

# The particle-without-particle approach to the self-force problem

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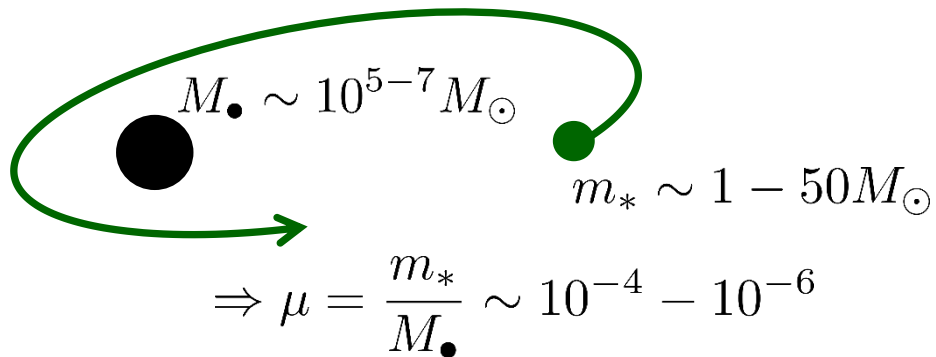
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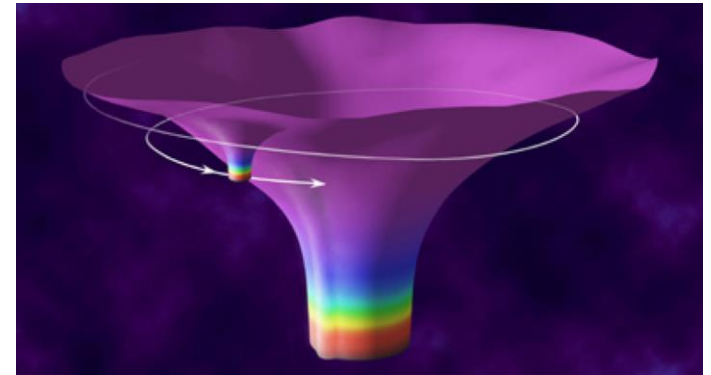
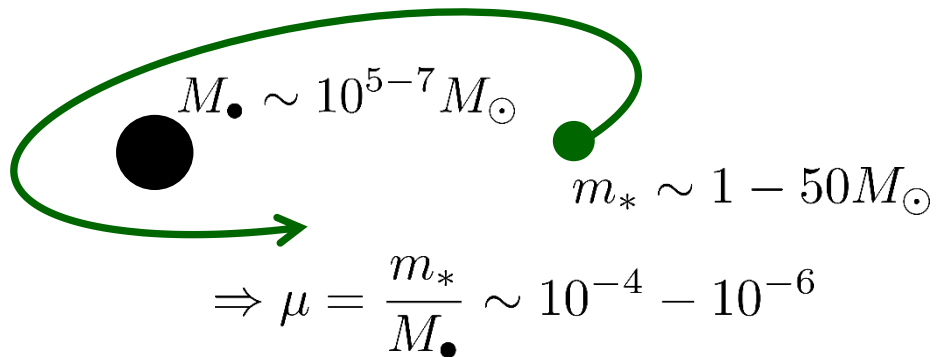
Paris, France — 27 June 2017

