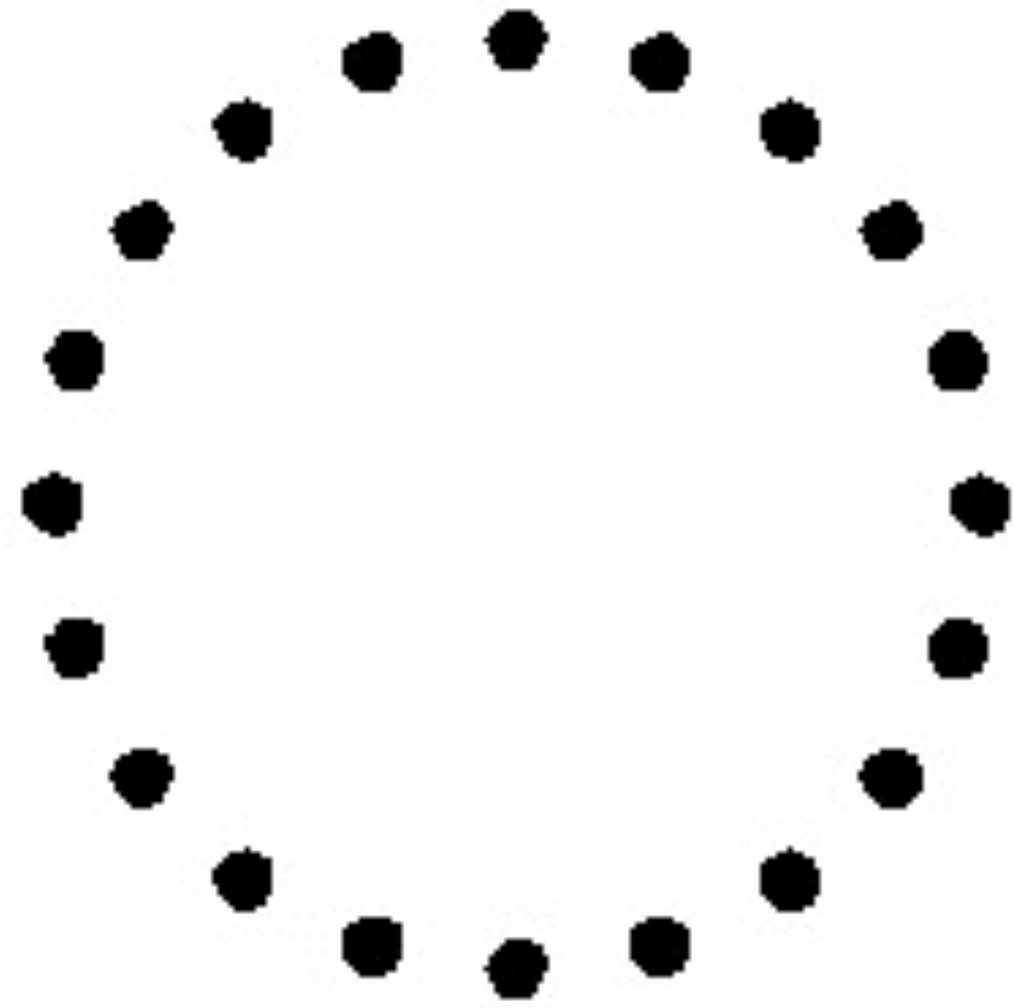


E mode

B mode

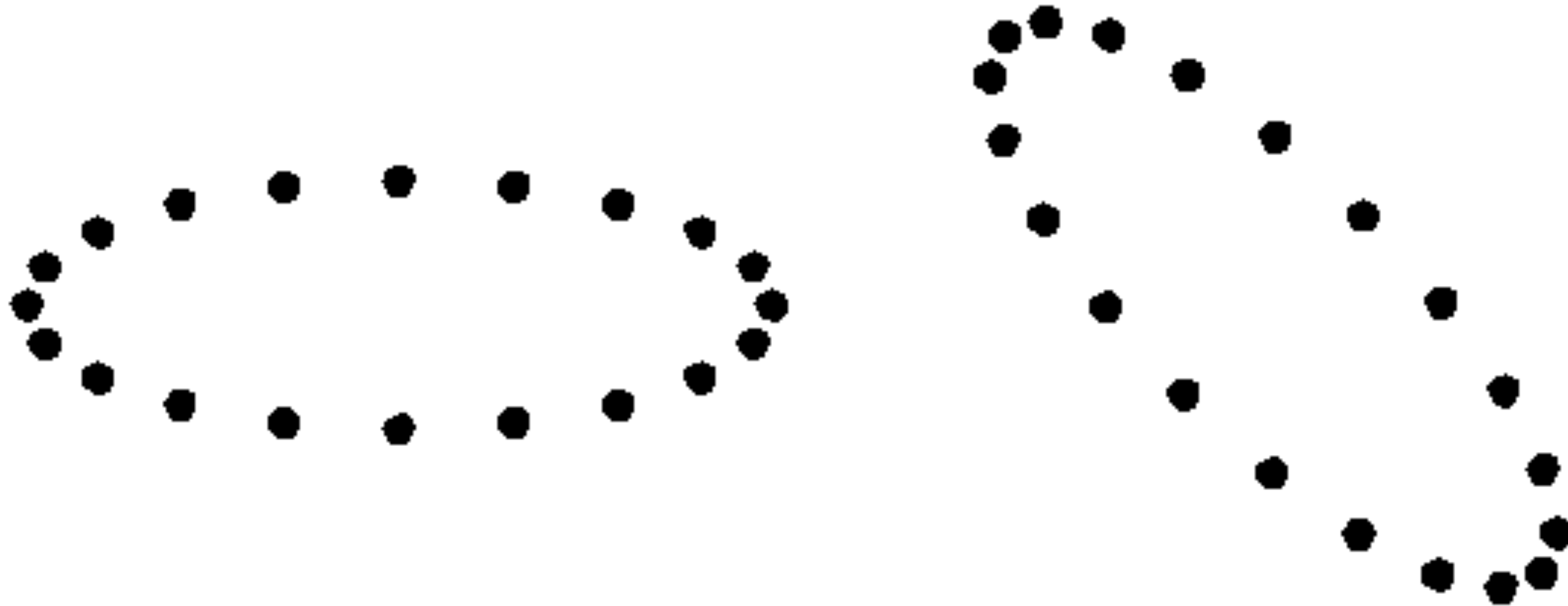
- Gravitational potential can generate the E-mode polarization, but not B-modes.
- **Gravitational waves** can generate both E- and B-modes!

Gravitational waves are coming toward you... What do you do?



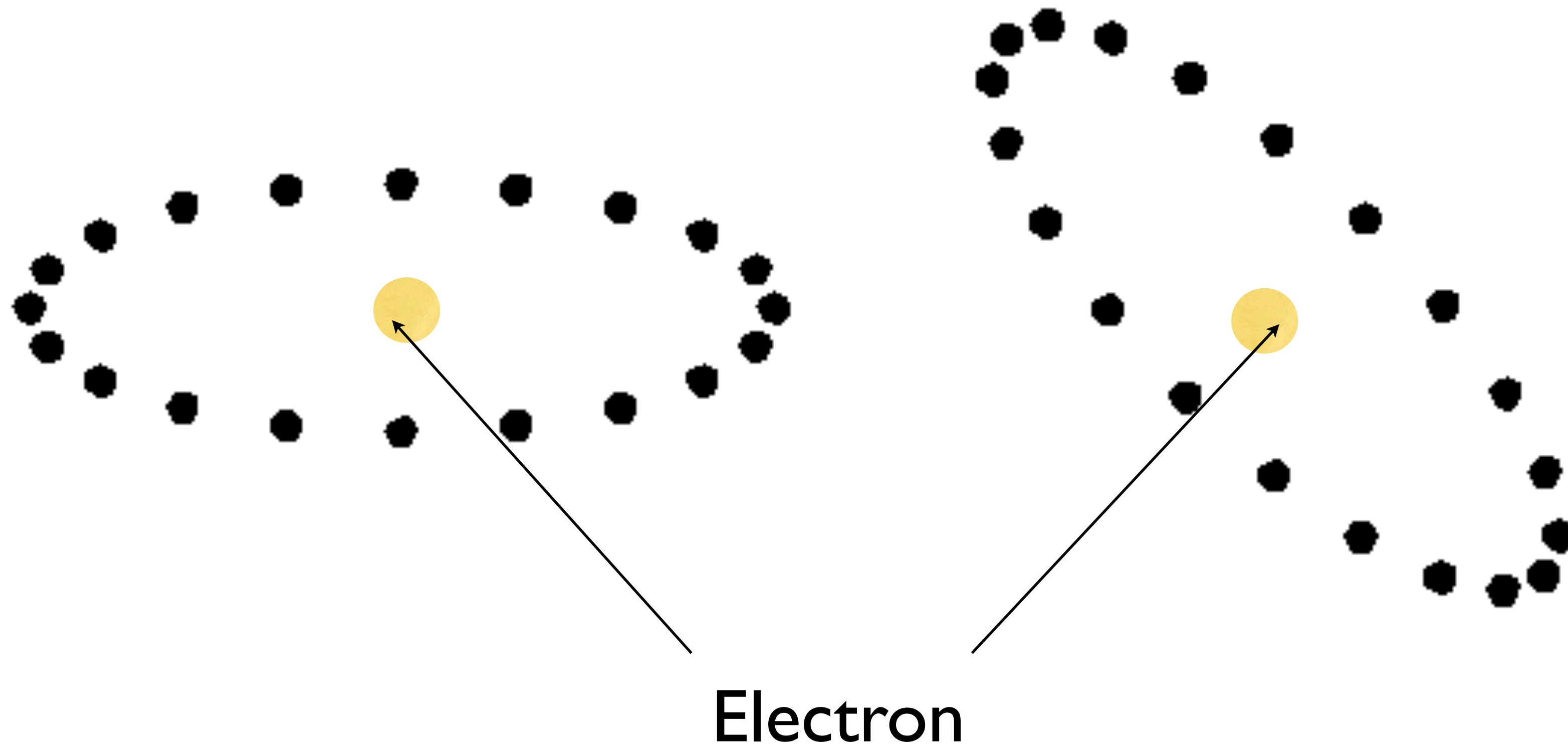
- Gravitational waves stretch space, causing particles to move.

Two Polarization States of GW

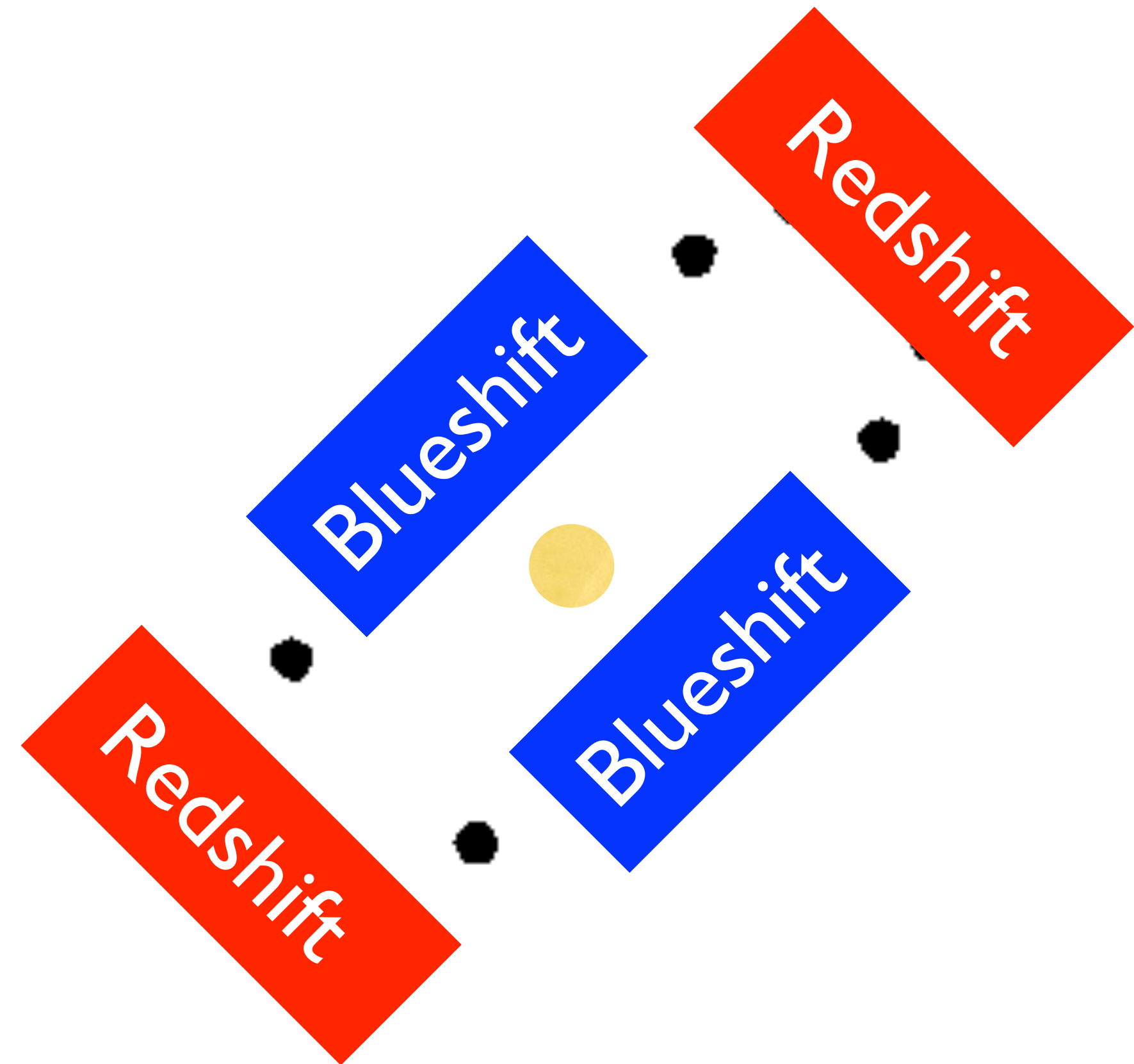
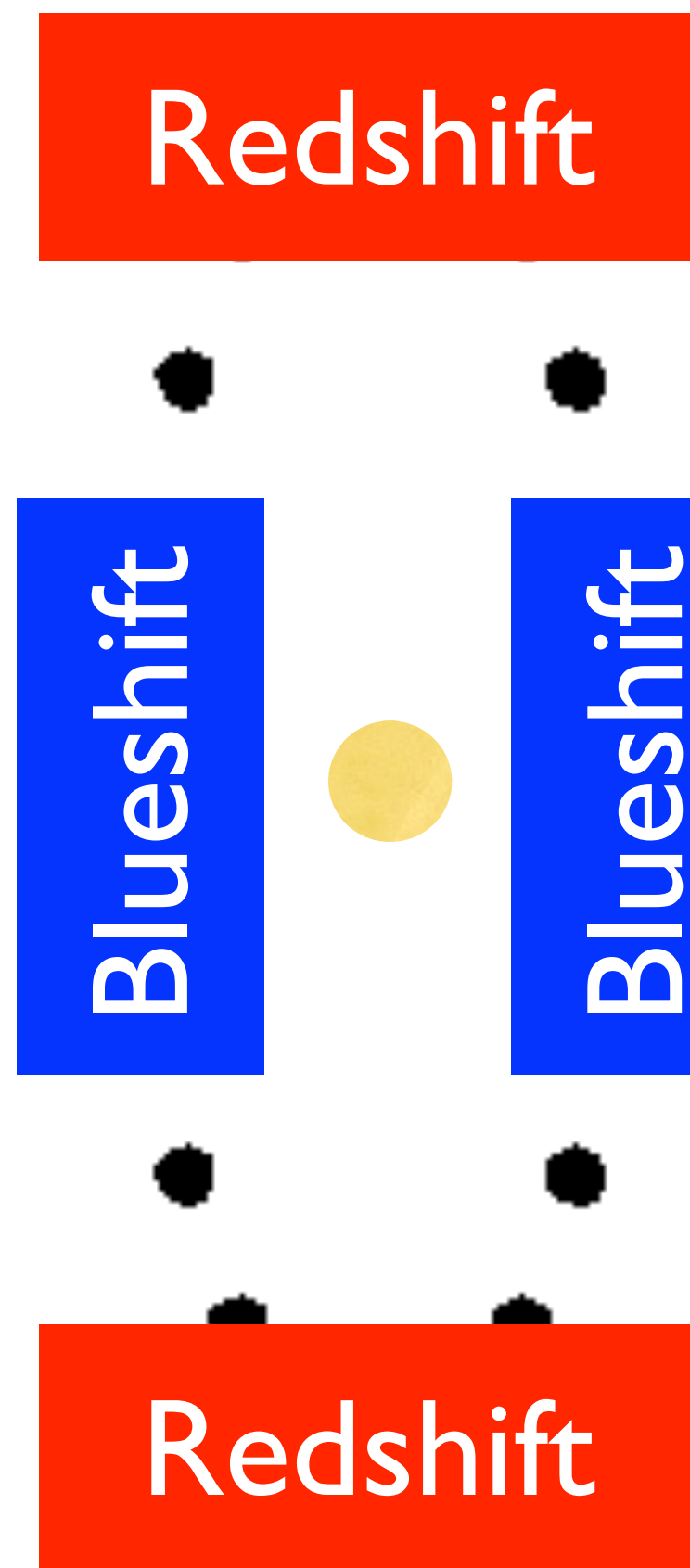


