

SDSS-III and BOSS: The Ly α Forest in 3-D

David Weinberg, Ohio State University SDSS-III Project Scientist

SDSS-I (2000-2005), SDSS-II (2005-2008)

SDSS-III (2008-2014), four surveys on three science themes.

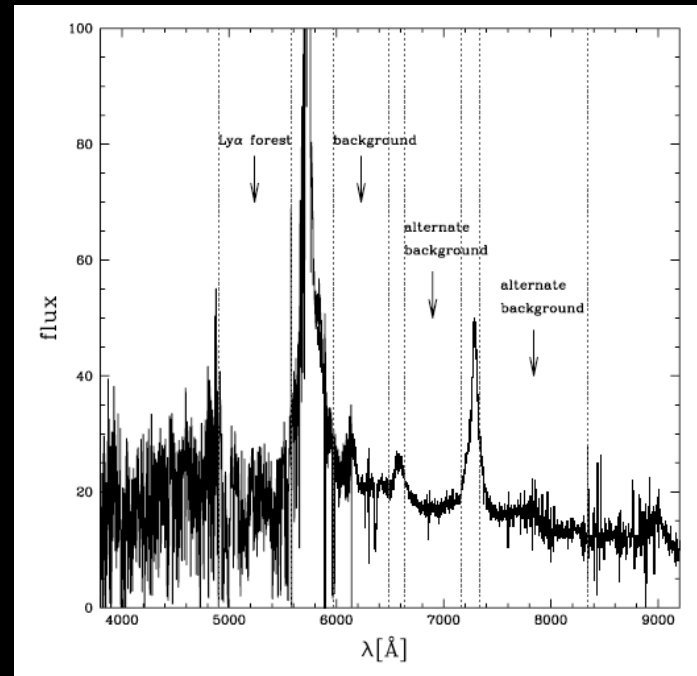
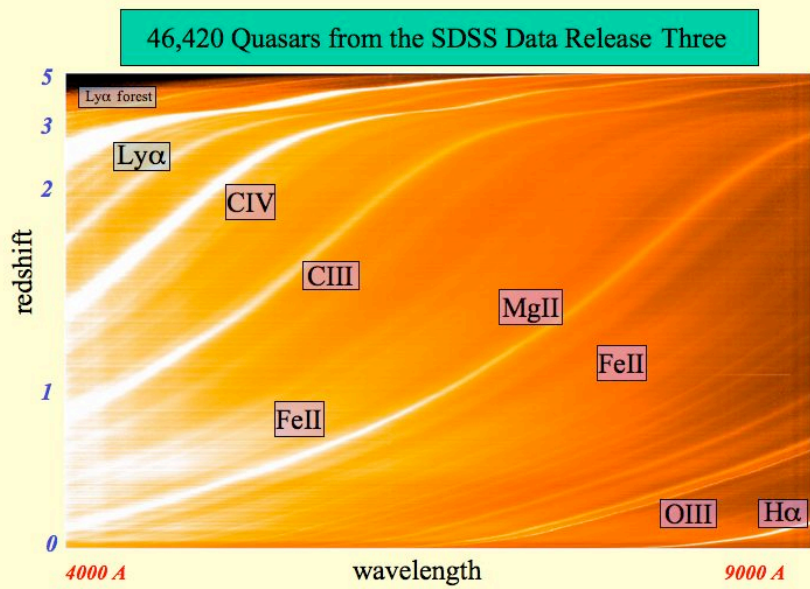
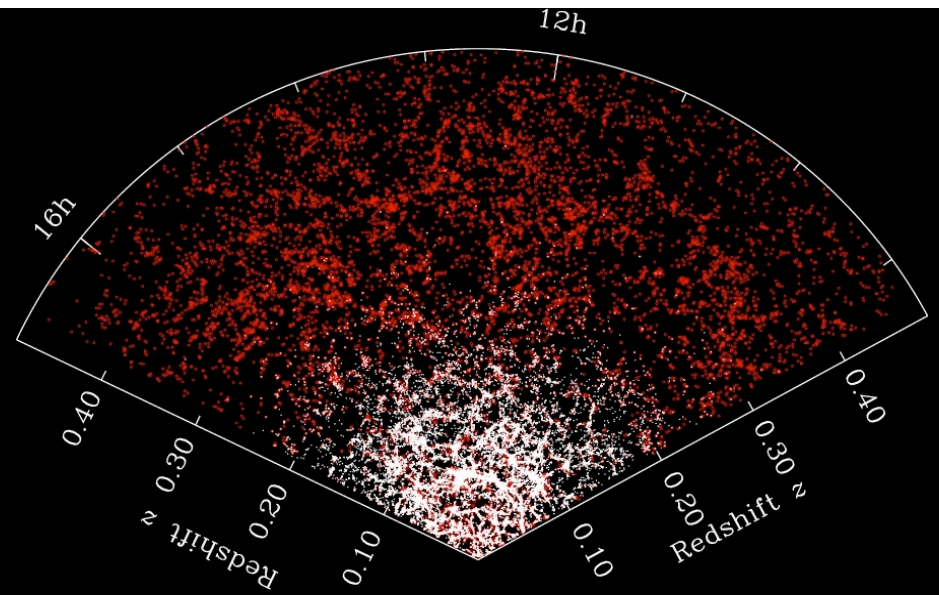
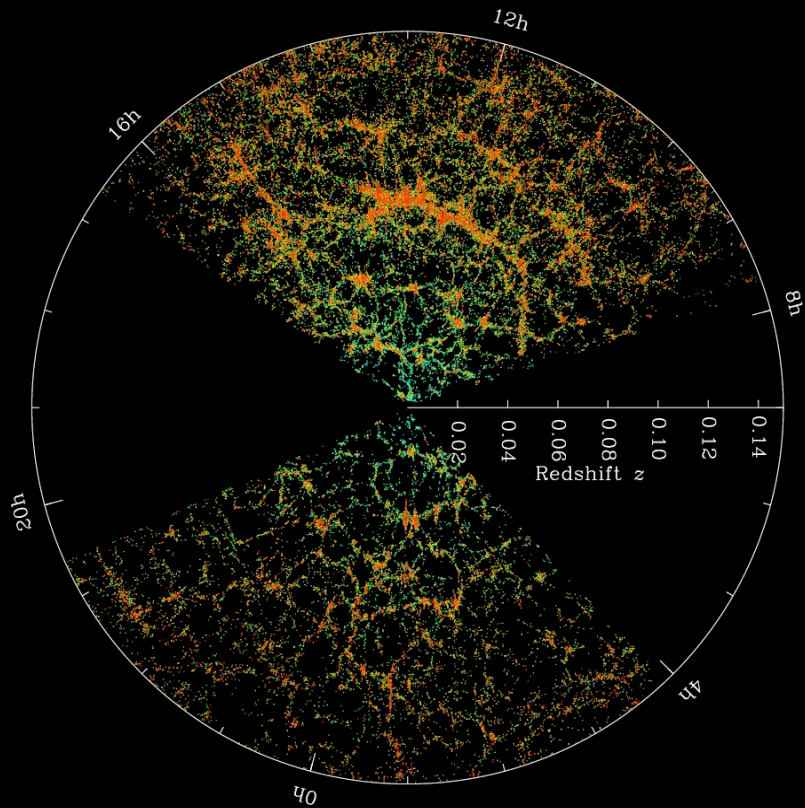
The Baryon Oscillation Spectroscopic Survey: Cosmology and dark energy with baryon acoustic oscillations (BAO). Redshifts of 1.5 million LRGs to $z=0.7$. Ly α forest spectra of 160,000 QSOs at $z\sim 2.5$.