# Introduction

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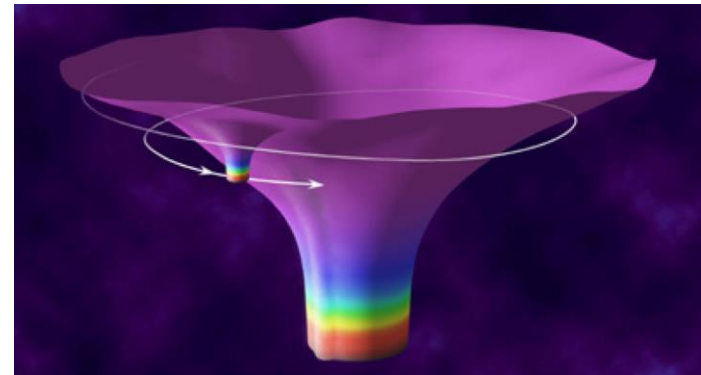
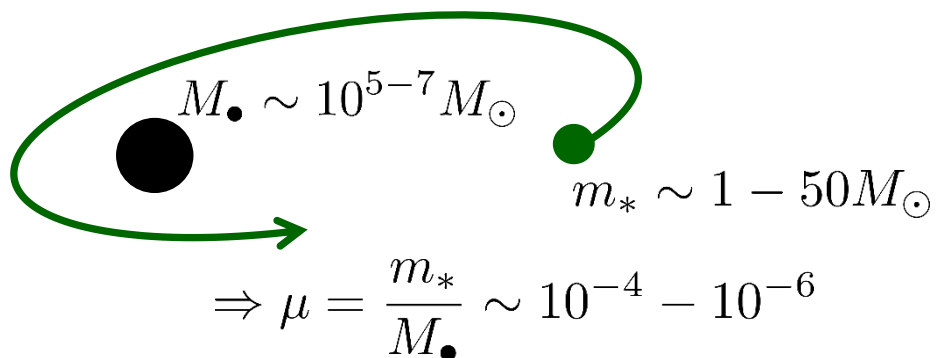
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[arXiv:1305.5720]

- The detection of GWs from EMRIs will rely crucially on an accurate computation of the *gravitational self-force*: the most popular approach is to model  $m_*$  as a point source (delta function) creating a backreaction in the geometry of  $M_\bullet$ .

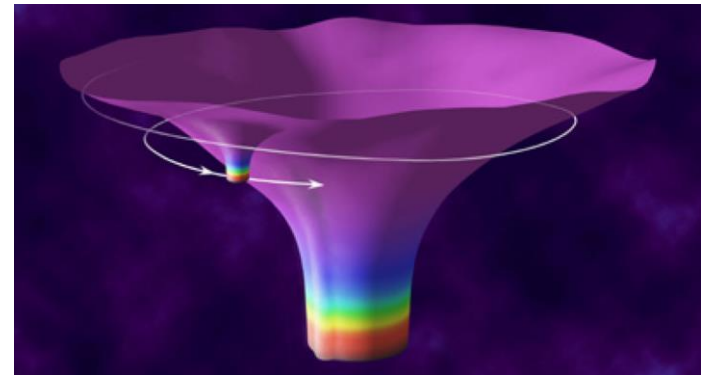
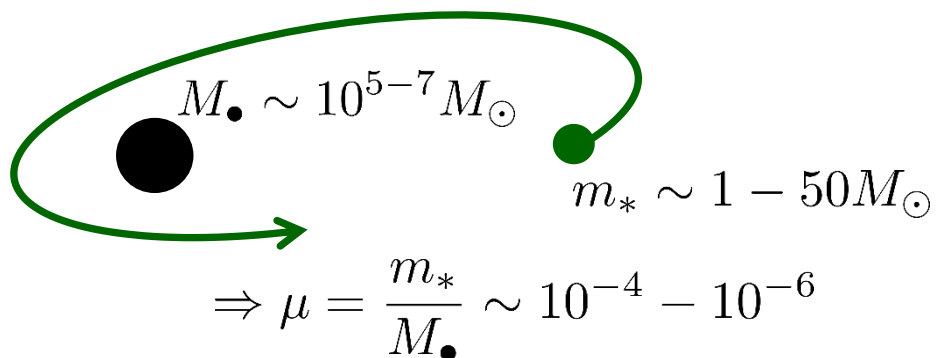
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- The theoretical formalism to compute the self-force has been largely established, e.g. [S. Gralla and R. Wald, CQG 25, 205009 (2008)], but its mathematical implementation is still under development; we use the *Particle-without-Particle (PwP)* method.