- This is great - this will automatically generate quadrupolar anisotropy around electrons!

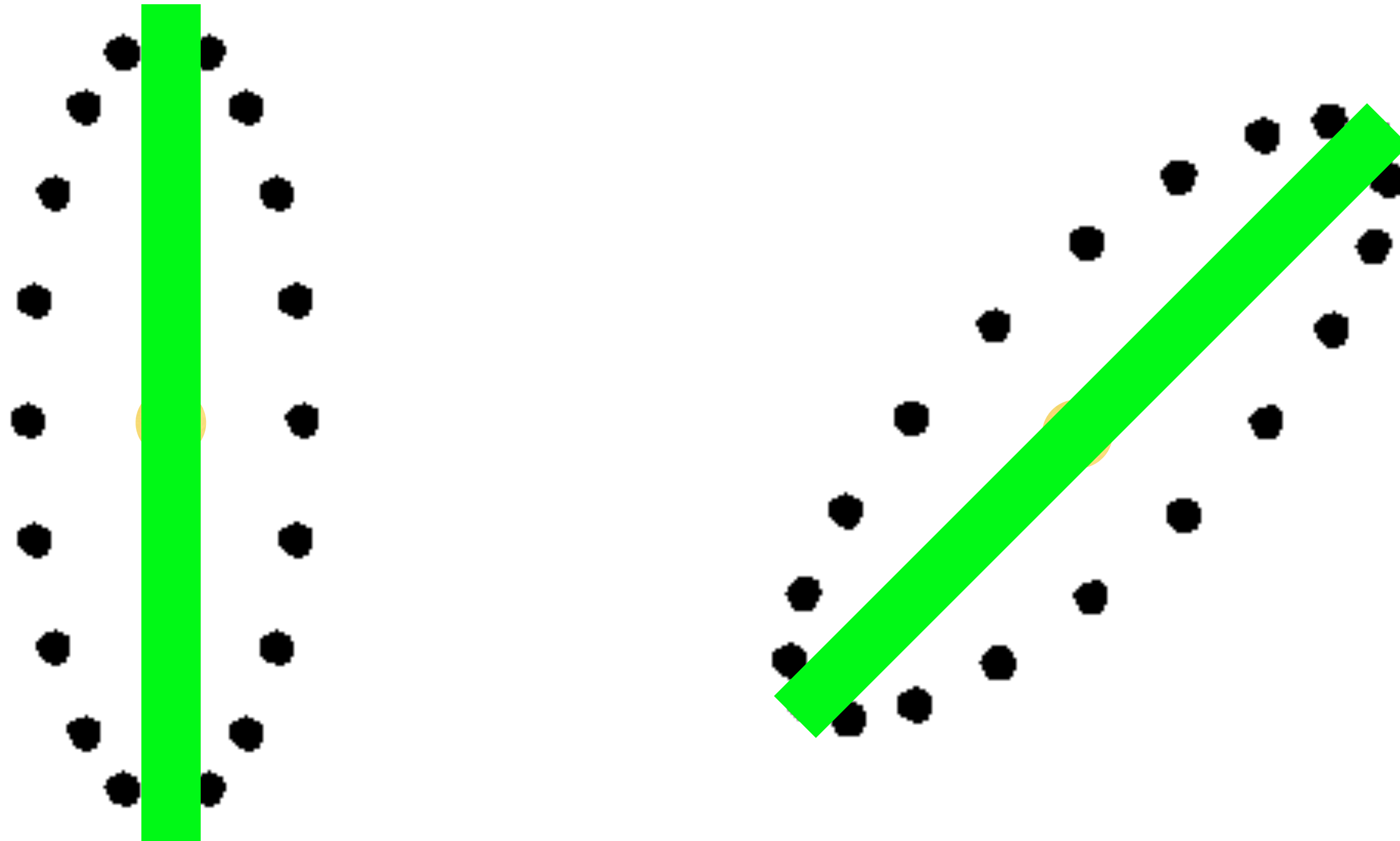
From GW to CMB Polarization



From GW to CMB Polarization



From GW to CMB Polarization



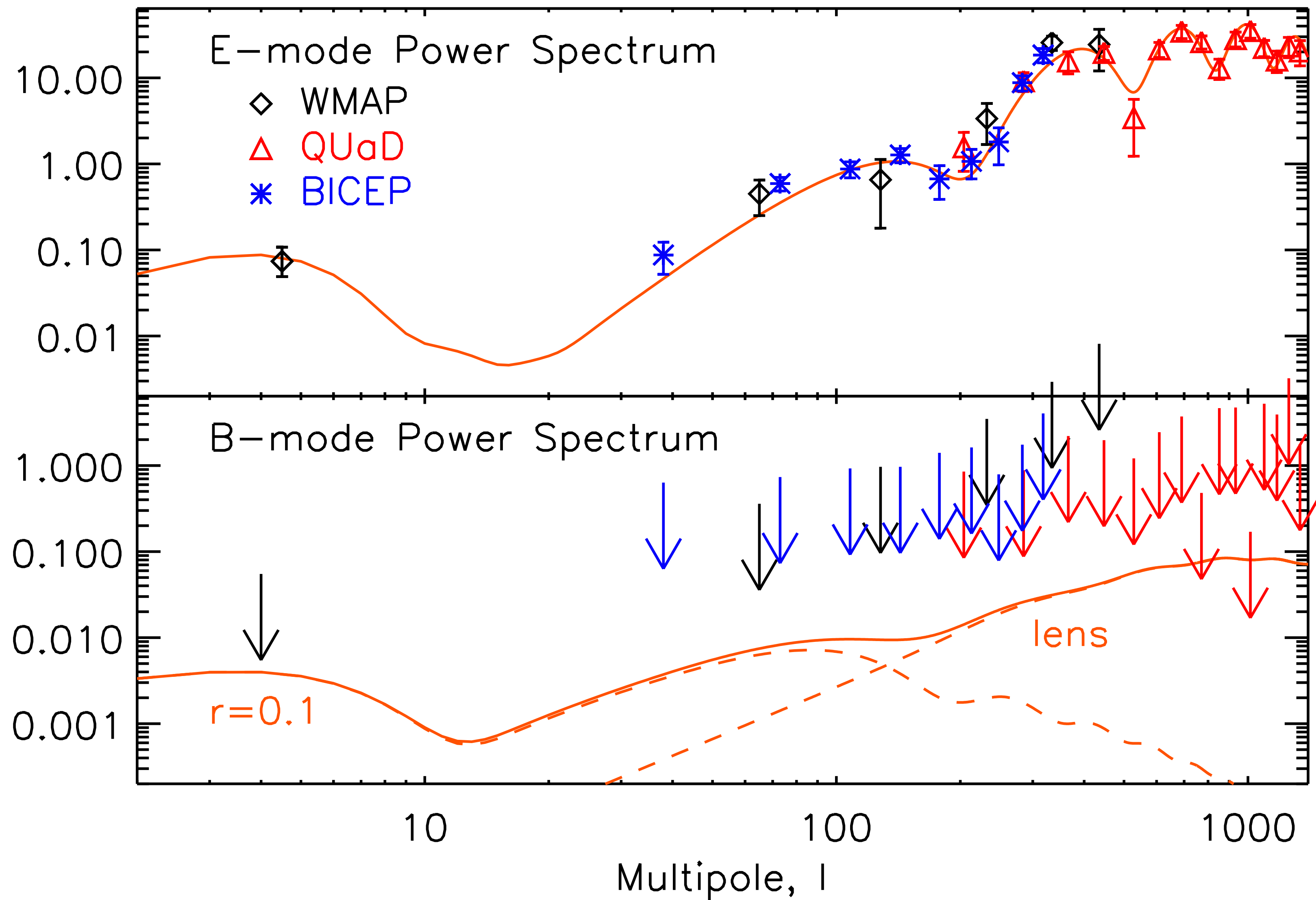
Gravitational waves can produce
both E- and B-mode polarization

“Tensor-to-scalar Ratio,” r

$$r = \frac{\text{[Power in Gravitational Waves]}}{\text{[Power in Gravitational Potential]}}$$

Theory of “Cosmic Inflation” predicts $r \ll 1$
– I will come back to this in a moment

Polarization Power Spectrum

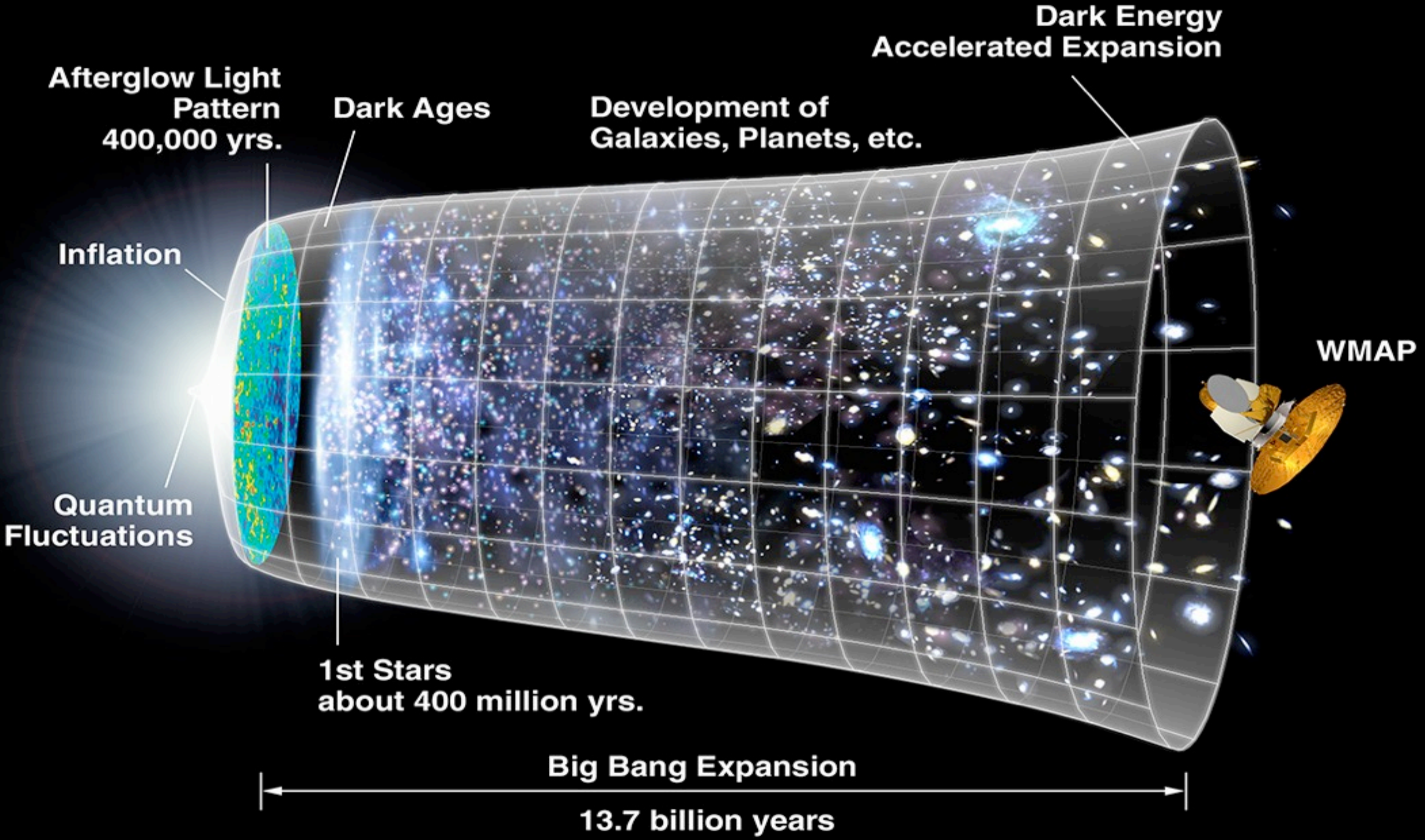


- No detection of B-mode polarization yet.
B-mode is the next holy grail!