SEGUE-2: Galactic structure and evolution. Optical spectra of 120,000 stars in the outer Galaxy, resolution $R \approx 2000$.

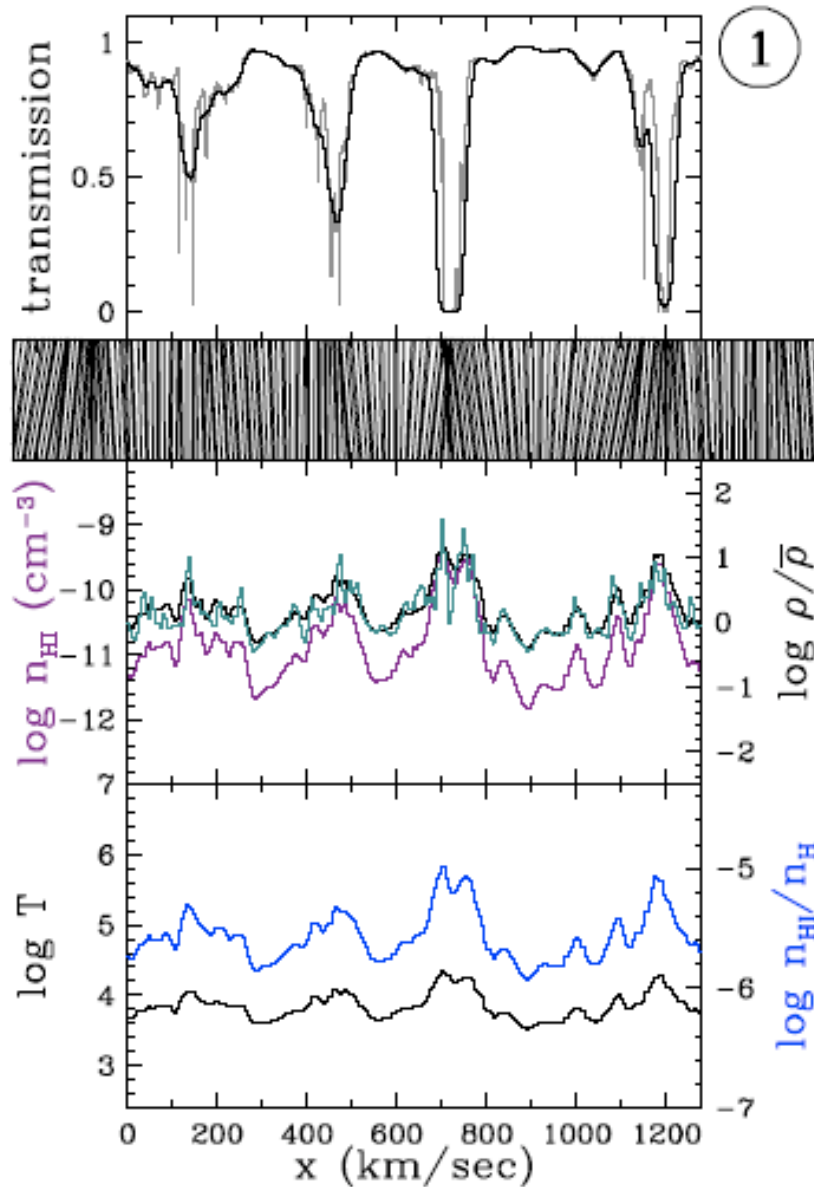
APOGEE: Galactic structure and evolution. H-band spectra of 100,000 stars throughout the Galaxy, resolution $R \approx 25,000$.

MARVELS: Extra-solar planets. Radial velocity monitoring of 10,000 stars, ~ 30 visits over 18 months, $10\text{-}20 \text{ m s}^{-1}$ precision.

Funding: Sloan Foundation, NSF, DOE, ~ 25 Participating Institutions.