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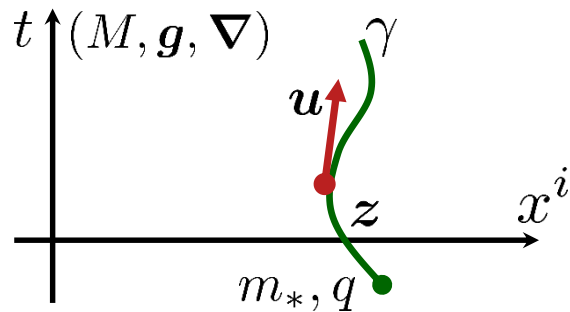
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- A helpful testbed for the gravitational self-force is the *scalar* self-force—we tackle this using the PwP method in the *frequency domain*.

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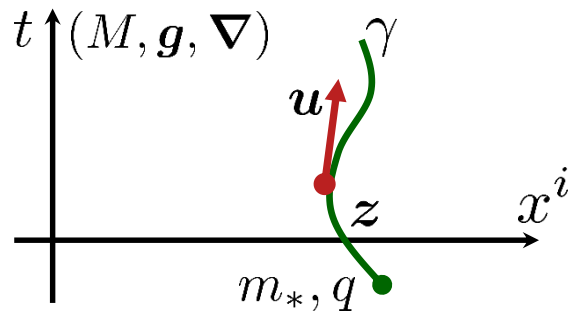
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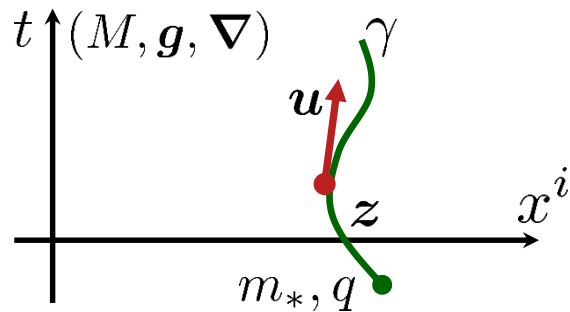


