

Clustering properties of LAEs in MUSYC (ECDF-S)

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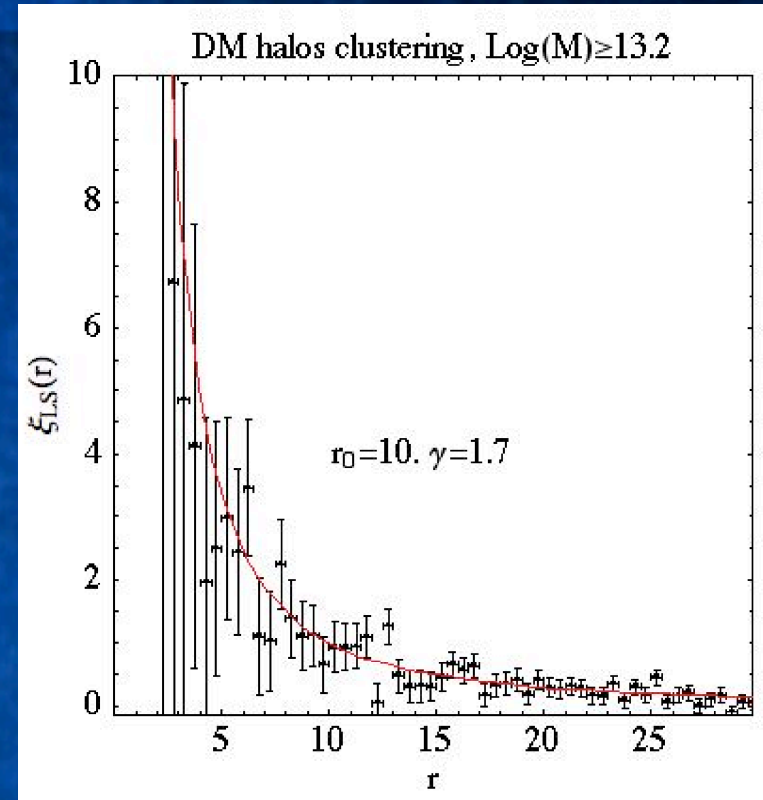
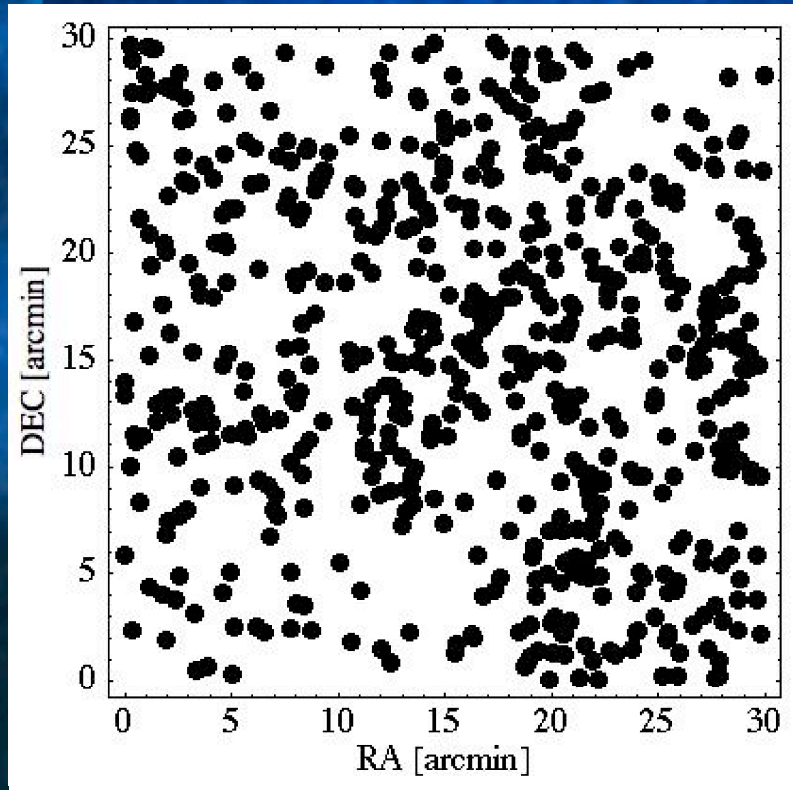
Outline

- Introduction
- Observational Dataset: MUSYC
- Clustering Analysis
 - Halo masses at $z=3.1$ and 2.1
 - Evolution
- Summary and Conclusions

Introduction

- Lyman Alpha Emitters have shown significant clustering strength at high redshift...
 - But the bulk of them seem to have fairly blue, low stellar mass SEDs...
 - MUSYC is optimized to detect star-forming galaxies at $2 < z < 3$
- ⇒ LAEs at $2 < z < 3$ in MUSYC are great targets for clustering analysis!!