Figures: M. Blanton, X. Fan, P. McDonald

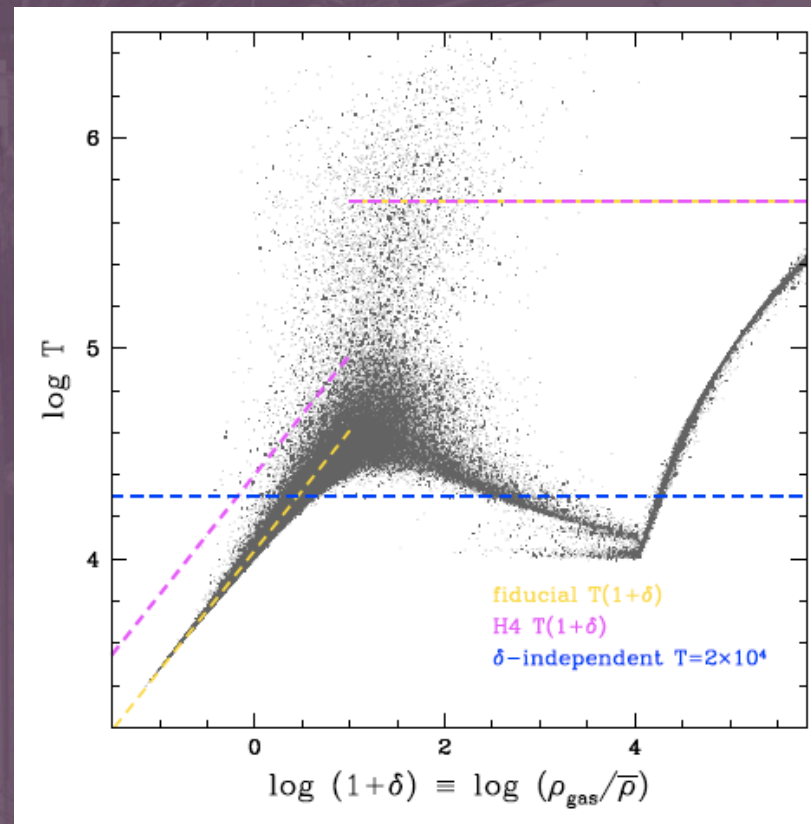


Fluctuating Gunn-Peterson effect:
 The Ly α forest arises in a smoothly
 fluctuating gas distribution.