Theory of the Very Early Universe

- The leading theoretical idea about the primordial Universe, called “**Cosmic Inflation**,” predicts:
(Guth 1981; Linde 1982; Albrecht & Steinhardt 1982; Starobinsky 1980)
- The expansion of our Universe **accelerated** in a tiny fraction of a second after its birth.
- Just like Dark Energy accelerating today’s expansion: the acceleration also happened at very, very early times!
- **Inflation stretches “micro to macro”**
- In a tiny fraction of a second, the size of an atomic nucleus ($\sim 10^{-15}\text{m}$) would be stretched to 1 A.U. ($\sim 10^{11}\text{m}$), at least.

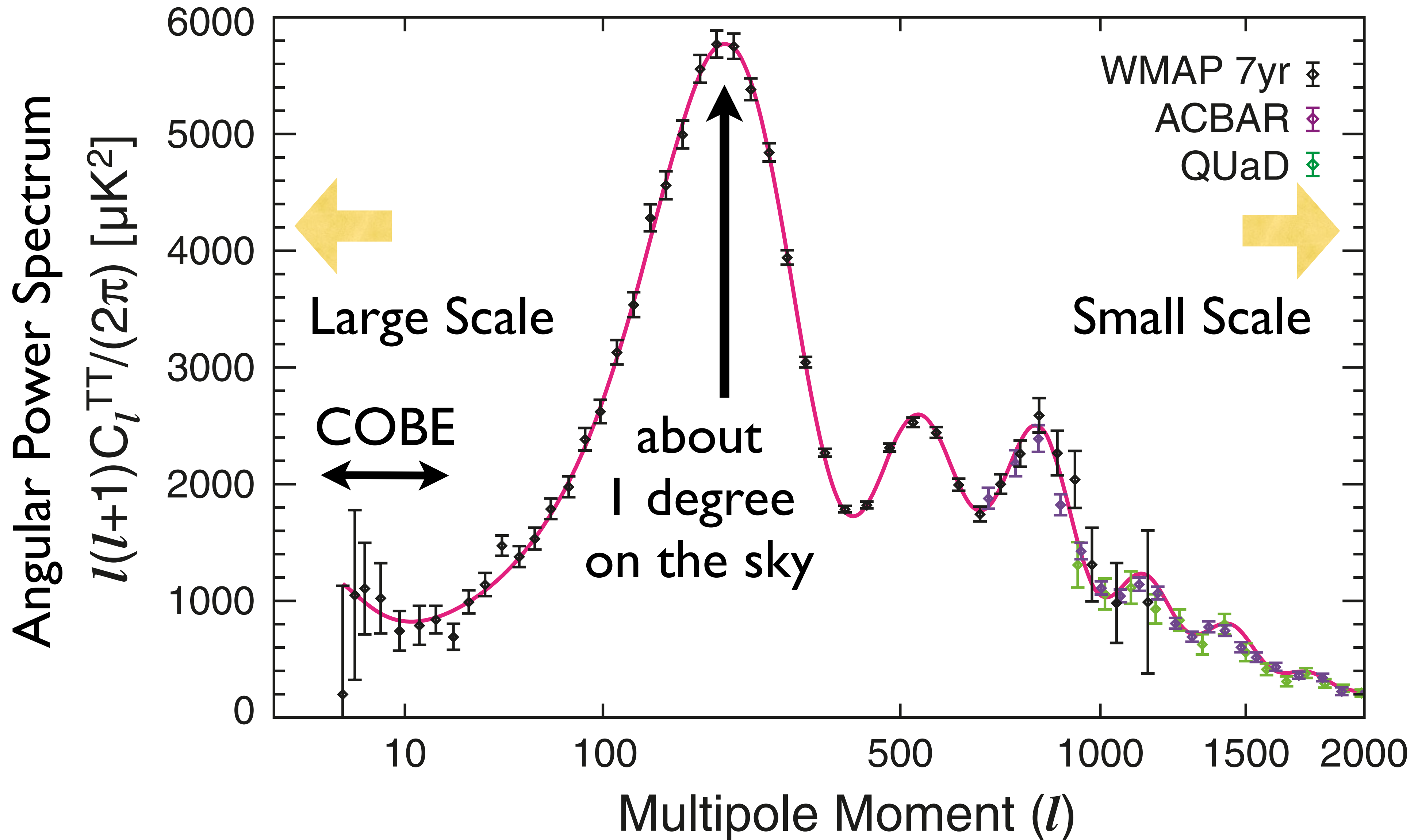
Cosmic Inflation = Very Early Dark Energy



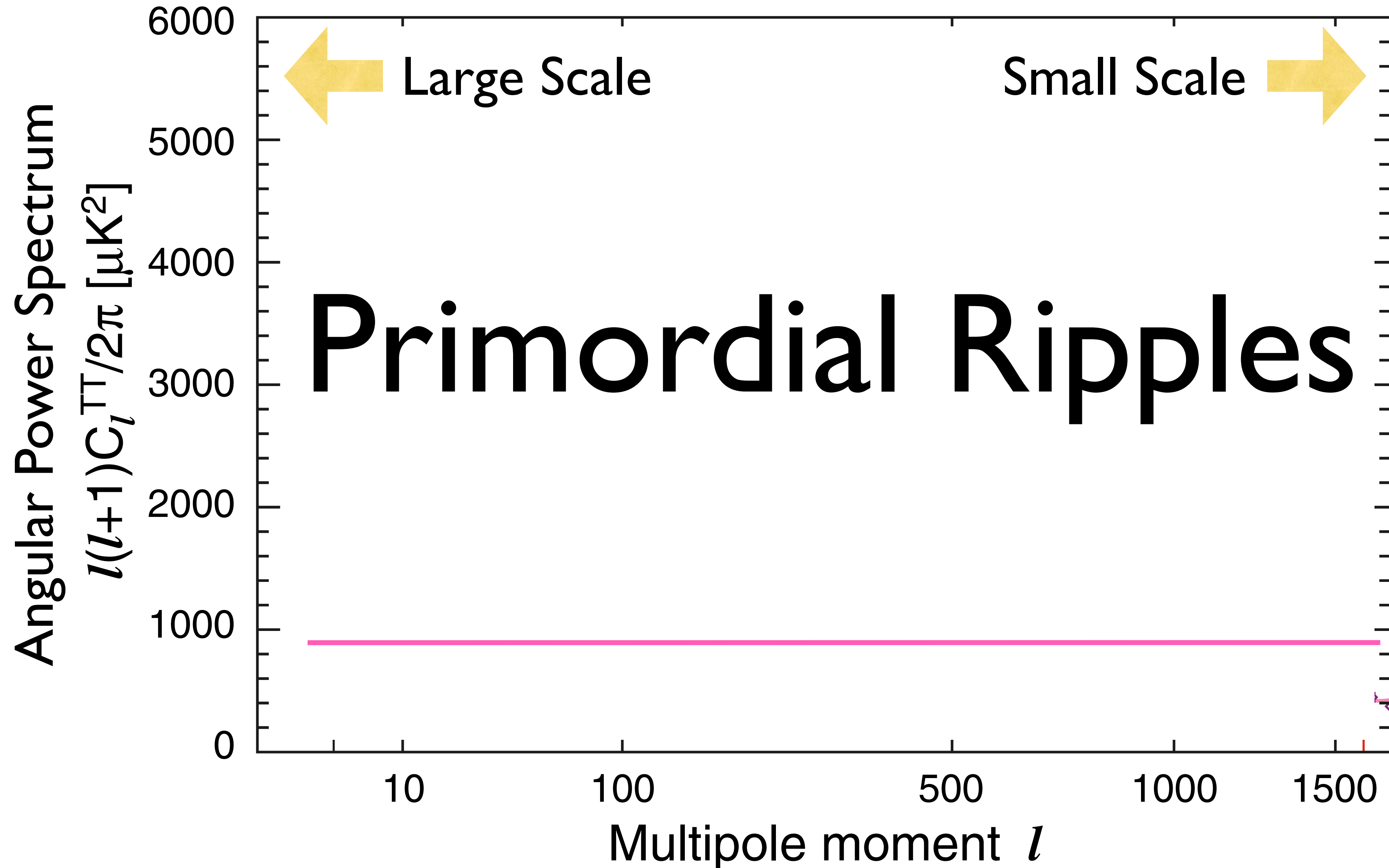
Origin of Fluctuations

- OK, back to the cosmic hot soup.
- The sound waves were created when we perturbed it.
- “We”? **Who?**
- Who actually perturbed the cosmic soup?
- **Who generated the original (seed) ripples?**