Halo Mass and clustering



Clustering measure \Rightarrow Two-point correlation function

MUSYC Dataset

- Optical and NIR photometry, 4 fields, 30'x30' each.
- Spectroscopy:
 - Magellan+IMACS, >500 with spec ID so far
- ECDF-S has unique 250ks Chandra data
- SDSS 1030 has 105ks XMM data

Field	# Obj.	U	B	V	R	I	z	J	H	K	NB5000	X-Ray [erg/(s cm ²)]
E-CDFS	84410	26.0	26.9	26.4	26.4	24.6	23.6	24.3	23.8	23.4	25.5	2x10 ⁻¹⁶
SDSS 1030+05	69619	25.8	26.0	26.2	26.0	25.4	23.7	24.1	23.9	23.3	24.8	8x10 ⁻¹⁶
Cast1255+01	60344	26.0	26.2	26.1	26.0	25.0	24.1	24.0	22.8	23.0	24.4	
E-HDFS	62968	26.0	26.1	26.0	25.8	24.7	23.6	24.3	23.4	23.4	24.1	

10x10 deep fields

LAE @ $z=3.1$ selection in MUSYC

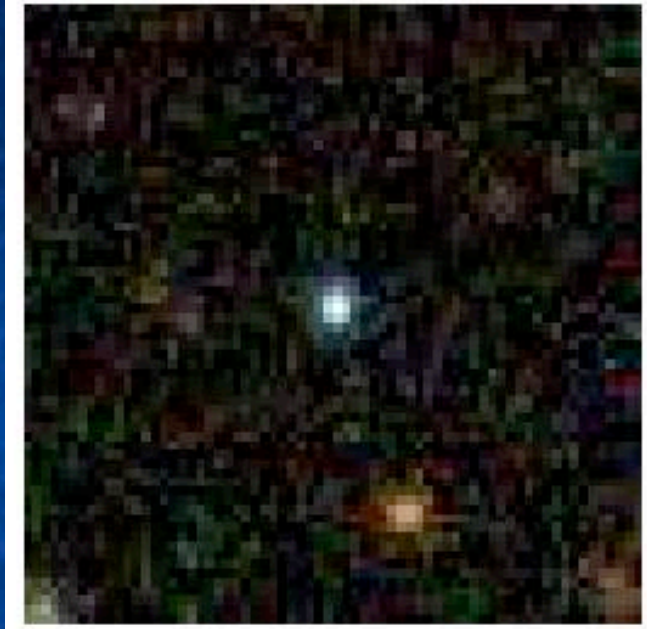
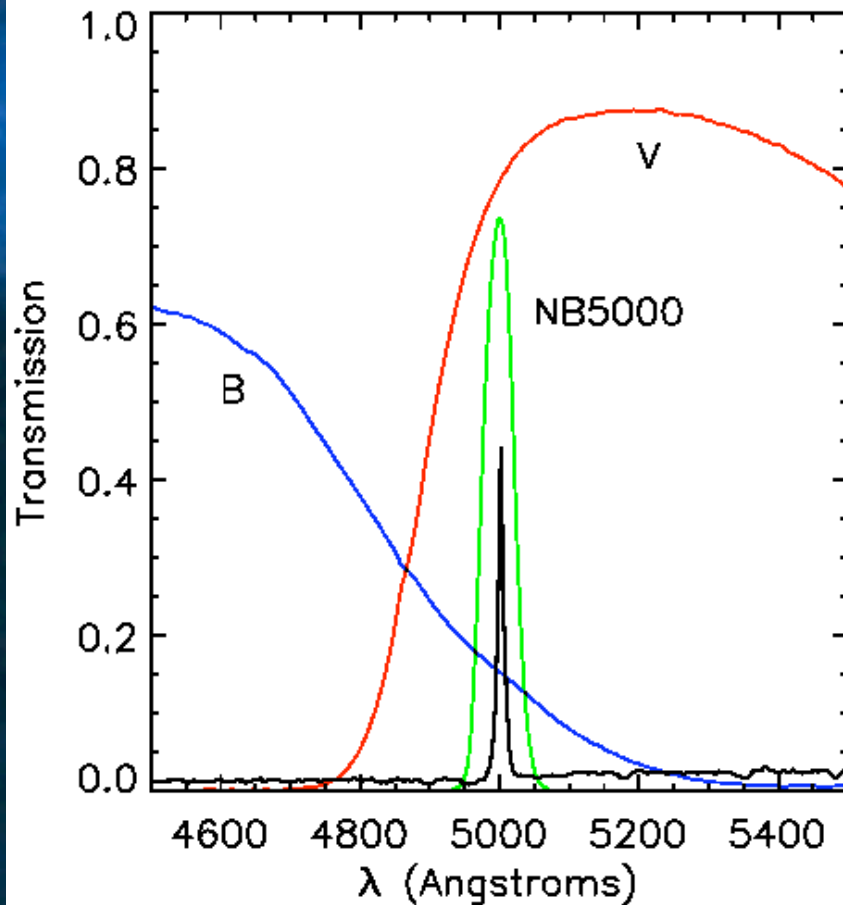
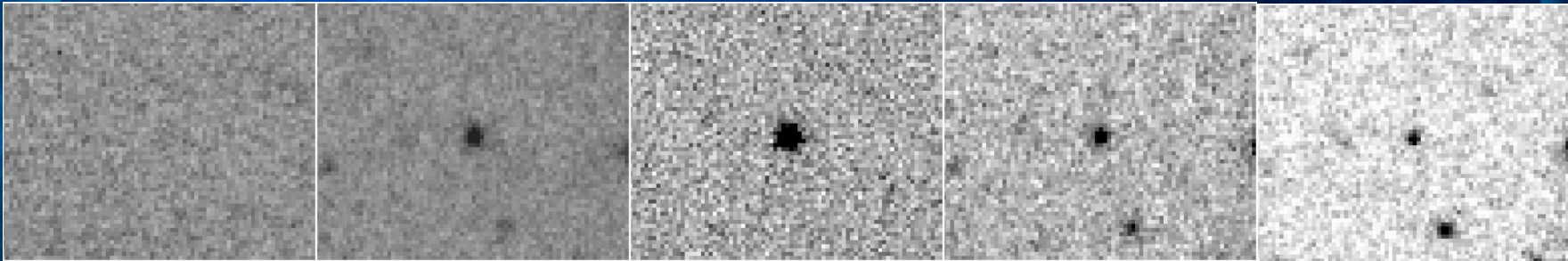
U