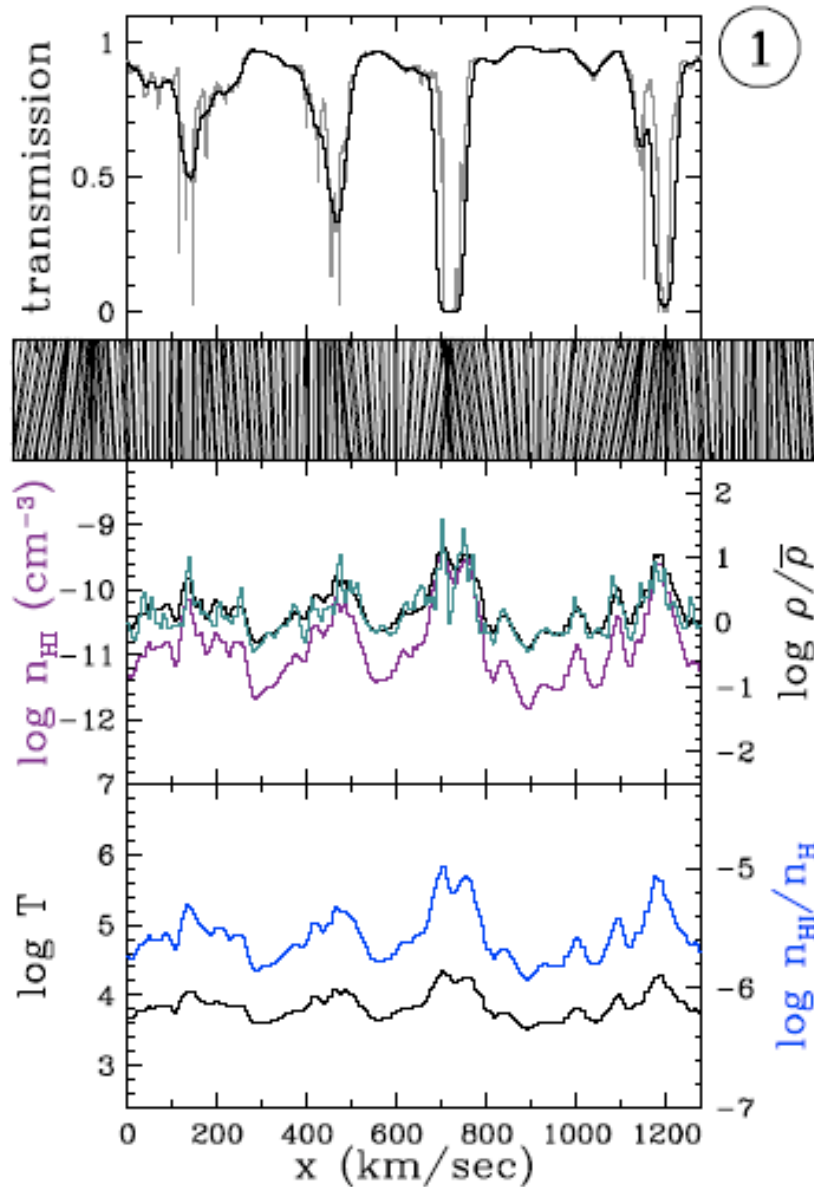
Approximately

$$F/F_c = e^{-\tau}$$

$$\tau \sim n_{\text{HI}} \sim n^2 T^{-0.7} = A (1+\delta)^\beta$$



M. Peeples et al, in prep.



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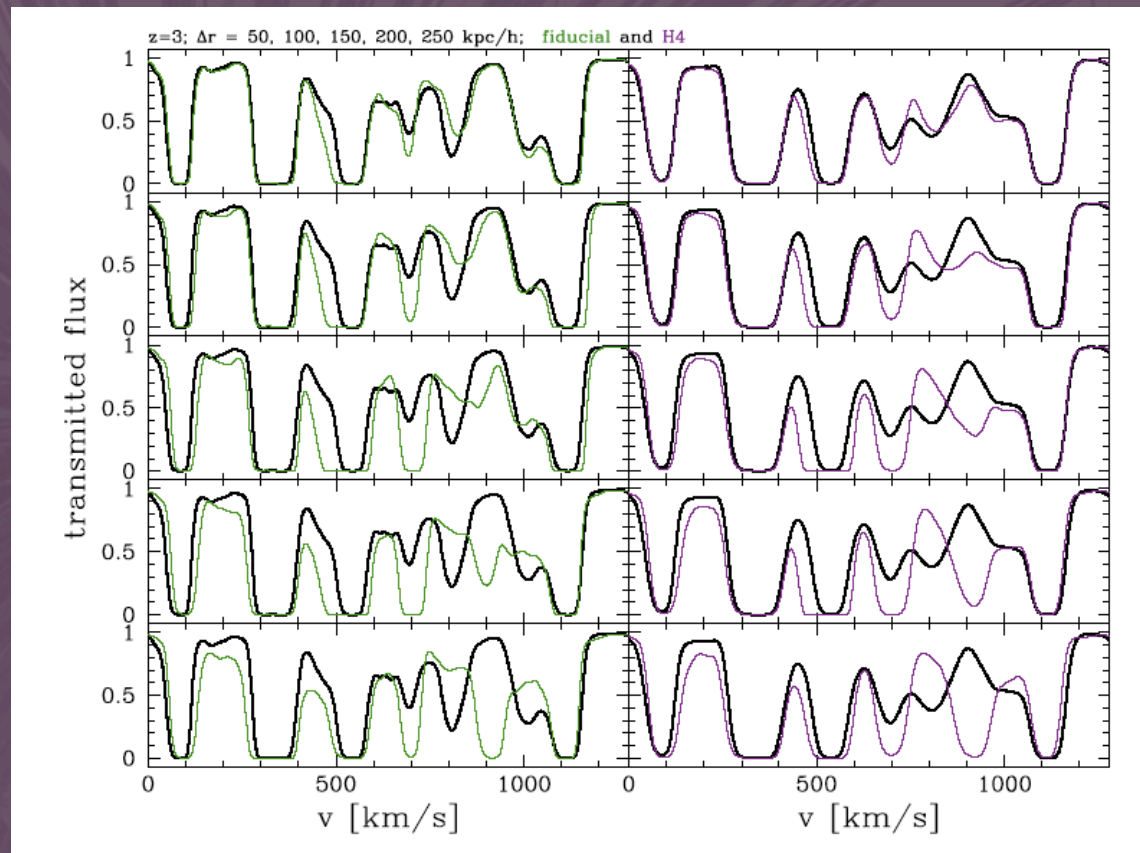
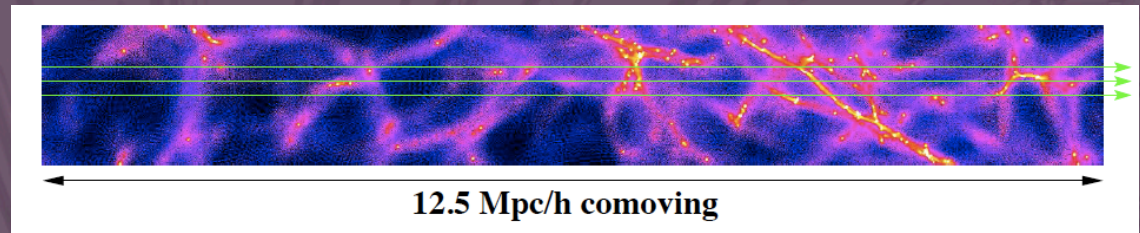
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Main Complications:

- Peculiar velocities
- Thermal broadening
- Pressure support of gas
- Shock heating

$\text{Ly}\alpha$ forest in 2.5-D

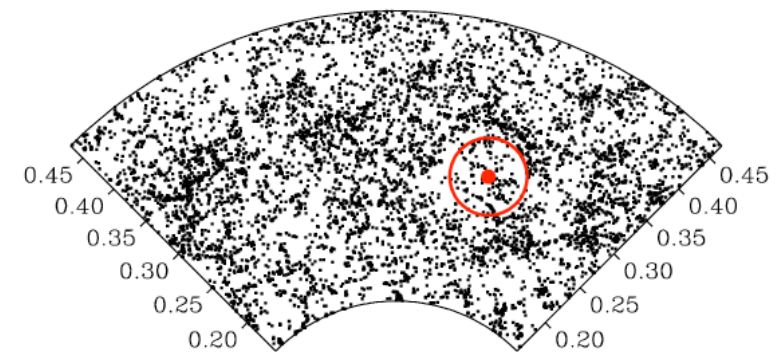
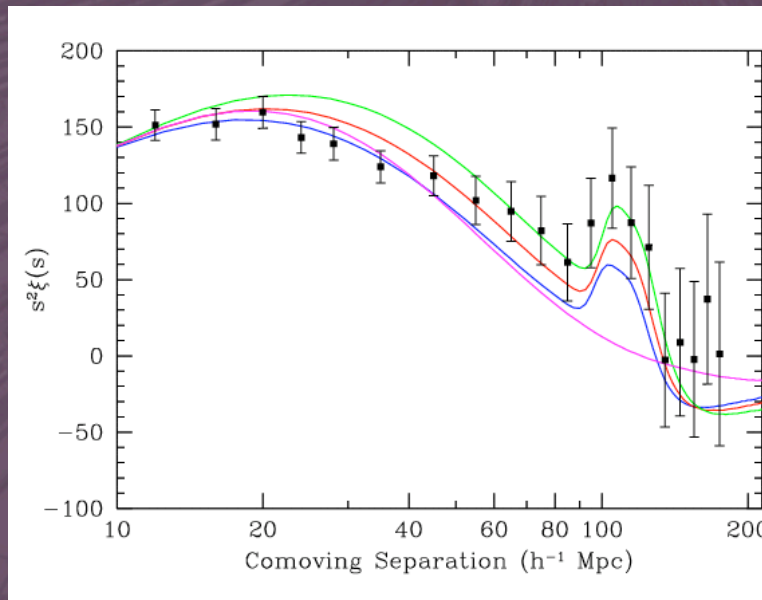
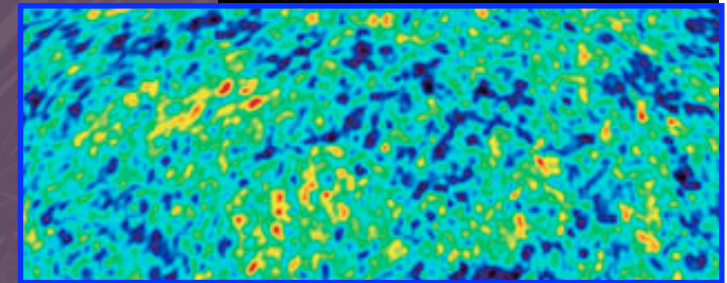
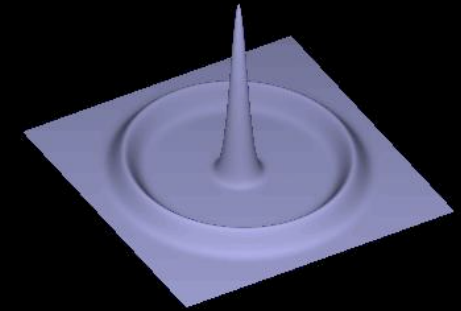
Absorption spectra towards close quasar pairs can probe the pressure-support (Jeans) scale of the diffuse IGM, and maybe cosmological parameters as well.