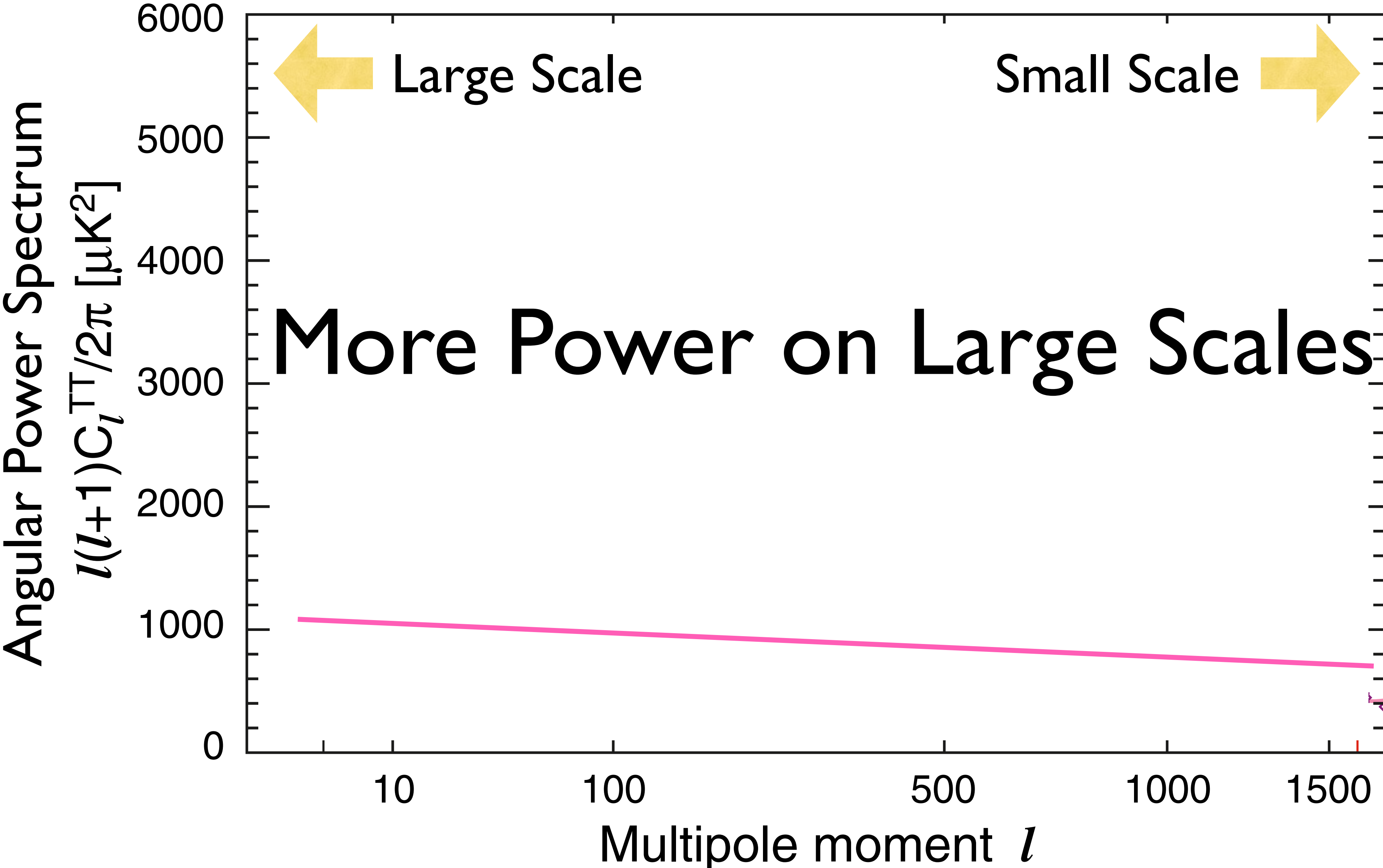
WMAP Power Spectrum



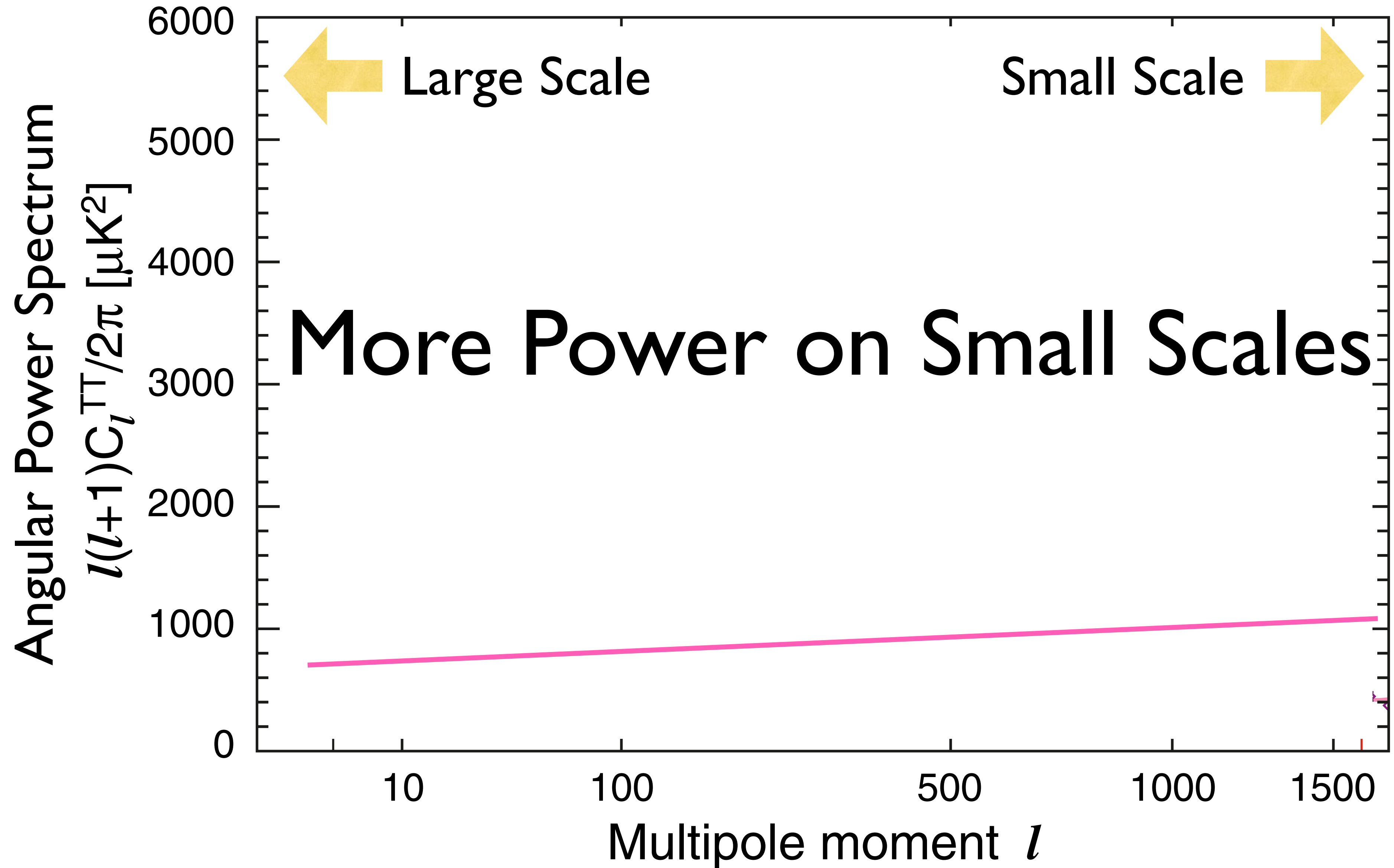
Getting rid of the Sound Waves



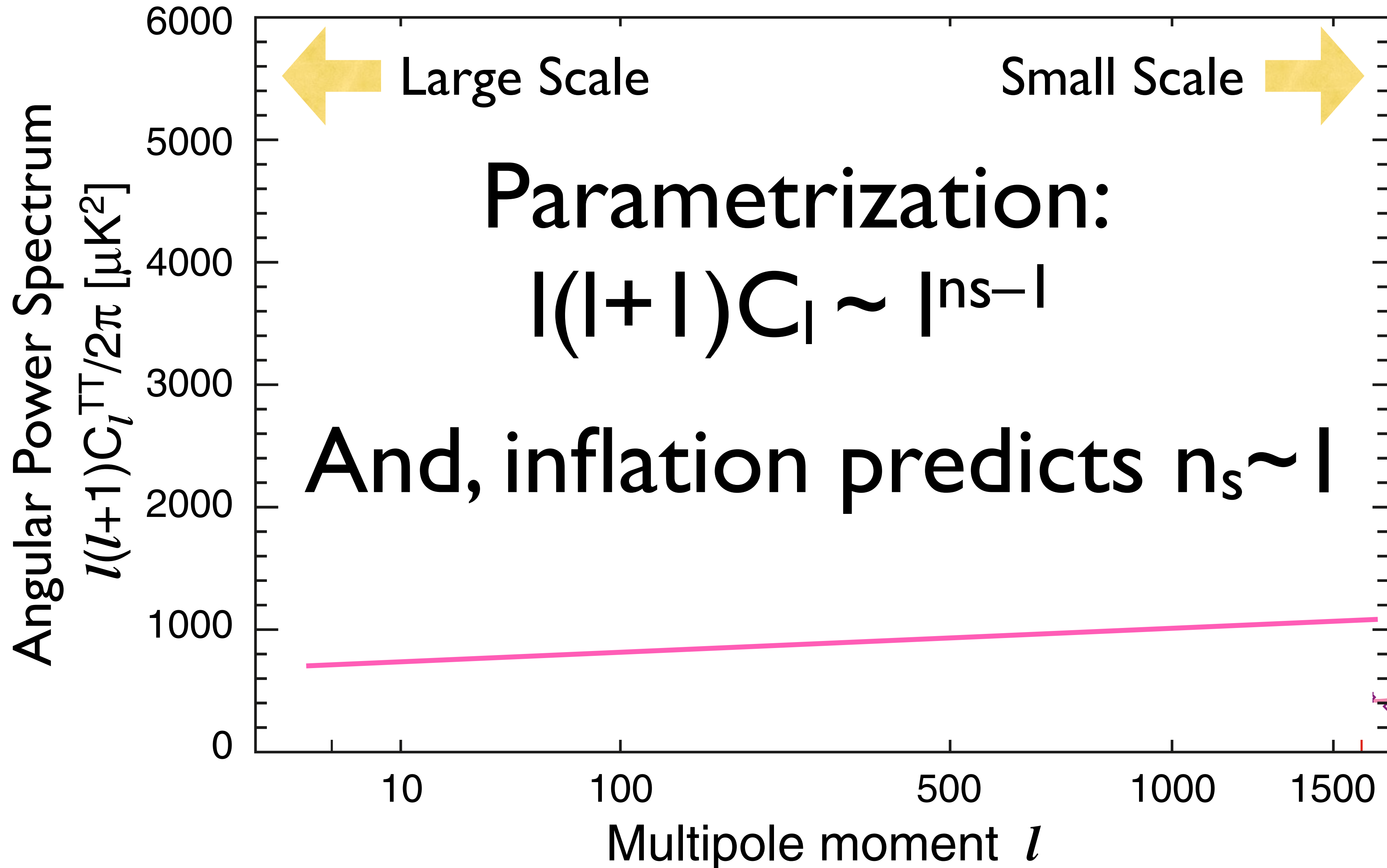
The Early Universe Could Have Done This Instead



...or, This.



...or, This.



Theory Says...

- The leading theoretical idea about the primordial Universe, called “**Cosmic Inflation**,” predicts:
 - The expansion of our Universe **accelerated** in a tiny fraction of a second after its birth.
 - the primordial ripples were created by **quantum fluctuations** during inflation, and
 - how the power is distributed over the scales is determined by the expansion history during cosmic inflation.
- Measurement of n_s gives us **this** remarkable information!

*Mukhanov & Chibisov (1981); Guth & Pi (1982); Starobinsky (1982); Hawking (1982);
Bardeen, Turner & Steinhardt (1983)*

(Scalar) Quantum Fluctuations

$$\delta\varphi = (\text{Expansion Rate})/(2\pi) \text{ [in natural units]}$$

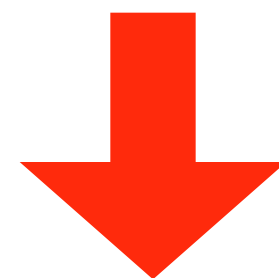
- Why is this relevant?
- The cosmic inflation (probably) happened when the Universe was a tiny fraction of second old.
 - Something like 10^{-36} second old
 - (Expansion Rate) $\sim 1/(\text{Time})$
 - which is a big number! ($\sim 10^{12}\text{GeV}$)
- *Quantum fluctuations were important during inflation!*

Stretching Micro to Macro

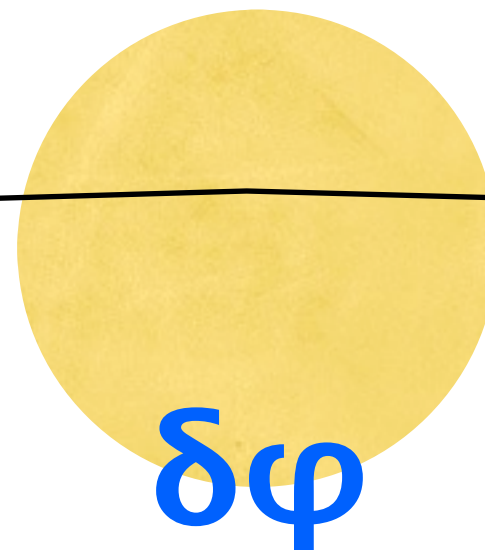
Macroscopic size at which gravity becomes important



Quantum fluctuations on microscopic scales



INFLATION!



Quantum fluctuations cease to be quantum, and become observable!

Inflation Offers a Magnifier for Microscopic World

- Using the *power spectrum of primordial fluctuations* imprinted in CMB, we can observe the quantum phenomena at the ultra high-energy scales that would never be reached by the particle accelerator.
- Measured value: $n_s = 0.968 \pm 0.012$ (68%CL)

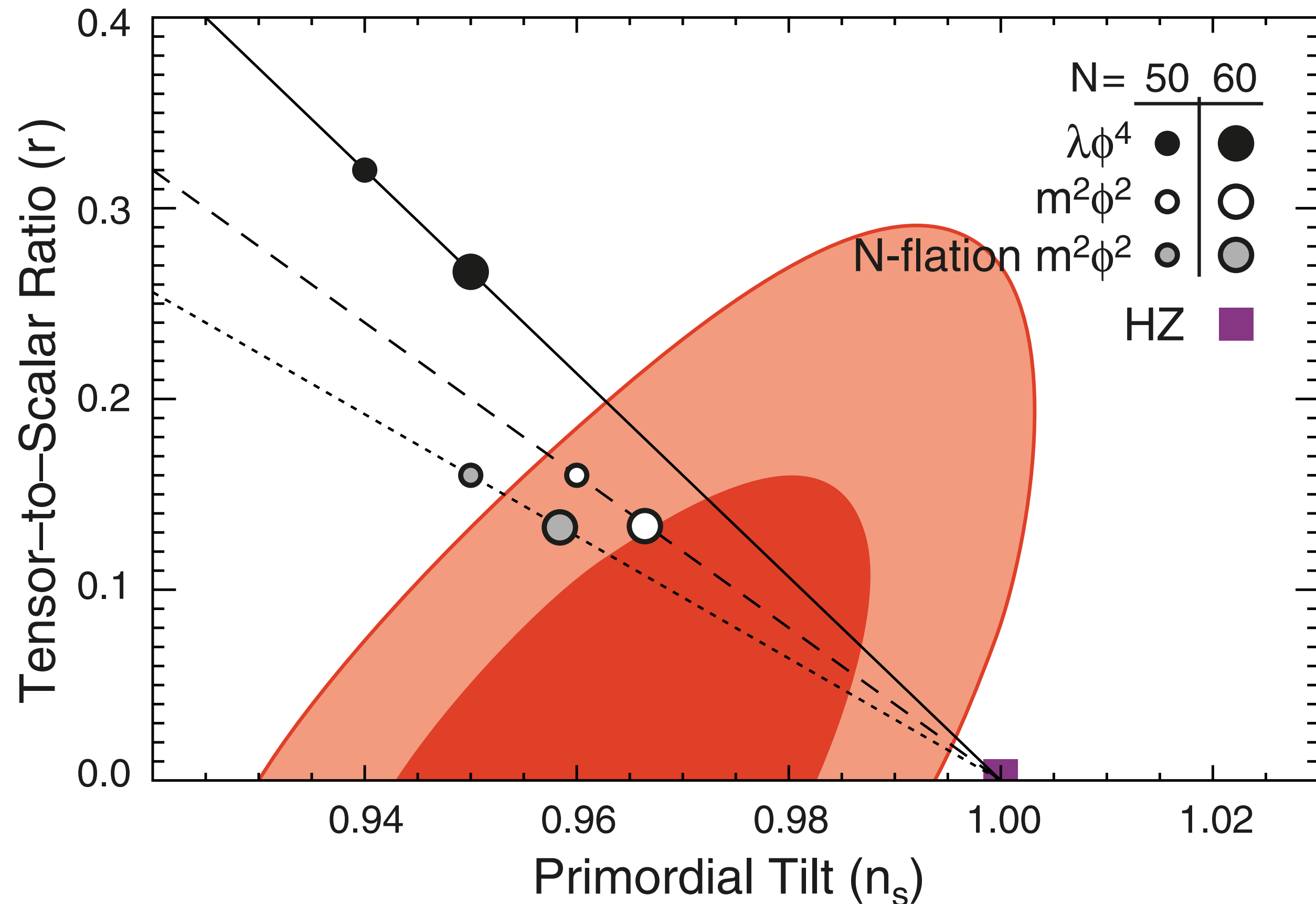
(Tensor) Quantum Fluctuations, a.k.a. Gravitational Waves

$$h = (\text{Expansion Rate}) / (2^{1/2} \pi M_{\text{planck}}) \text{ [in natural units]}$$

[h = “strain”]

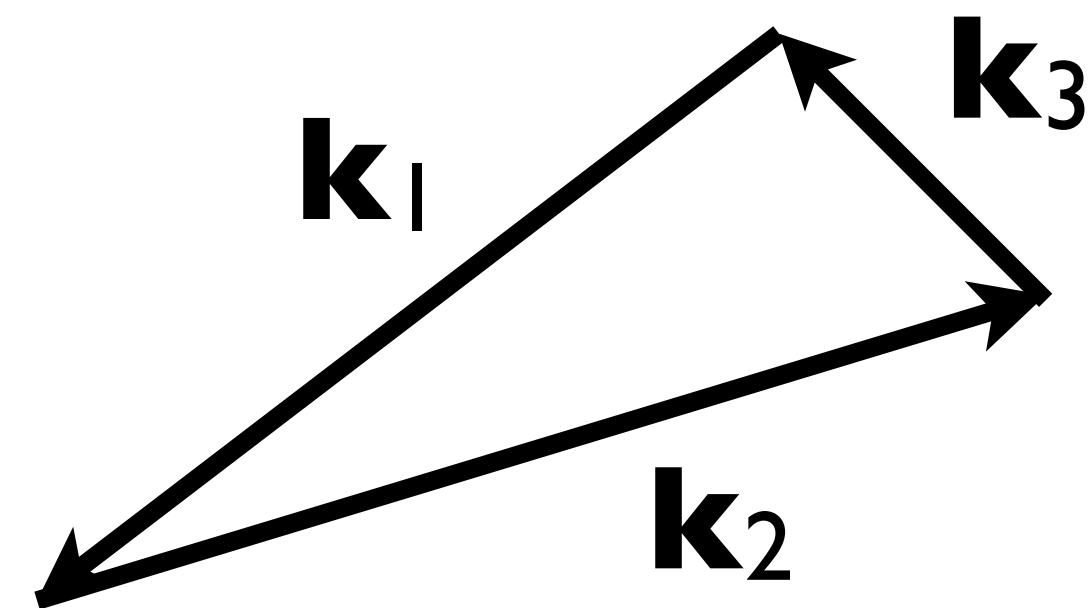
- Quantum fluctuations also generate ripples in space-time, i.e., gravitational waves, by the same mechanism.
- Primordial gravitational waves generate temperature anisotropy in CMB, as well as polarization in CMB with a distinct pattern called “**B-mode polarization.**”

Probing Inflation (2-point Function)



- Joint constraint on the primordial tilt, n_s , and the tensor-to-scalar ratio, r .
- $r < 0.24$ (95%CL)

Bispectrum

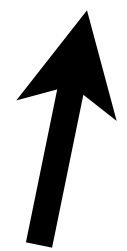


- Three-point function!