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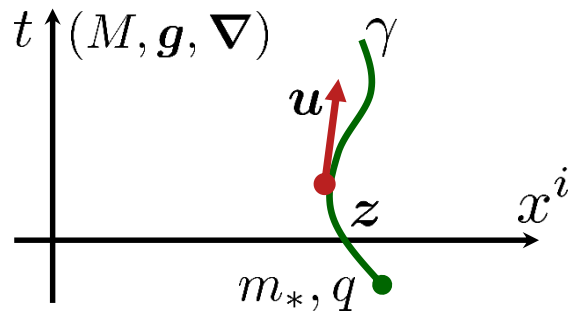
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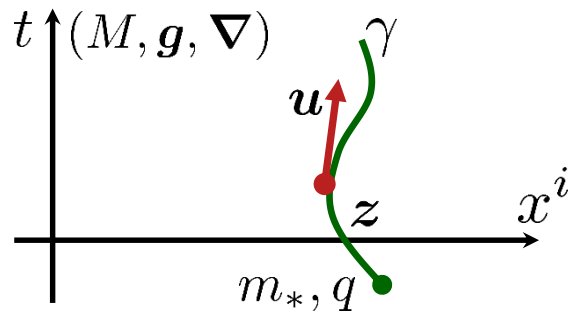
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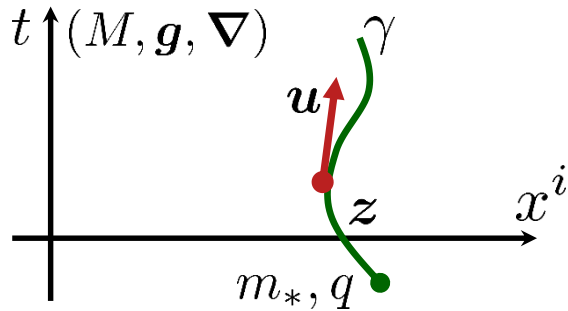
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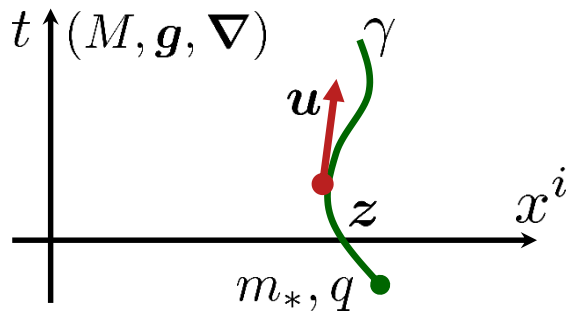
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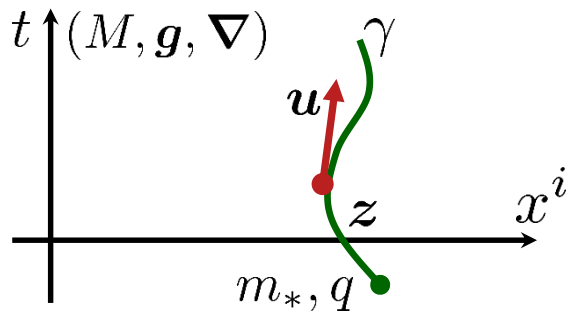
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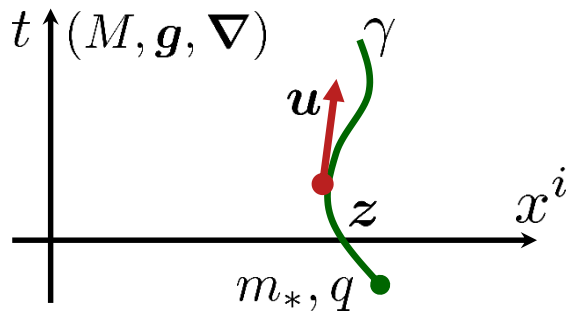
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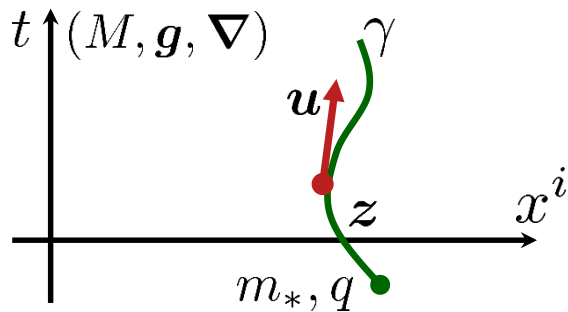
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- Once the field is solved for, its singular part must be subtracted (via "mode-sum regularization" [L. Barack and A. Ori, PRD 61, 061502 (2000)]).