M. Peeples et al., in prep.

Baryon Acoustic Oscillations

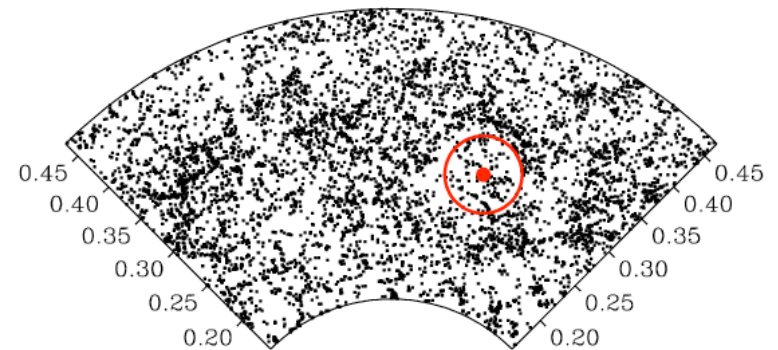
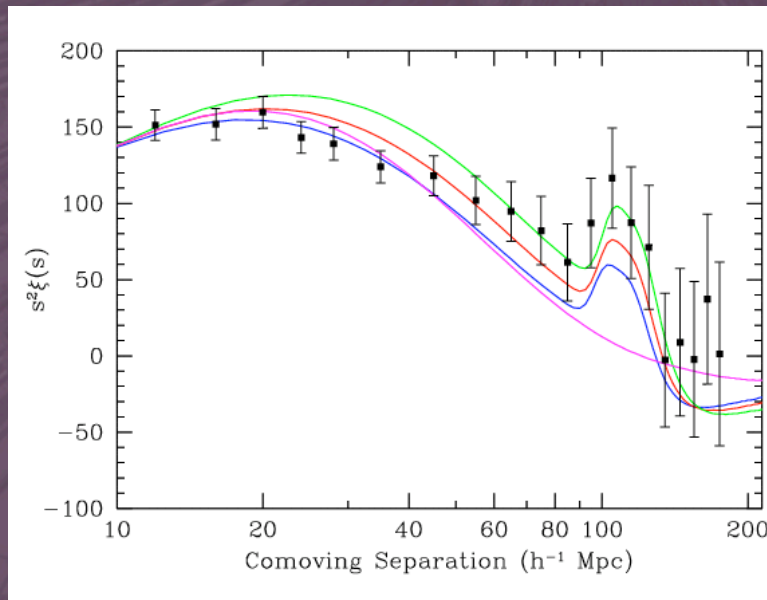
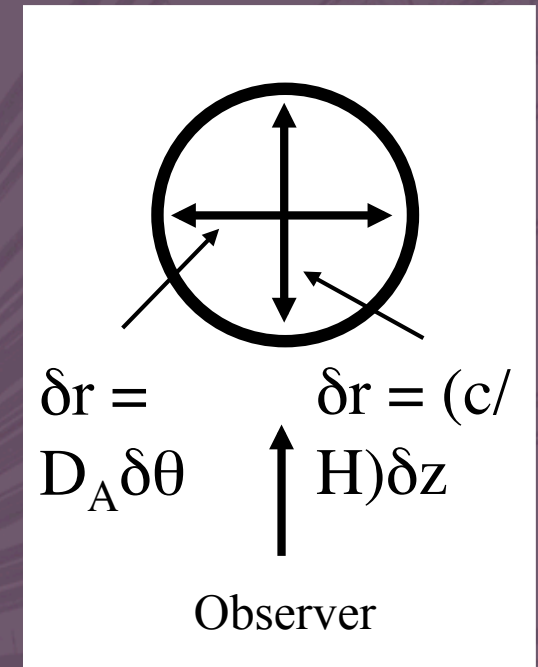
- Sound waves travel in hot early universe. Imprint characteristic scale in the CMB and (more weakly) on the galaxy distribution.
- Detected by SDSS in clustering of luminous red galaxies.
- Physical scale determined by CMB constraints plus straightforward physics.



Eisenstein et al. 2005

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- Provides standard ruler 150 Mpc long for measuring $D_A(z)$ and $H(z)$.