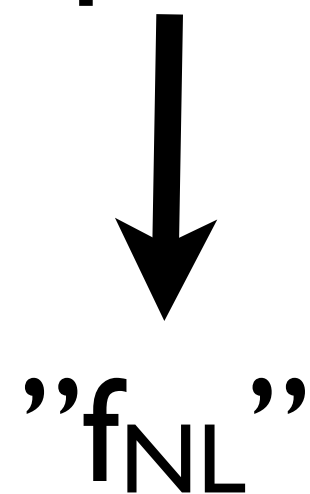
- $B_\zeta(\mathbf{k}_1, \mathbf{k}_2, \mathbf{k}_3)$

$$= \langle \zeta_{\mathbf{k}_1} \zeta_{\mathbf{k}_2} \zeta_{\mathbf{k}_3} \rangle = (\text{amplitude}) \times (2\pi)^3 \delta(\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3) F(k_1, k_2, k_3)$$

model-dependent function

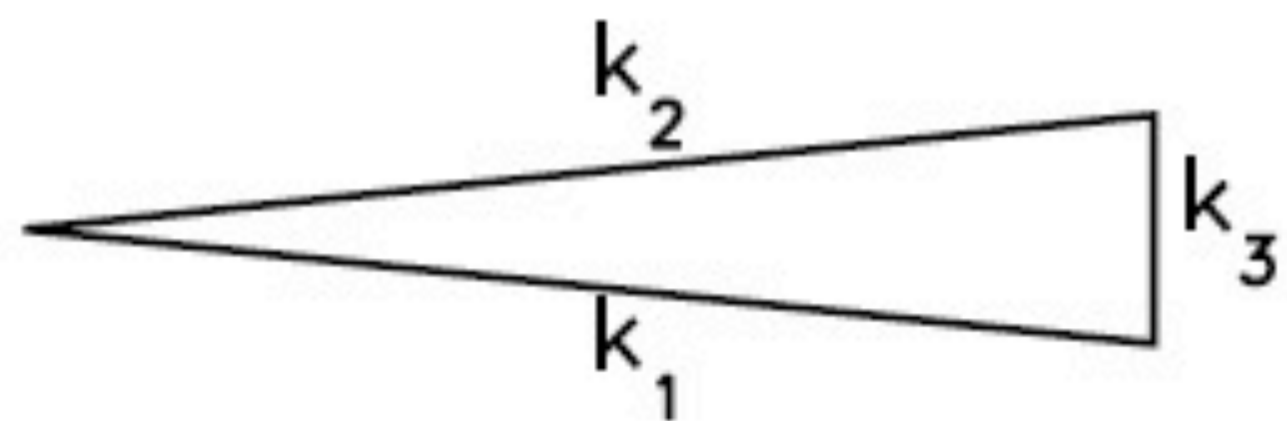


Primordial fluctuation

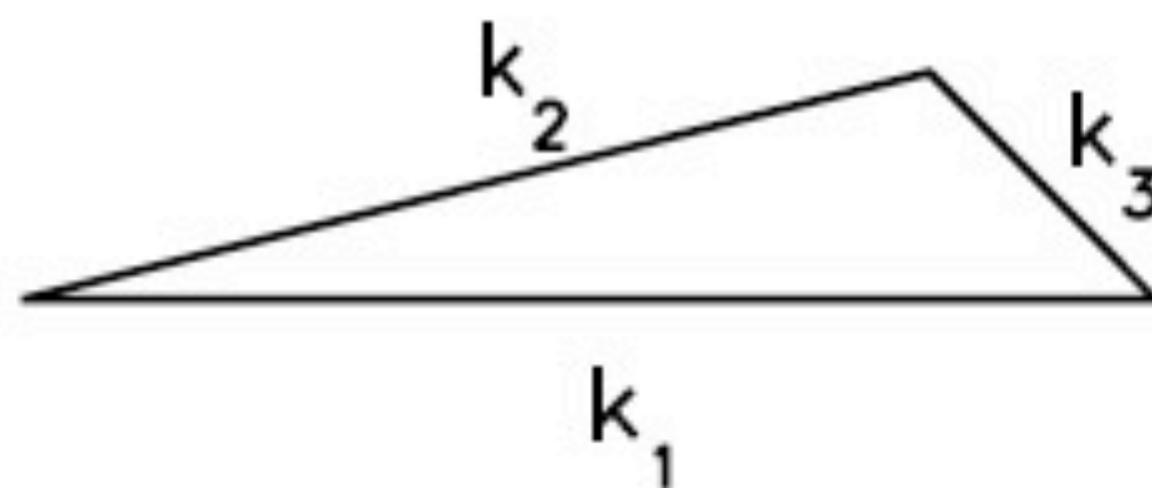


"f_{NL}"

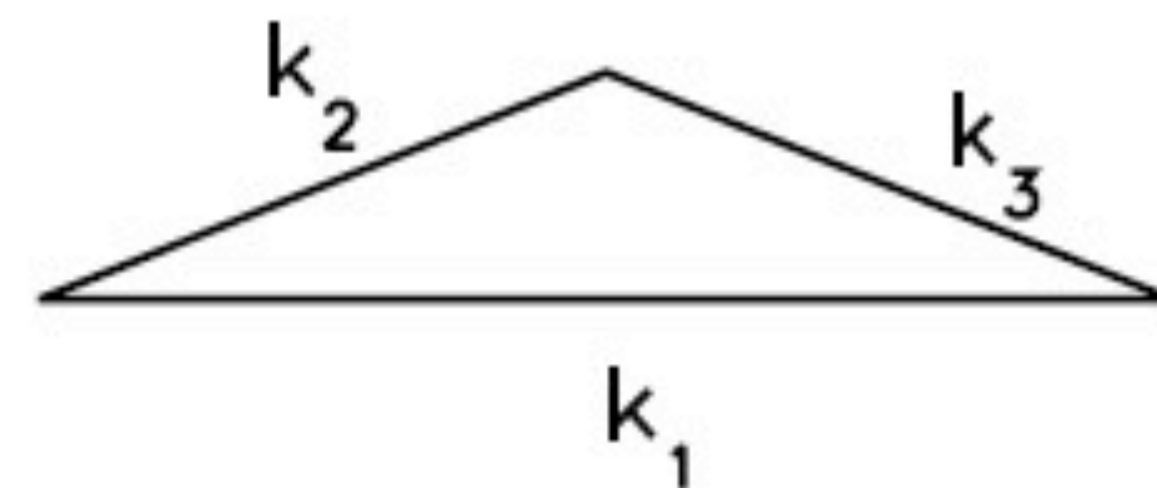
(a) squeezed triangle
($k_1 \approx k_2 \gg k_3$)



(b) elongated triangle
($k_1 = k_2 + k_3$)

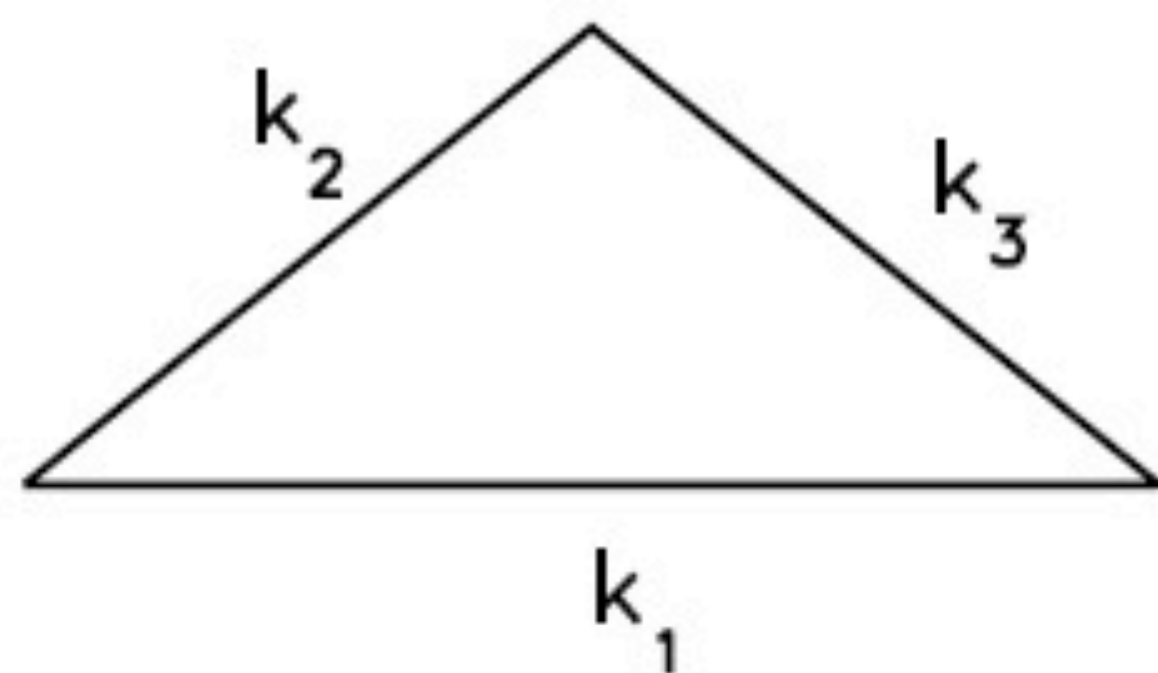


(c) folded triangle
($k_1 = 2k_2 = 2k_3$)

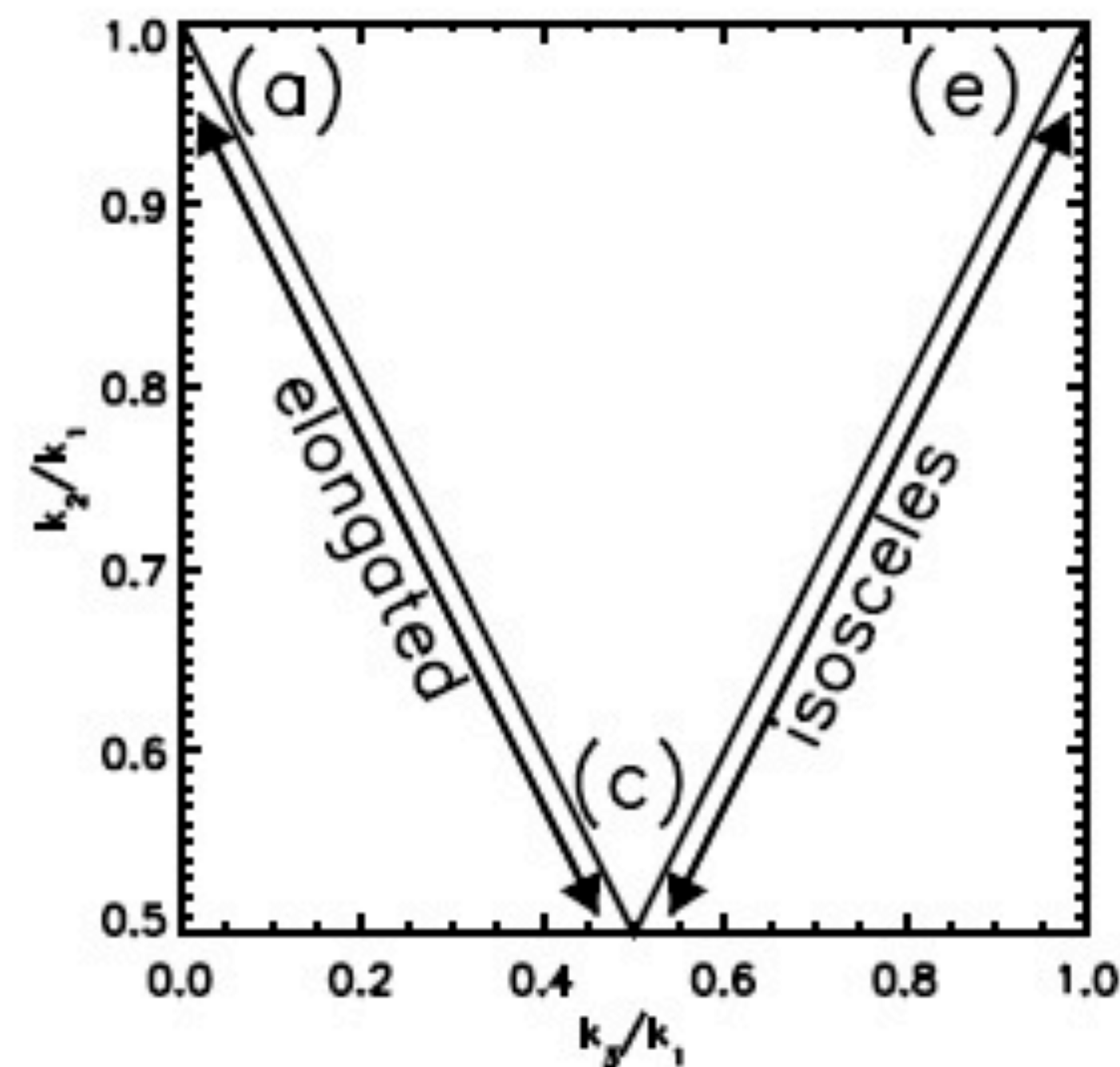
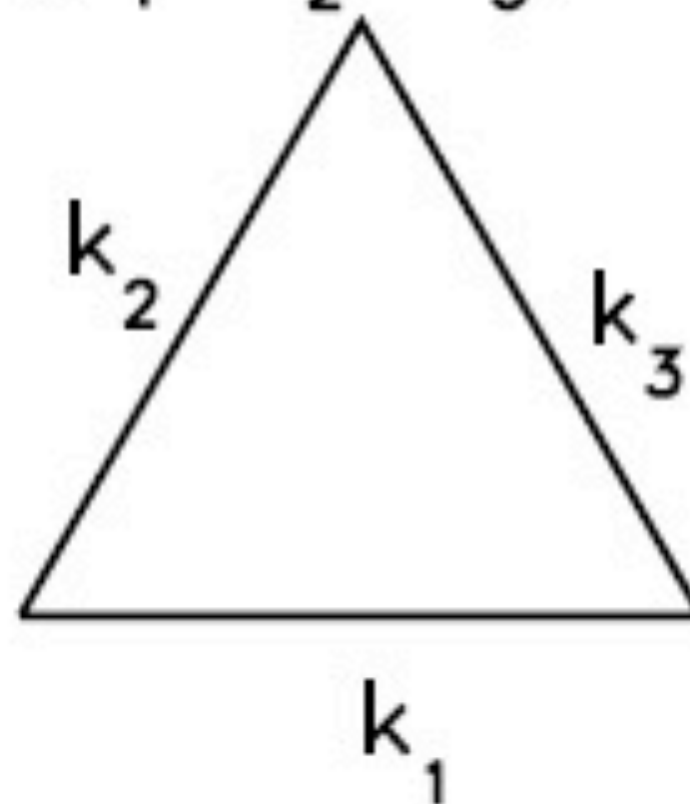