B

NB5000

V

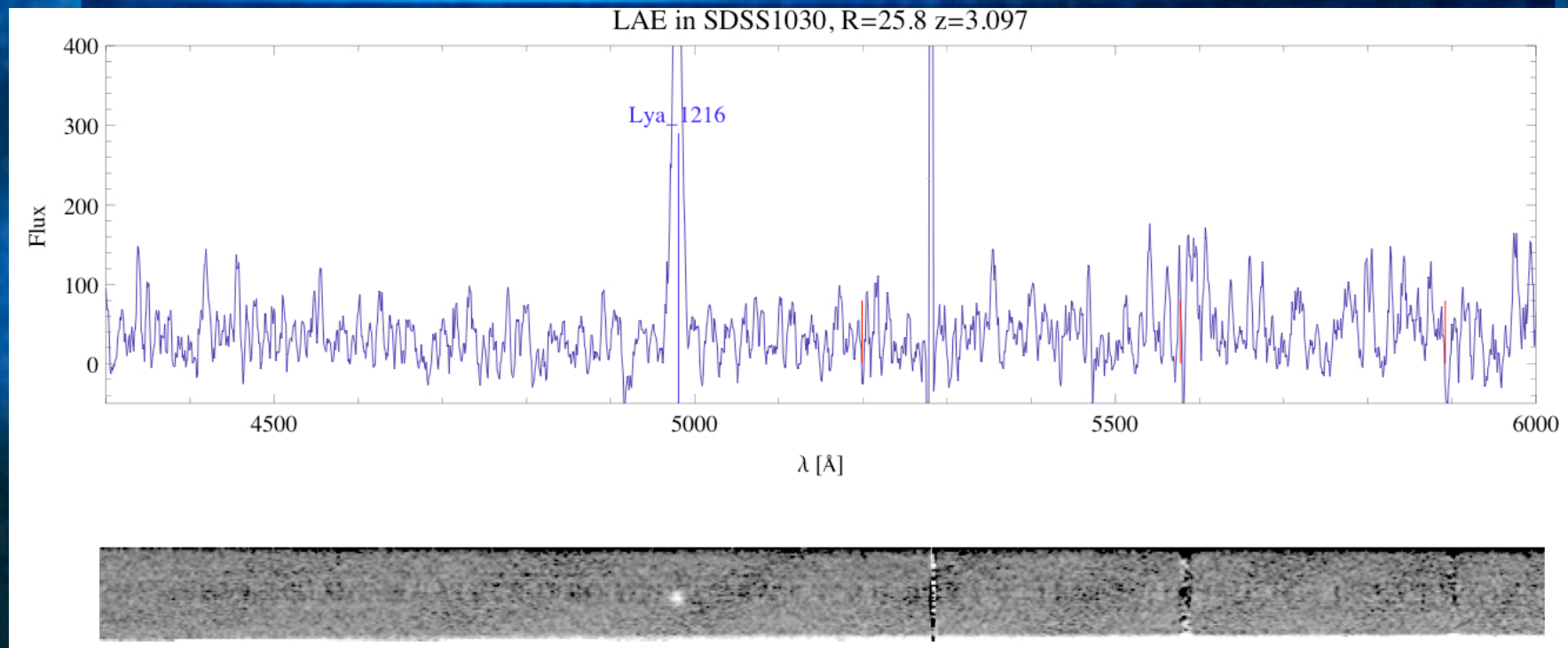
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Sample of 160 LAEs, for more details,
see:

[Gronwall et al 2007, ApJ 667, 79](#)

LAE spectrum (IMACS)

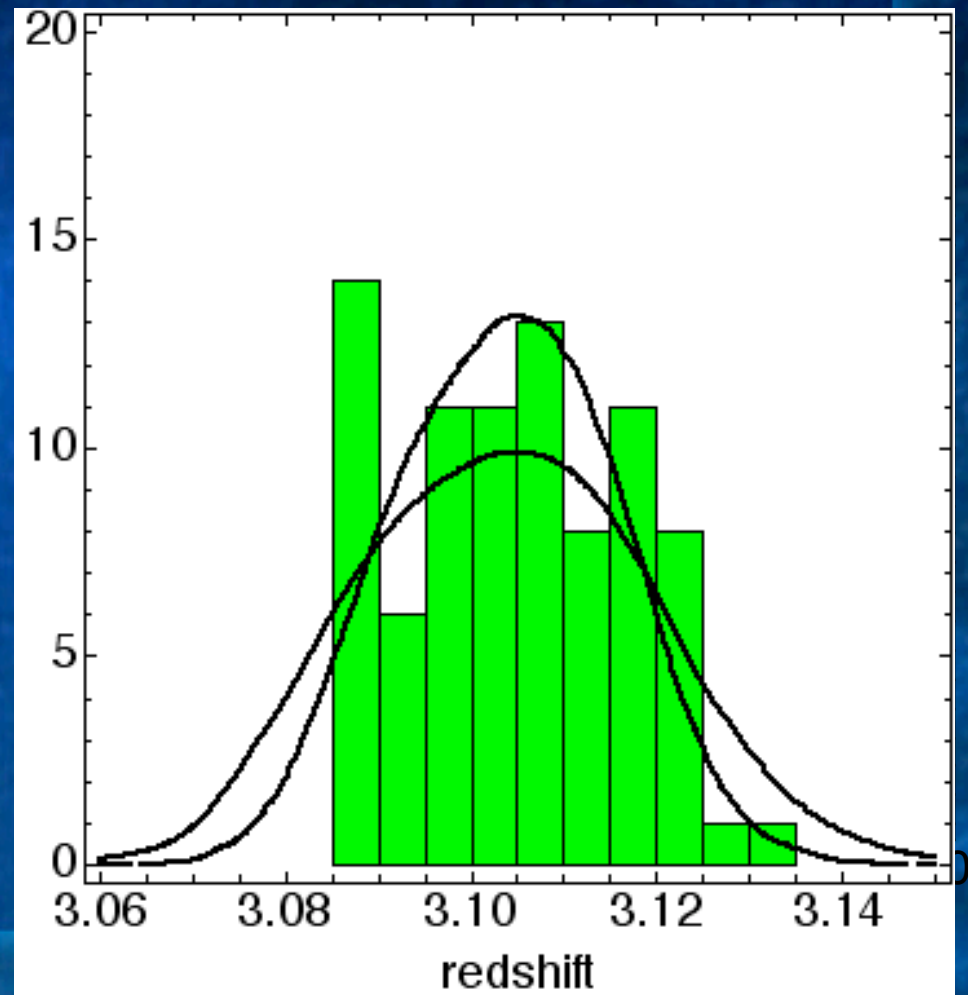


SED properties ($z=3.1$)