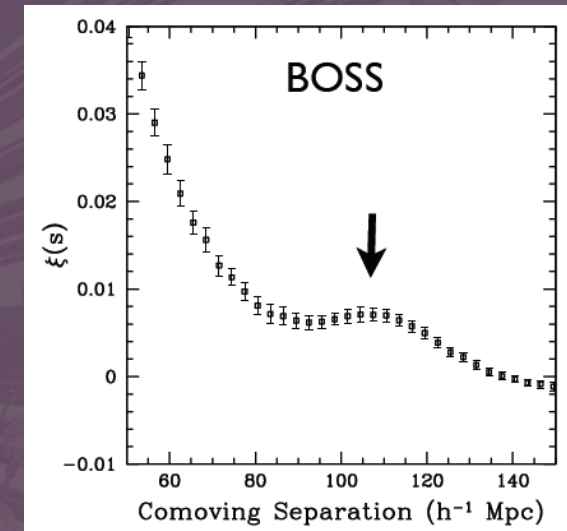


Eisenstein et al. 2005

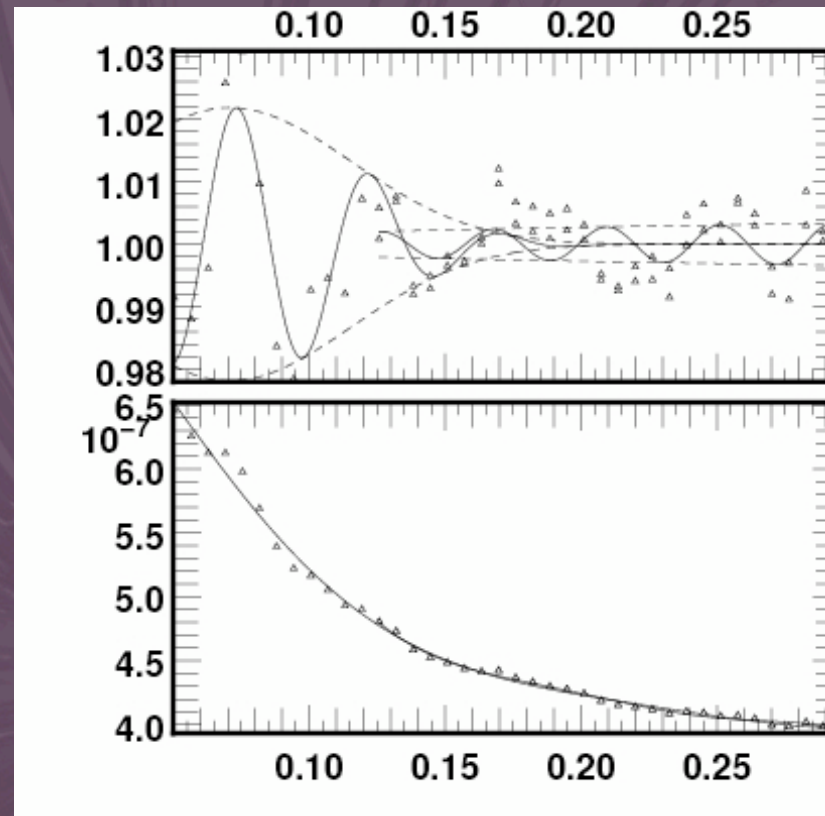
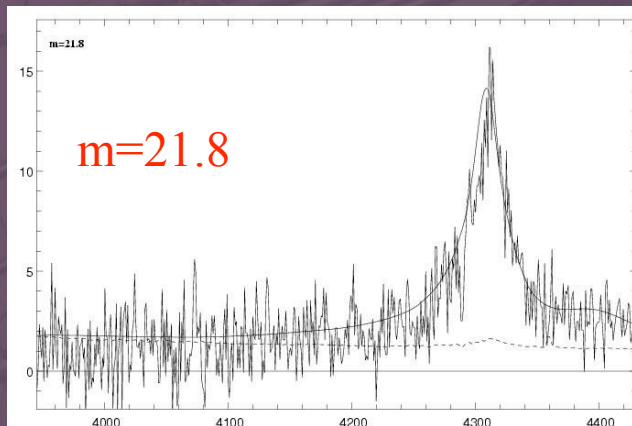
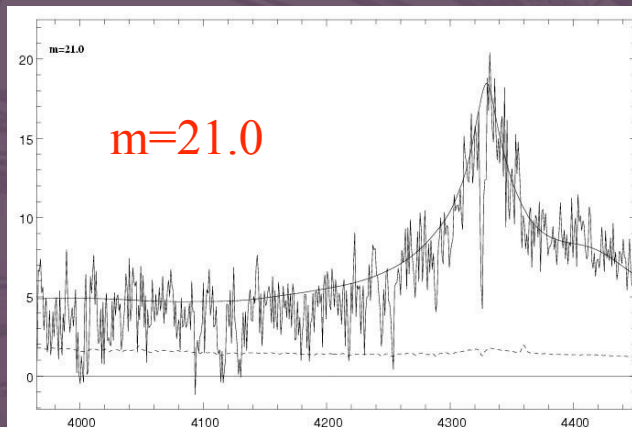
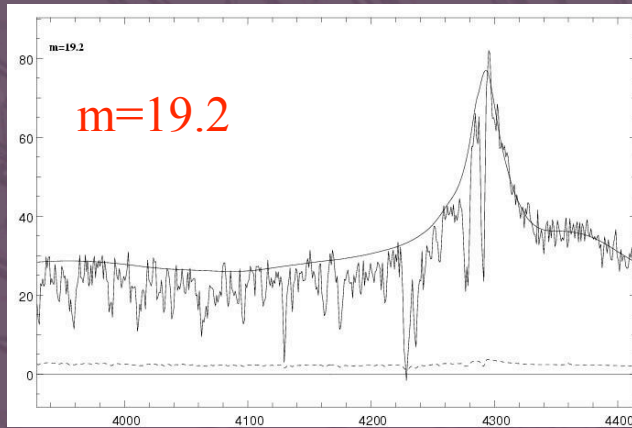
The Baryon Oscillation Spectroscopic Survey

Turn BAO into a precision tool for studying cosmic acceleration.

- Redshift survey of 1.5 million LRGs to $z=0.7$.
- Ly α forest spectra of 160,000 QSOs at redshifts
 - $2.1 < z < 3.0$. Limiting magnitude $g \approx 22$.
 - $\Sigma = 20 \text{ deg}^{-2} \Rightarrow 13.4 \text{ arc-min separation} = 16.4 \text{ h}^{-1} \text{ comoving Mpc at } z=2.4$.



- Forecast D_A precision: 1.0% (LRGs), 1.5% (Ly α F)
- Forecast $H(z)$ precision: 1.8% (LRGs), 1.5% (Ly α F)
- Uses SDSS spectrographs upgraded with higher throughput optics, more sensitive CCDs, 1000 2-arcsec fibers
- David Schlegel (PI), Martin White (SS), Natalie Roe (IS)

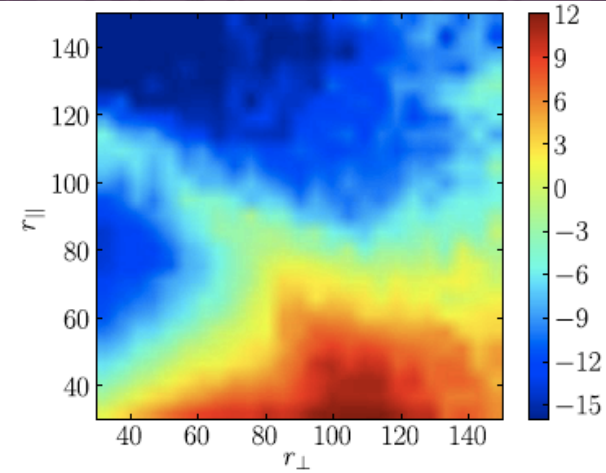
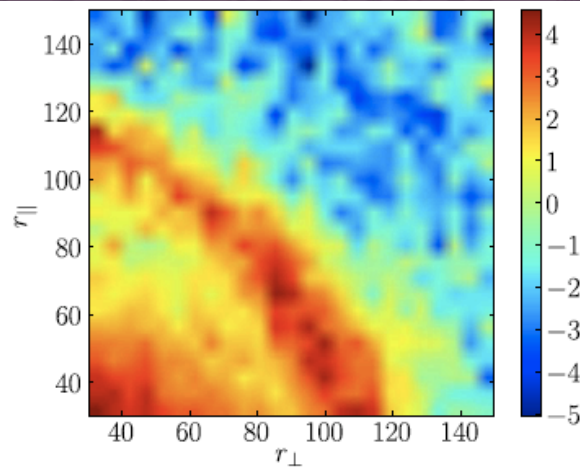