MOST IMPORTANT

(d) isosceles triangle
($k_1 > k_2 = k_3$)



(e) equilateral triangle
($k_1 = k_2 = k_3$)



Probing Inflation (3-point Function)

- Inflation models predict that primordial fluctuations are very close to Gaussian.
- In fact, **ALL SINGLE-FIELD** models predict a particular form of 3-point function to have the amplitude of $f_{\text{NL}}=0.02$.
- Detection of $f_{\text{NL}} > 1$ would rule out ALL single-field models!
- No detection of 3-point functions of primordial curvature perturbations. The 68% CL limit is:
 - **$f_{\text{NL}} = 32 \pm 21$ (1σ)**
 - The WMAP data are consistent with the prediction of **simple single-field inflation** models: $1-n_s \approx r \approx f_{\text{NL}}$

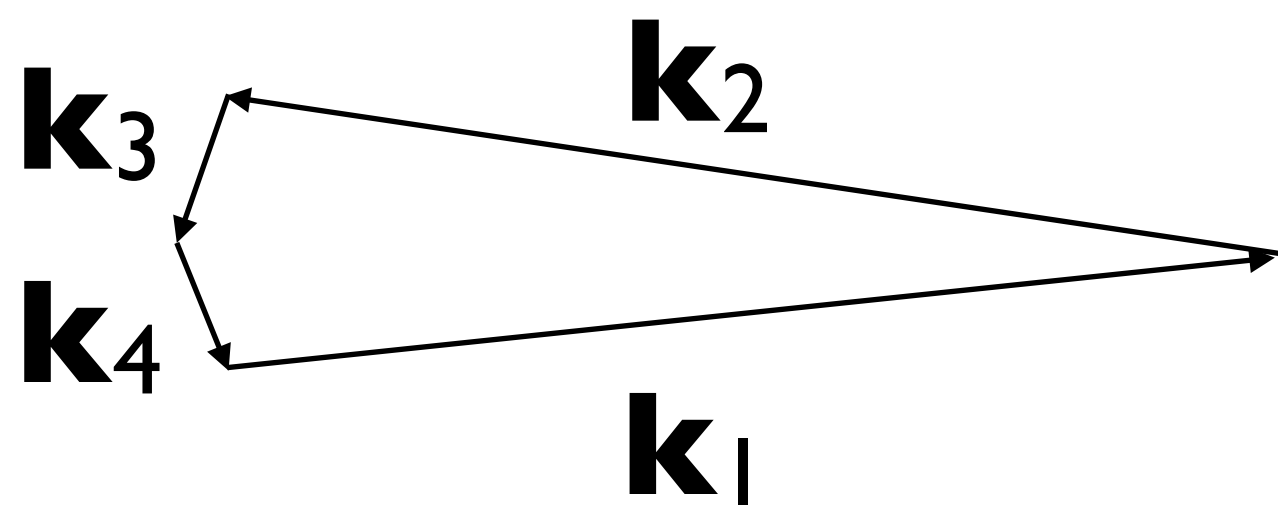
Summary

- CMB is the fossil light of the Big Bang.
- We could determine the age, composition, expansion rate, etc., from CMB.
- We could even push the boundary farther back in time, probing the origin of fluctuations in the very early Universe: inflationary epoch at ultra-high energies.
- Next Big Thing: **Primordial gravitational waves.**
- The 3-point function: **Powerful test of inflation.**

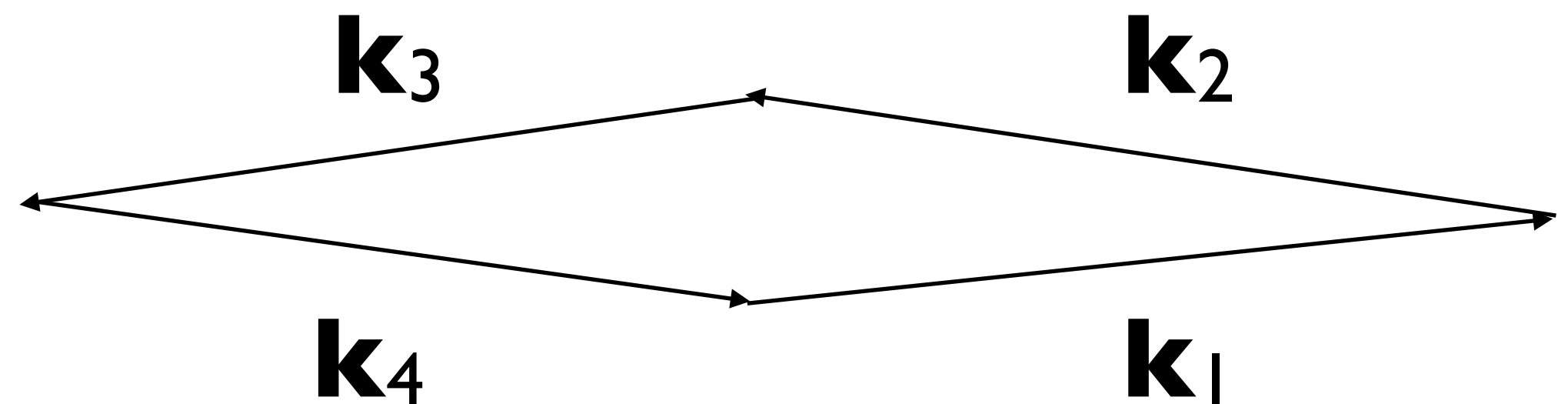
Trispectrum

- $T_{\zeta}(\mathbf{k}_1, \mathbf{k}_2, \mathbf{k}_3, \mathbf{k}_4) = (2\pi)^3 \delta(\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3 + \mathbf{k}_4)$
 $\{g_{NL}[(54/25)P_{\zeta}(k_1)P_{\zeta}(k_2)P_{\zeta}(k_3) + \text{cyc.}]$
 $+ T_{NL}[P_{\zeta}(k_1)P_{\zeta}(k_2)(P_{\zeta}(|\mathbf{k}_1 + \mathbf{k}_3|) + P_{\zeta}(|\mathbf{k}_1 + \mathbf{k}_4|)) + \text{cyc.}]\}$

*The local form consistency relation,
 $T_{NL} = (6/5)(f_{NL})^2$, may not be respected –
 additional test of multi-field inflation!*