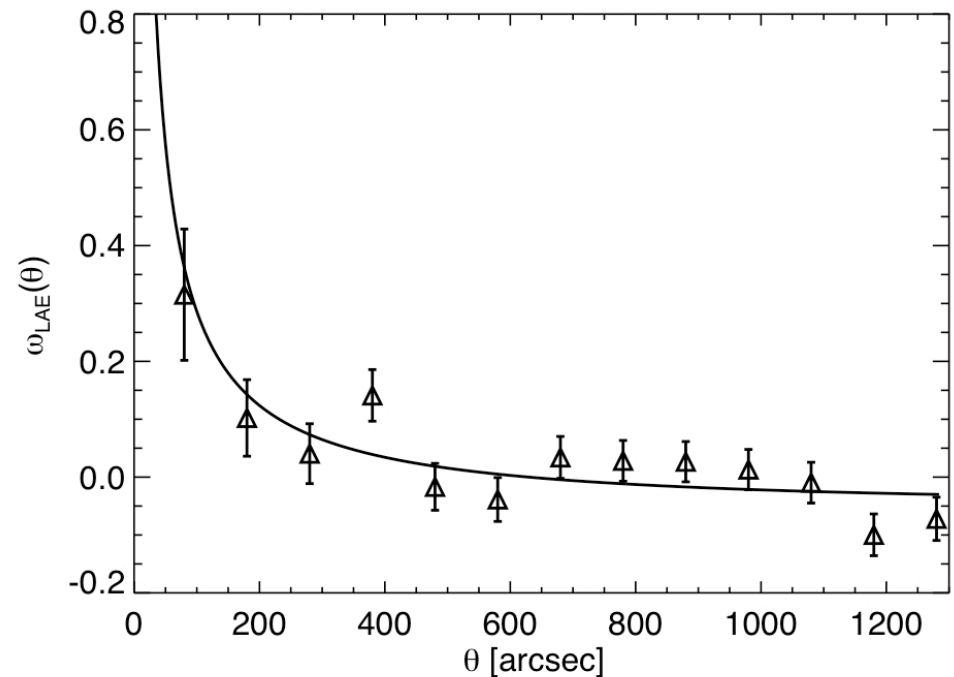
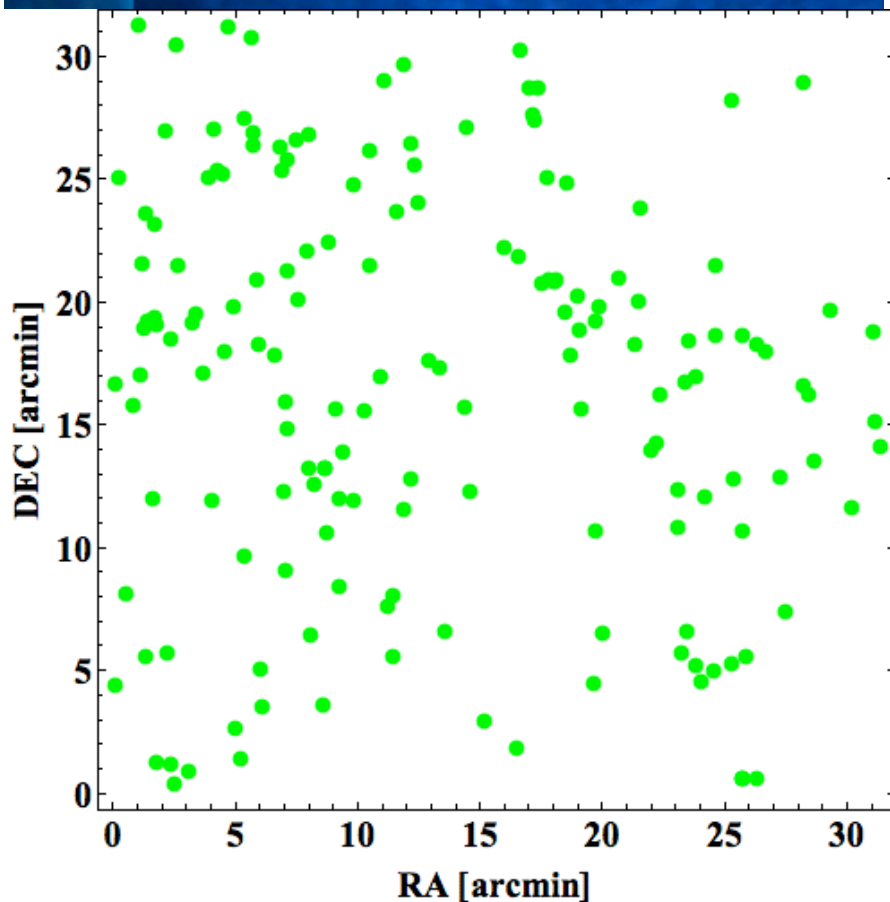
- From SED fitting (Gawiser et al. 2007), bulk of LAEs at $z\sim 3$ seem to have:
 - Low stellar masses, $\sim 10^9 M_{\text{sun}}$
 - Low SFR, $\sim 2 M_{\text{sun}}/\text{yr}$
 - Nearly no dust