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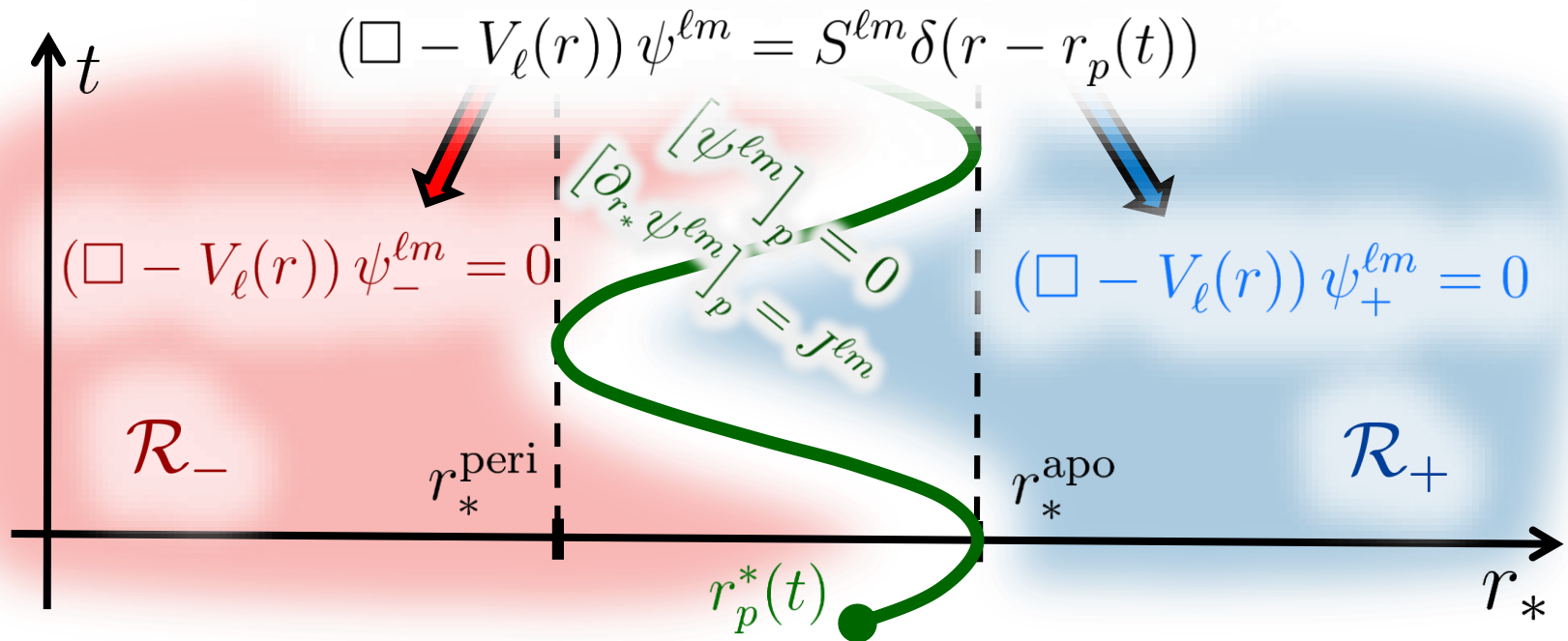
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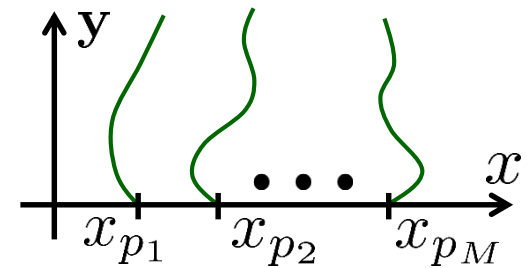


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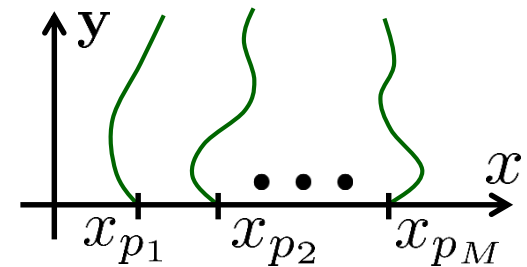
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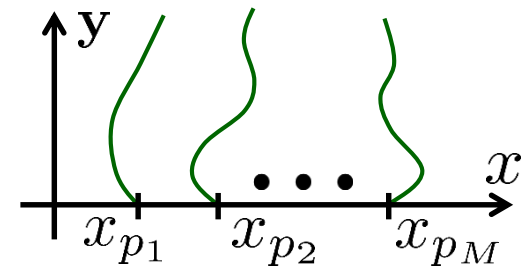