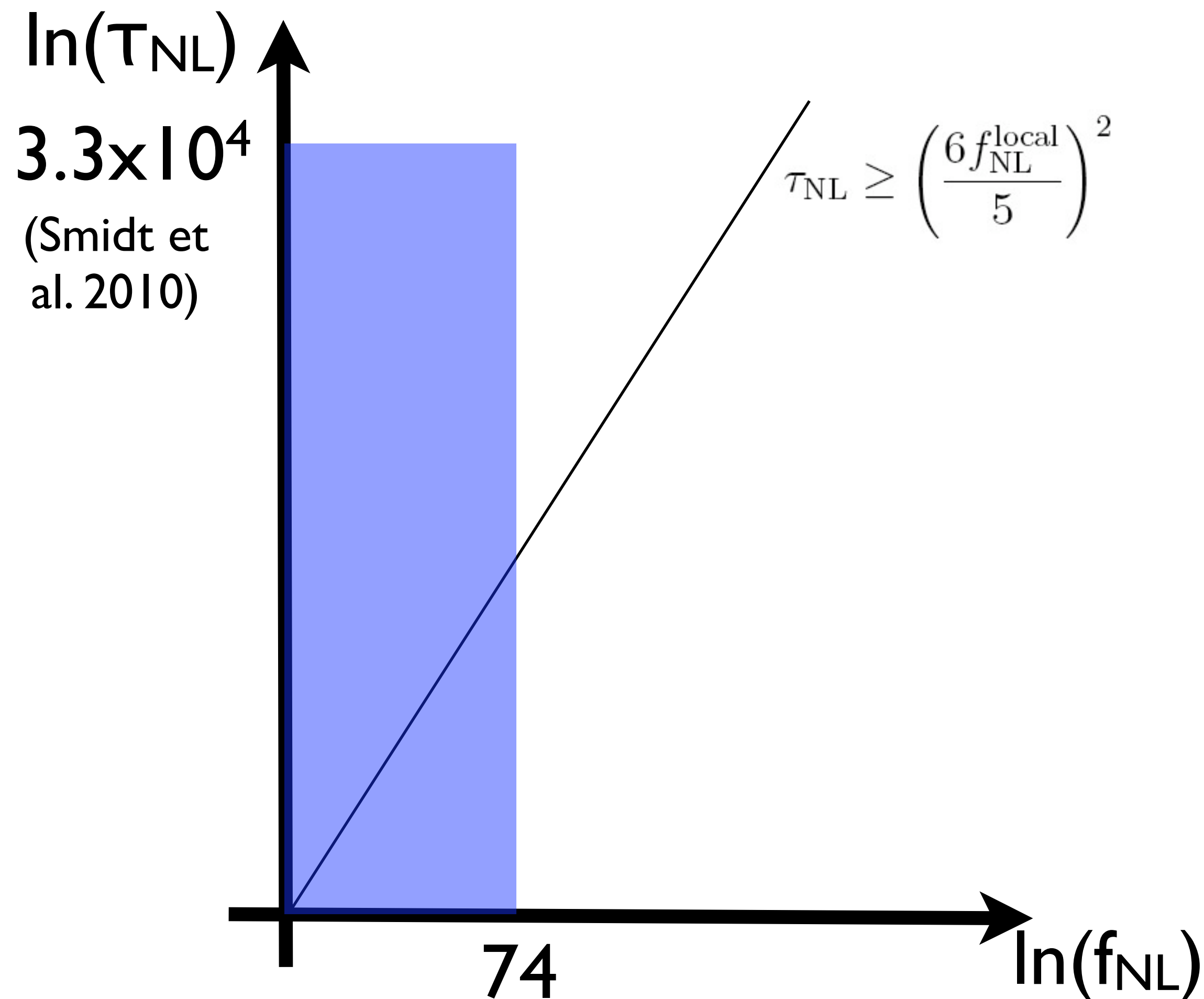


g_{NL}



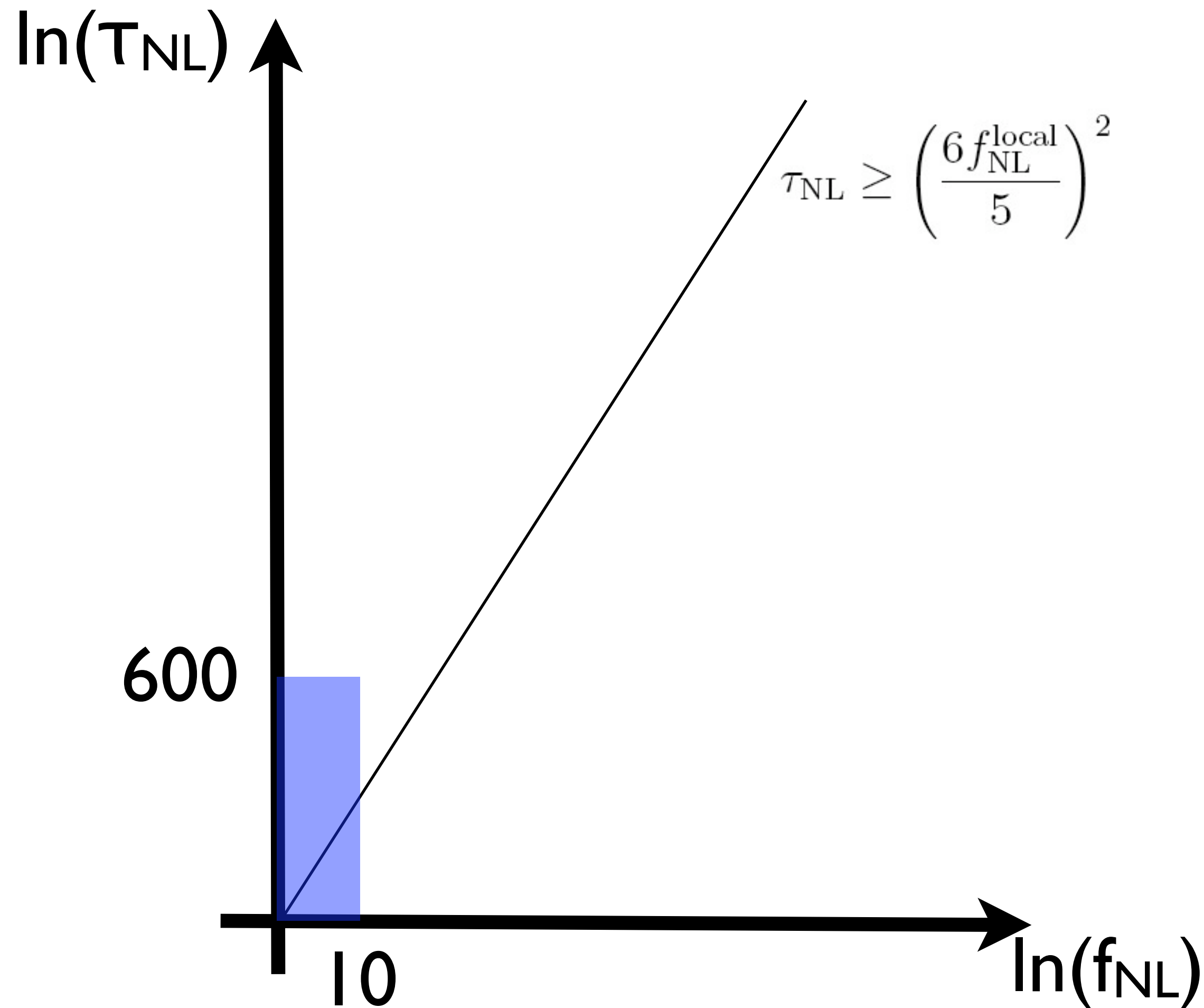
T_{NL}

The diagram that you should take away from this talk.



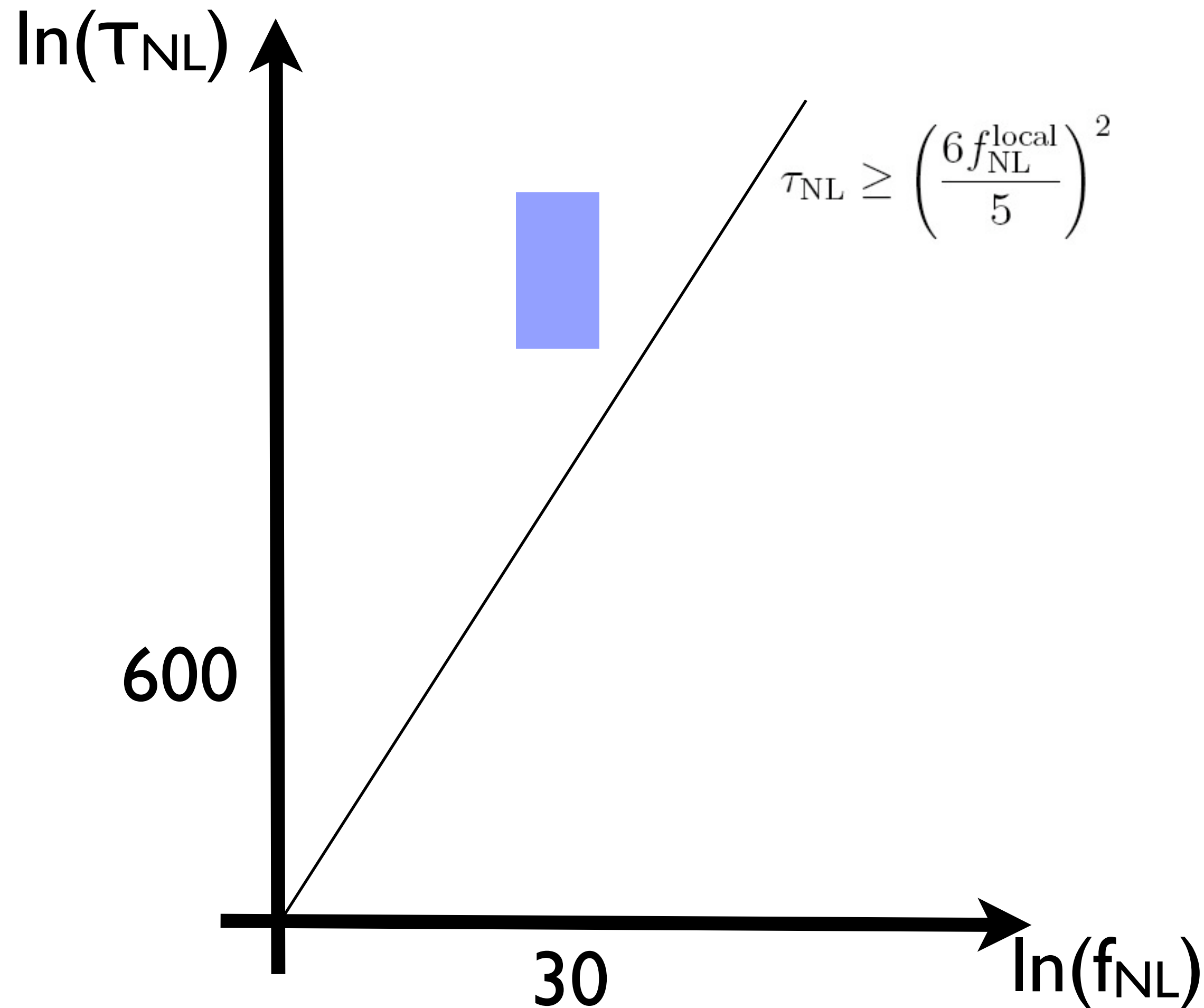
- The current limits from WMAP 7-year are consistent with single-field or multi-field models.
- So, let's play around with the future.

Case A: Single-field Happiness



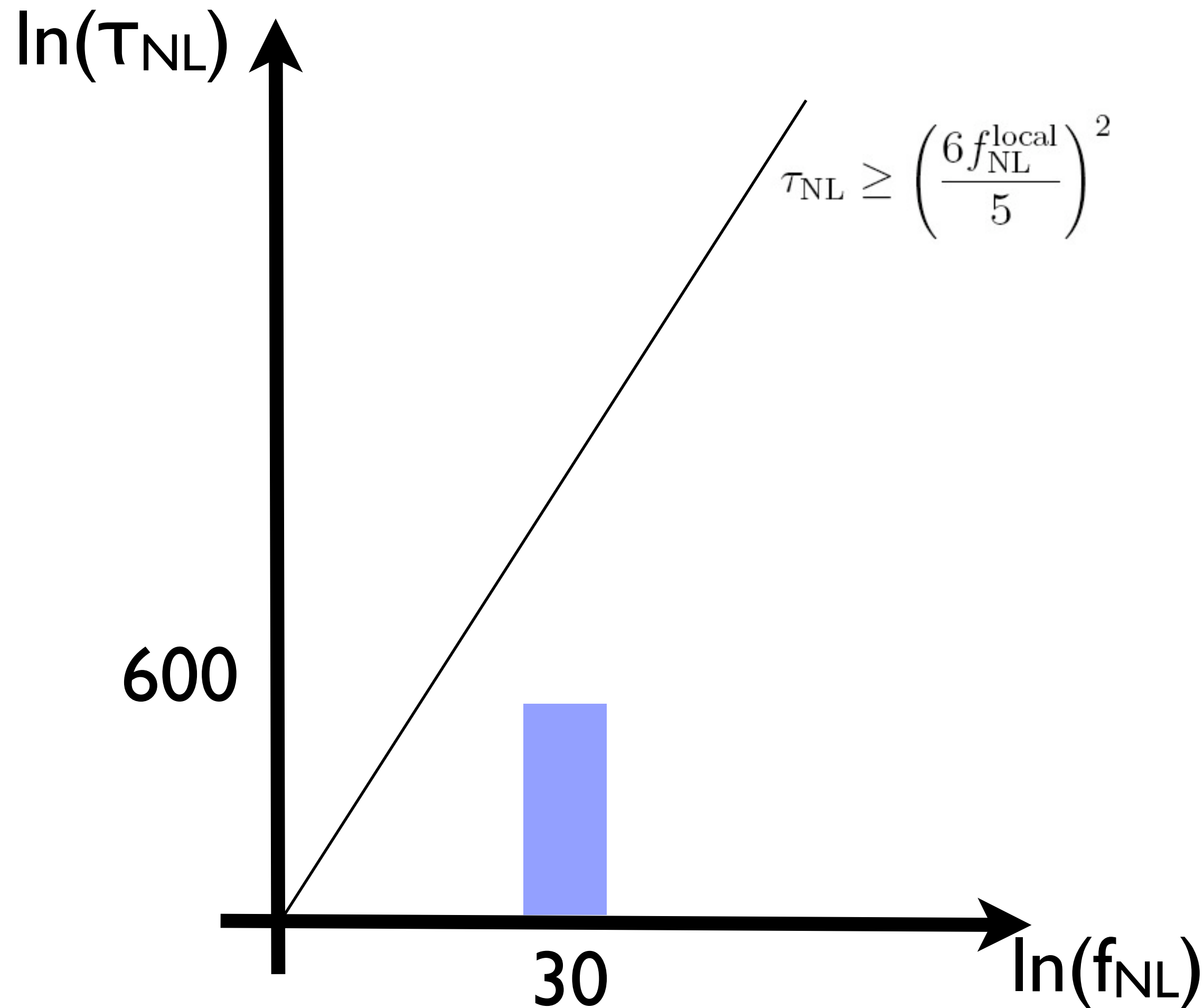
- No detection of anything after Planck. Single-field survived the test (for the moment: the future galaxy surveys can improve the limits by a factor of ten).

Case B: Multi-field Happiness



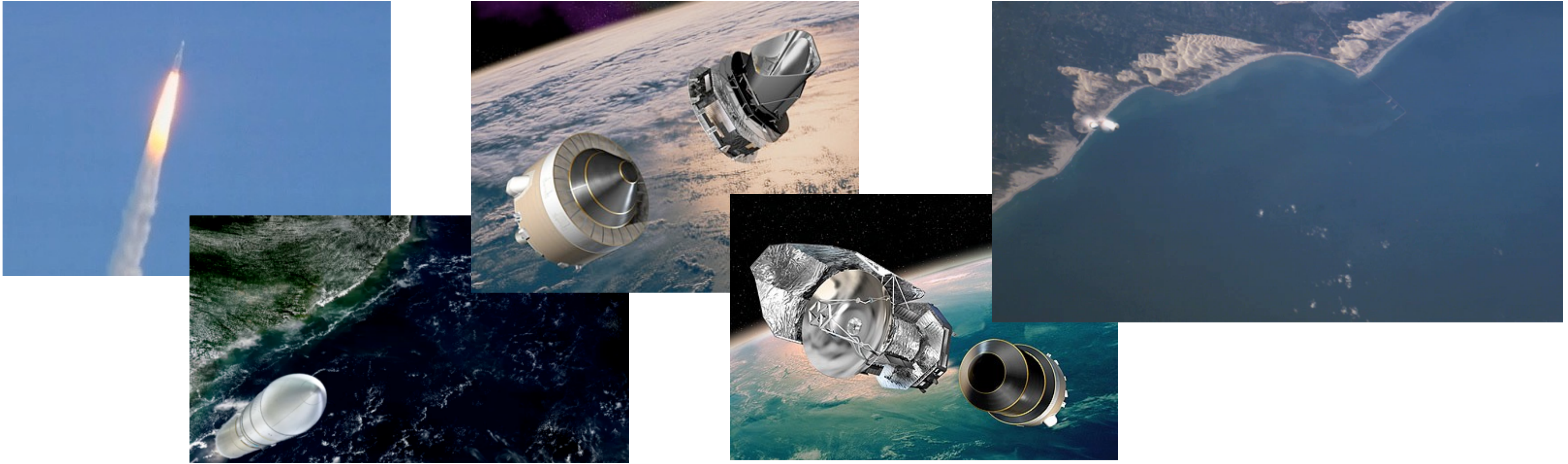
- f_{NL} is detected. Single-field is dead.
- But, τ_{NL} is also detected, in accordance with the Suyama-Yamaguchi inequality, as expected from most (if not all - left unproven) of multi-field models.