Spectroscopic confirmations at $z=3.1$

- From 3 MUSYC fields:
⇒ 74 LAEs



LAEs at $z=3.1$

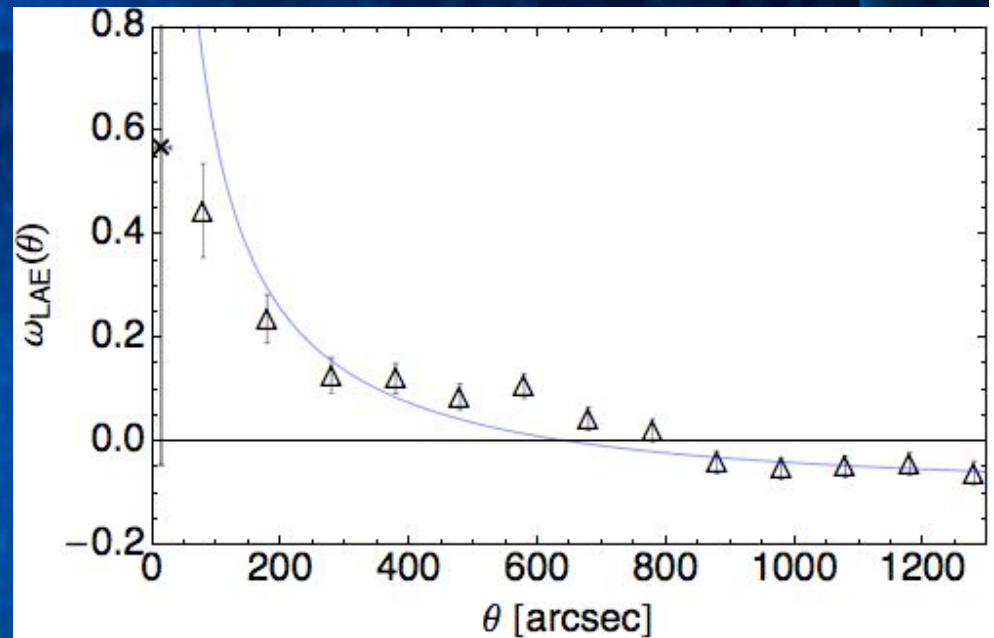
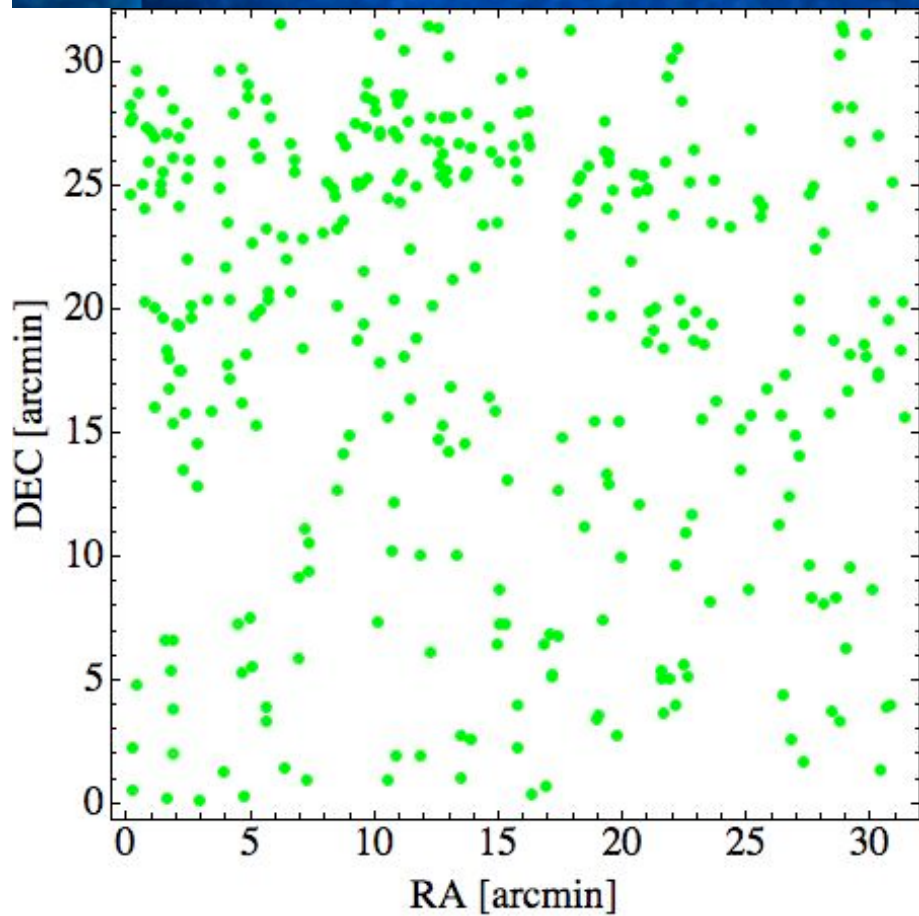