E. Rollinde, J.-M. Le Goff,
C. Pichon, P. Petitjean

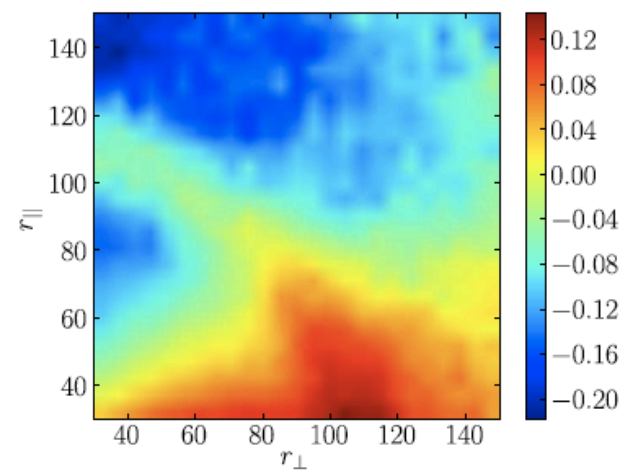
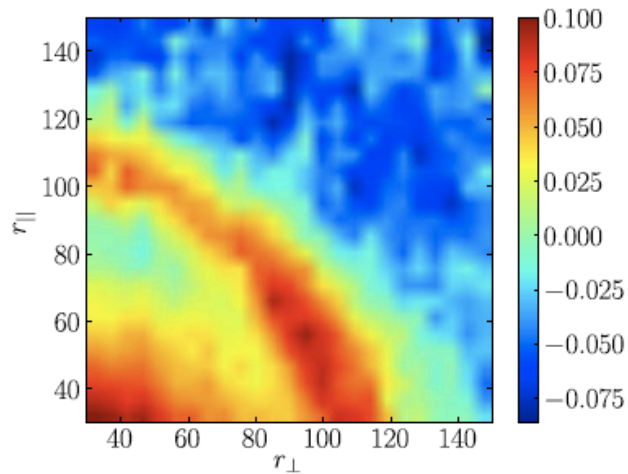
Real Space