Case C: Madness



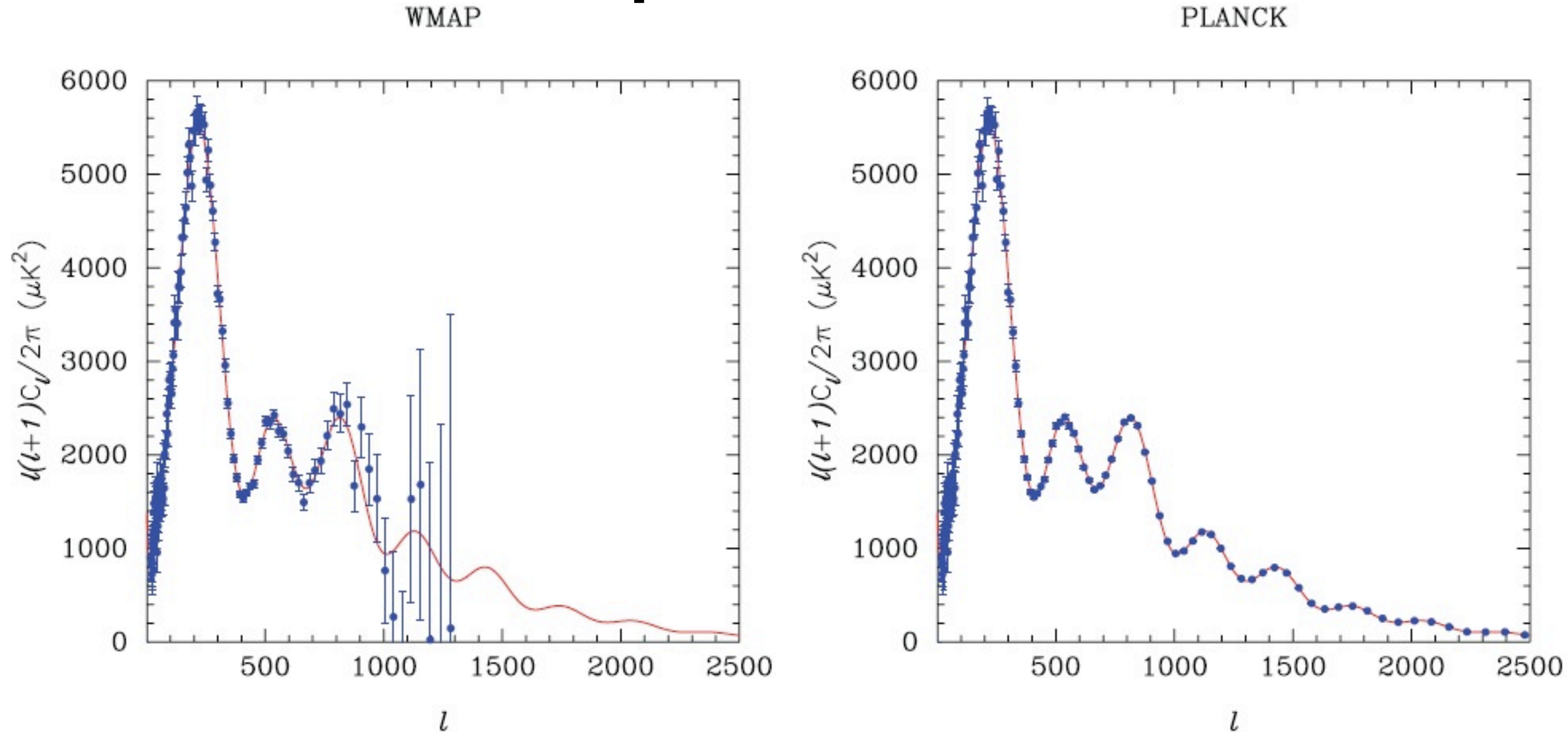
- f_{NL} is detected. Single-field is dead.
- But, τ_{NL} is **not** detected, inconsistent with the Suyama-Yamaguchi inequality.
- (With the caveat that this may not be completely general) BOTH the single-field and multi-field are gone.

Planck Launched!



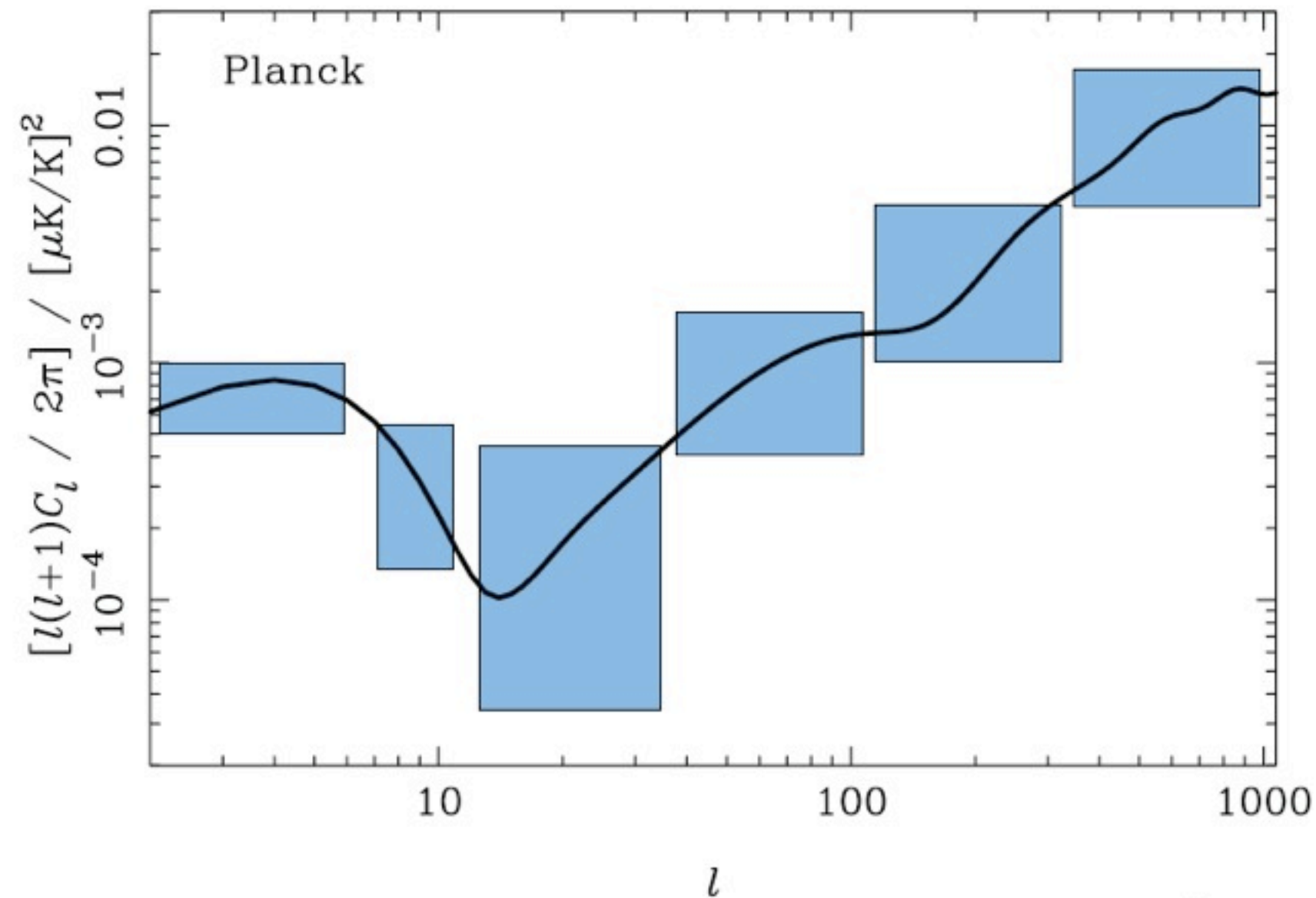
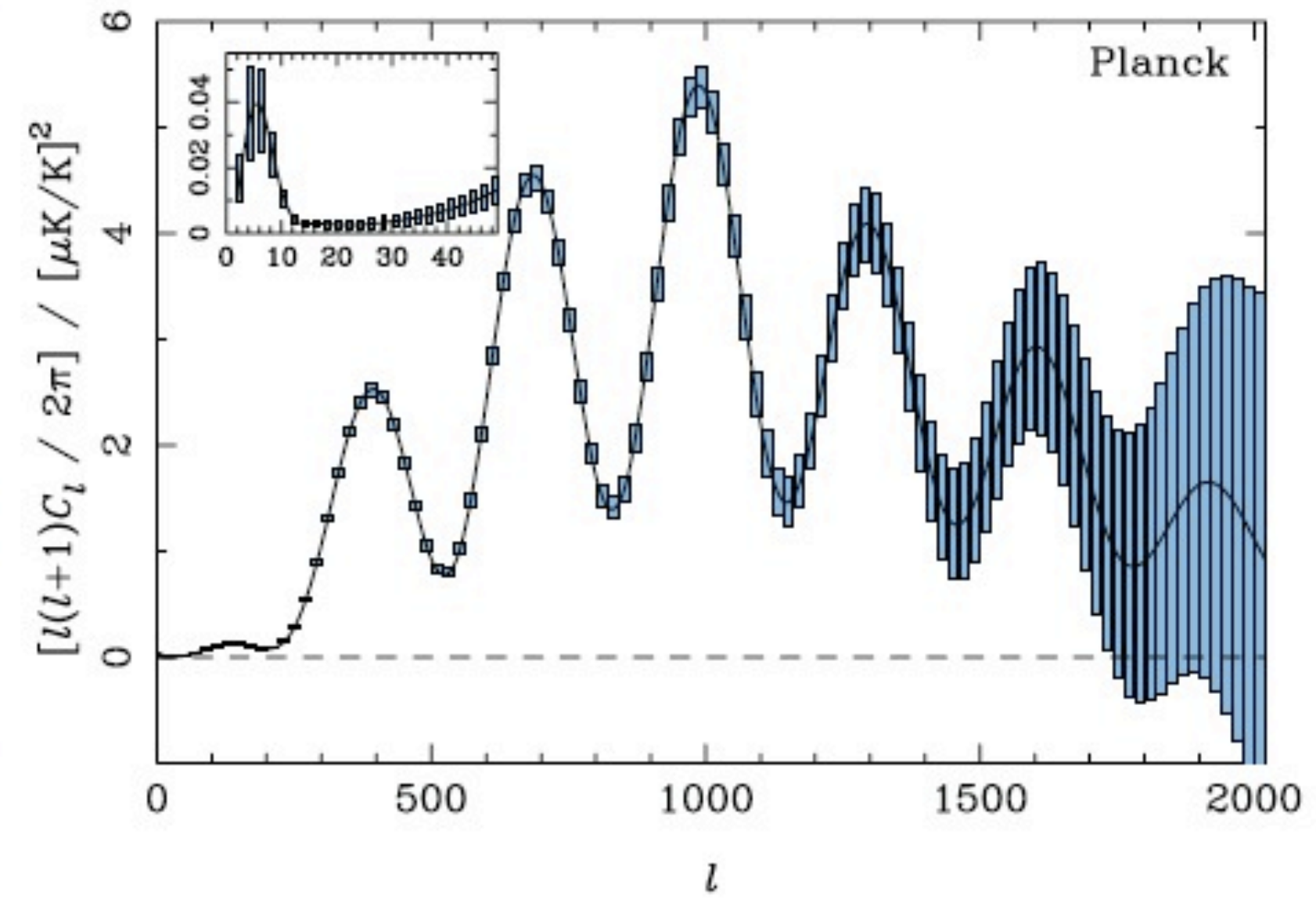
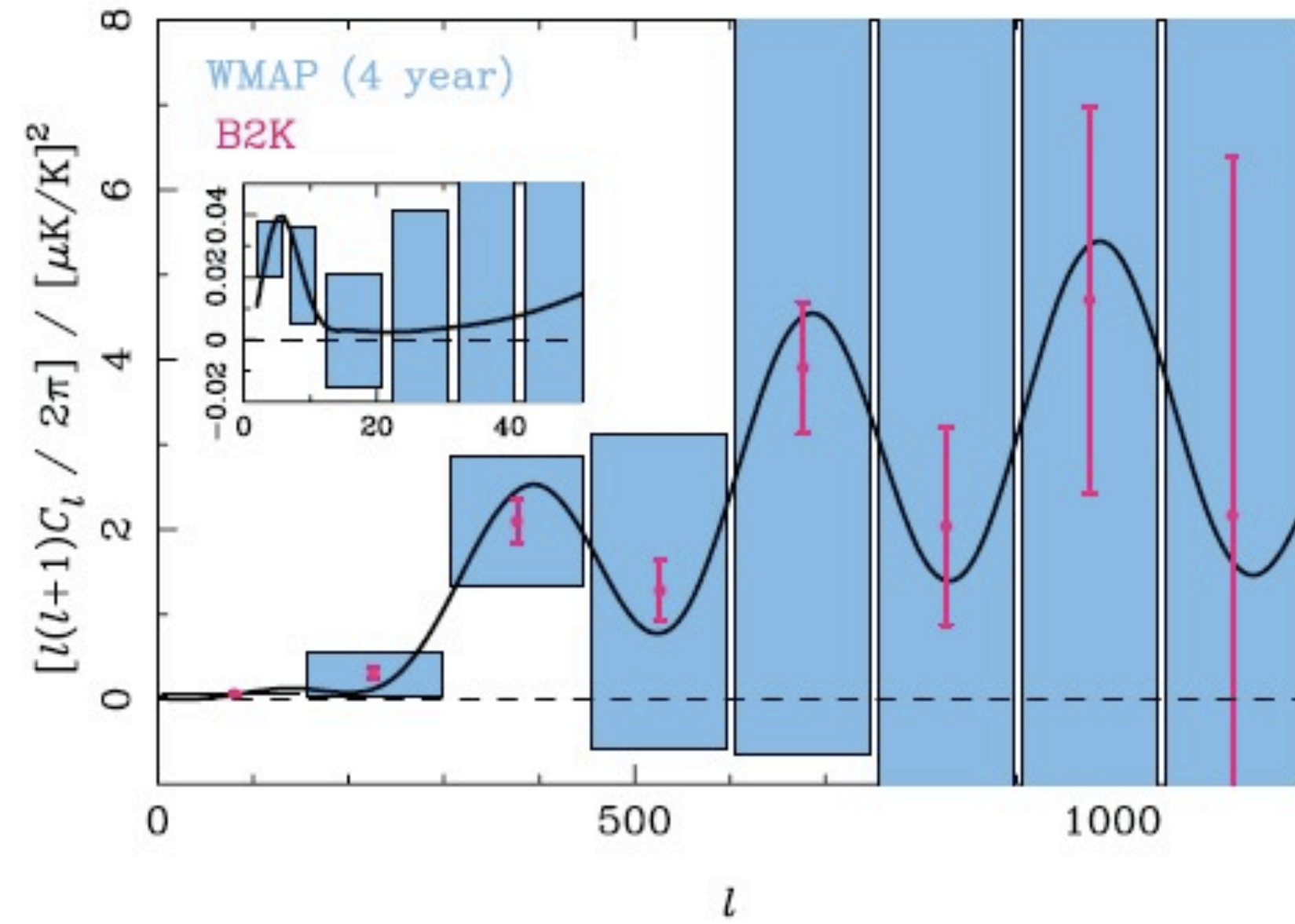
- The Planck satellite was successfully launched from French Guiana on May 14, 2009.
- Separation from the Herschel satellite was also successful.
- Planck has mapped the full sky already - results expected to be released in December, 2012.

Planck: Expected C_l Temperature



- WMAP: $l \sim 1000 \Rightarrow$ Planck: $l \sim 3000$

Planck: Expected C_l Polarization



- (Above) E-modes
- (Left) B-modes ($r=0.3$)