160 LAEs

Best fit: $r_0 = 3.6 \pm 1$ Mpc \Rightarrow bias = 1.9 ± 0.4
 $\Rightarrow \log M = 10.8^{+0.5}_{-0.9}$ oc. frac. $\sim 5\%$

Gawiser, Francke et al. 2007, ApJ 671, 278

LAEs at $z=2.1$



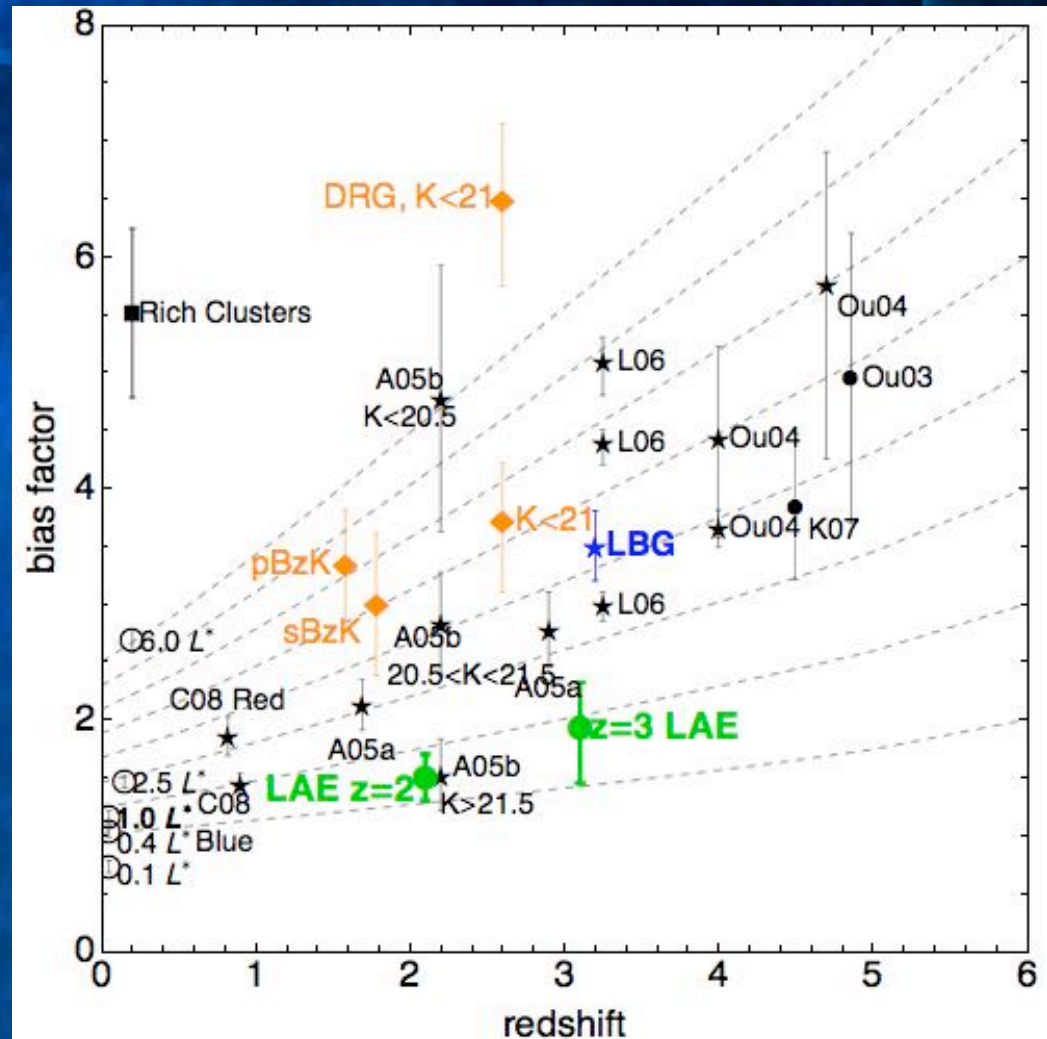
350 LAEs

Best fit: $r_0 = 3.7 \pm 0.6$ Mpc \Rightarrow bias = 1.5 ± 0.2
 \Rightarrow $\log M = 11.0^{+0.4}_{-0.4}$ oc. frac. $\sim 10\%$

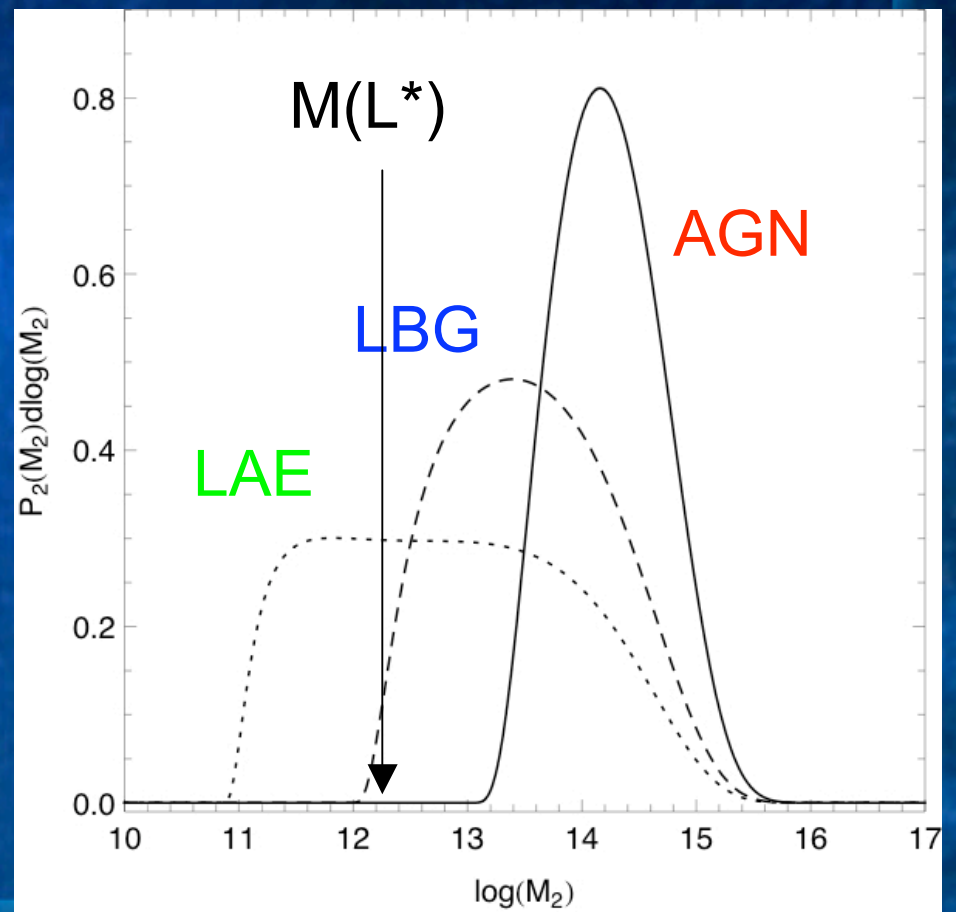
See Lucia Guaita's Poster

LAE clustering in context

- Both LAE samples are in the lowest mass range at $2 < z < 3$
- Could be progenitors of many L^* galaxies at $z=0$



Distribution at $z=0$ in terms of halo mass...



Summary/Conclusions

- LAEs show fairly weak clustering both at $z=3.1$ and 2.1 in ECDFS-MUSYC
- Appear to be progenitors of many of today's L^* galaxies
- Stay tuned... data from more fields coming to confirm and expand this result.
 - Narrow band imaging of other MUSYC fields
(Wider, more bright LAEs, better stats)
 - Follow-up spectroscopy
(Control better contamination frac. at $z=2.1$)