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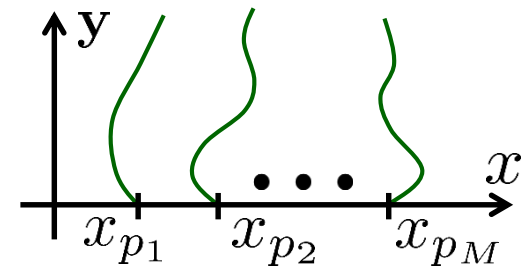
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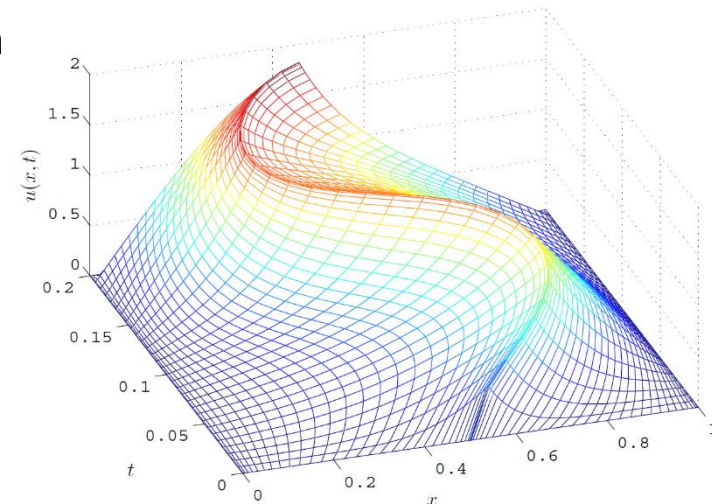
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- **Example:** Heat equation with constant source supported at a sinusoidally moving particle (using a Chebyshev-Lobatto grid in space and a standard finite-difference scheme in time).



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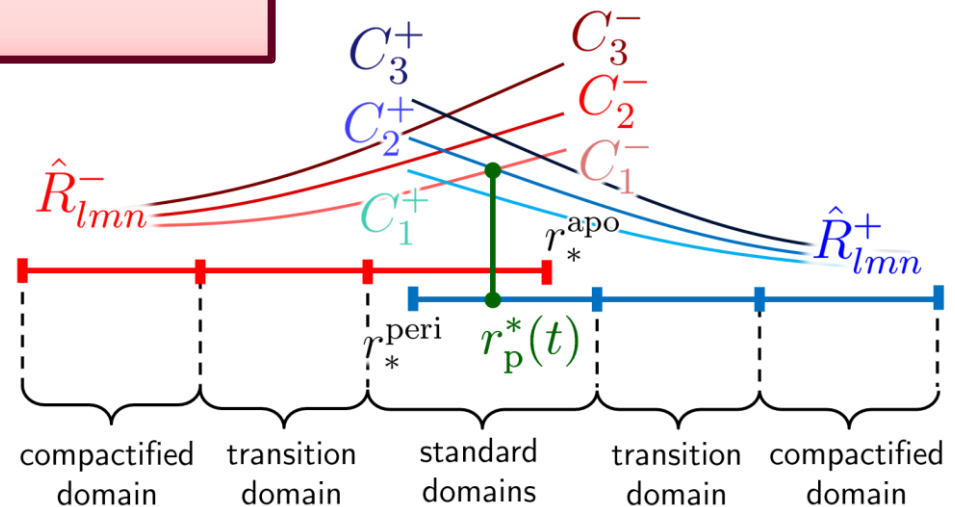
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- We use a pseudospectral collocation method to find the homogeneous numerical solutions  $\hat{R}_{\ell mn}^{-}$  and  $\hat{R}_{\ell mn}^{+}$  for arbitrary BCs, then use the jump conditions to get the true solution,  $R_{\ell mn} = C_{\ell mn}^{-} \hat{R}_{\ell mn}^{-} \Theta_p^{-} + C_{\ell mn}^{+} \hat{R}_{\ell mn}^{+} \Theta_p^{+}$



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- For future work, another objective is to also extend this method to *rotating black holes (Kerr)*.



**Thanks for your attention!**