Redshift Space

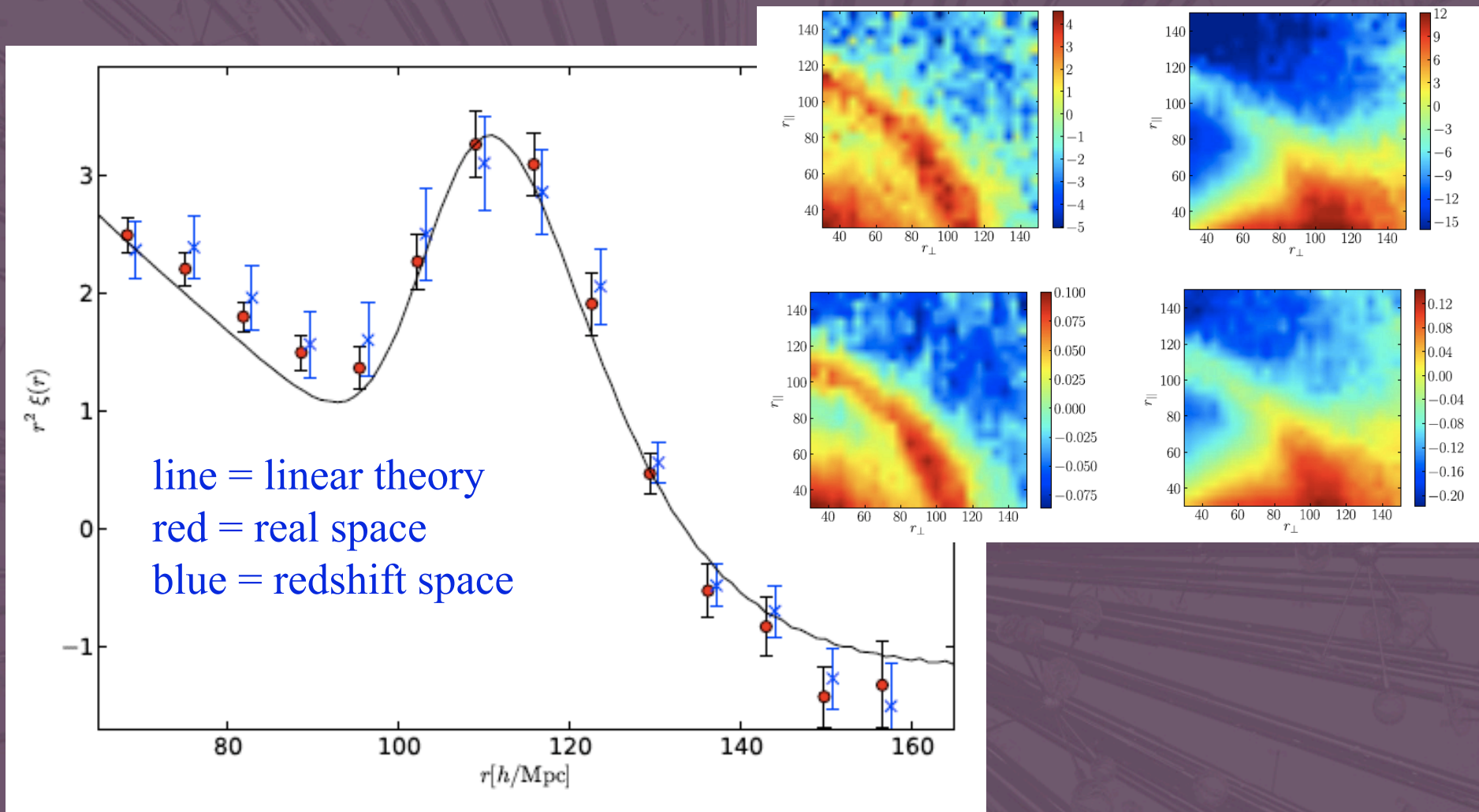
Mass



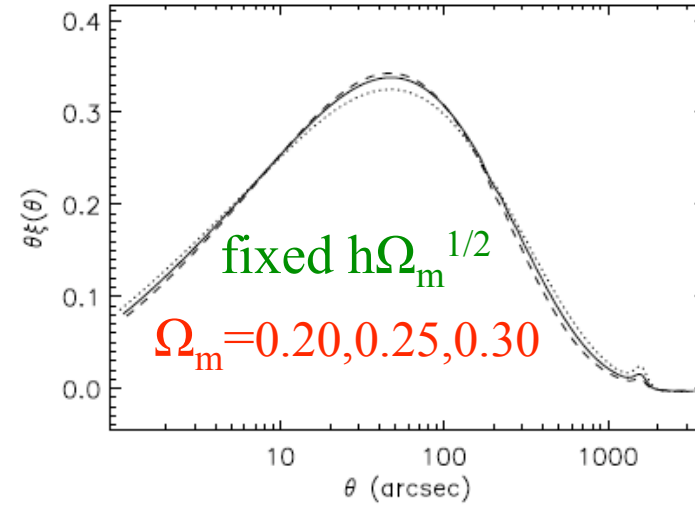
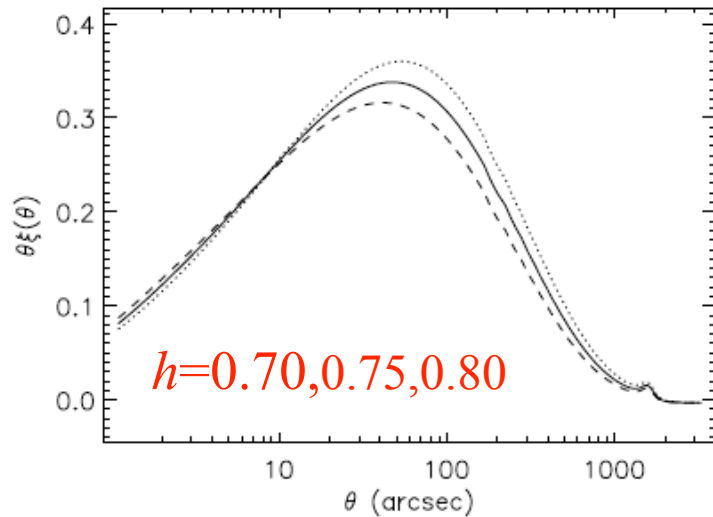
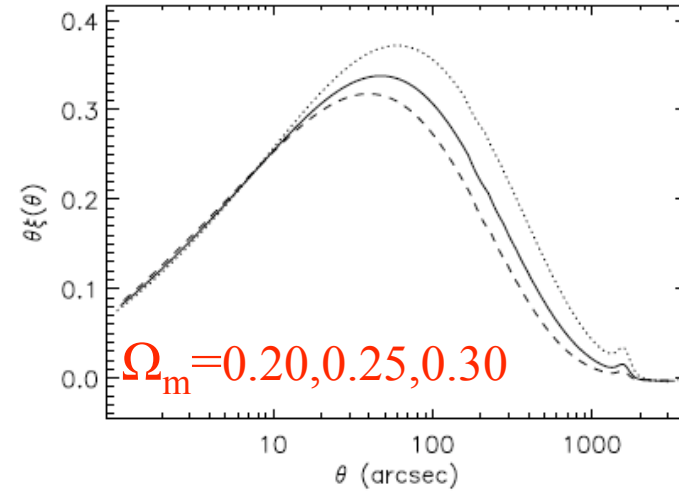
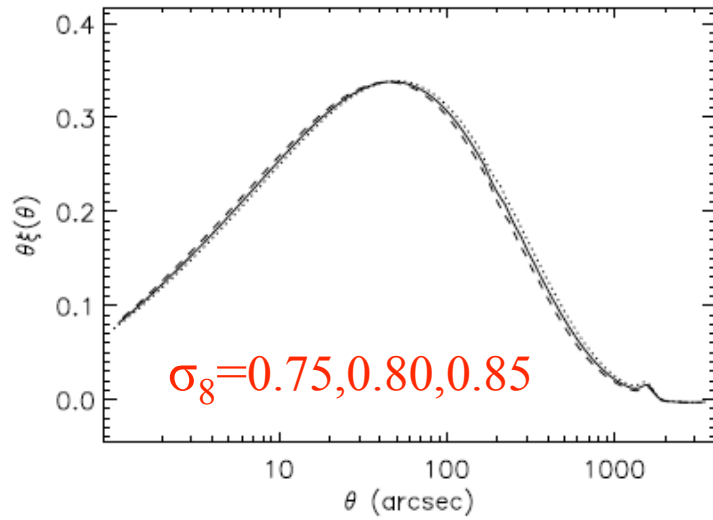
$\text{Ly}\alpha$ forest flux



Slosar, Ho, White, Louis 2009, arXiv:0906.2414
180,000 skewers through a $1.5 h^{-1}$ Gpc box



Slosar, Ho, White, Louis 2009



E. Rozo: Sensitivity of intermediate scales to cosmological and IGM parameters, based on log-normal analytic model.

Issues we are investigating:

- How to select 20 high- z QSOs per deg^2 from available data
- QSO continuum determination.
- Metal-line contamination.
- Spatial variations of UV background.
- Shape and spatial variations of IGM equation-of-state.
- Instrument and data pipeline artifacts.
- Optimal extraction of signal from data.

So far, systematics for BAO measurement look small compared to expected statistical errors.

Summary: The Ly α Forest in 3-D

- Close quasar pairs can measure pressure smoothing scale of IGM
- SDSS-III: BAO cosmology, Galactic structure, exoplanets
One year down, five years to go.
- BOSS: 1.5 million LRGs to $z=0.7$; 160,000 QSOs $2.1 < z < 3.0$,
typical QSO sightline separation of $15\text{-}20 h^{-1}$ Mpc comoving
- BAO signal clearly present in large volume Ly α F simulations
- BOSS makes the Ly α F a 3-D phenomenon, many applications

