

# Strong galaxy-galaxy lenses in COSMOS

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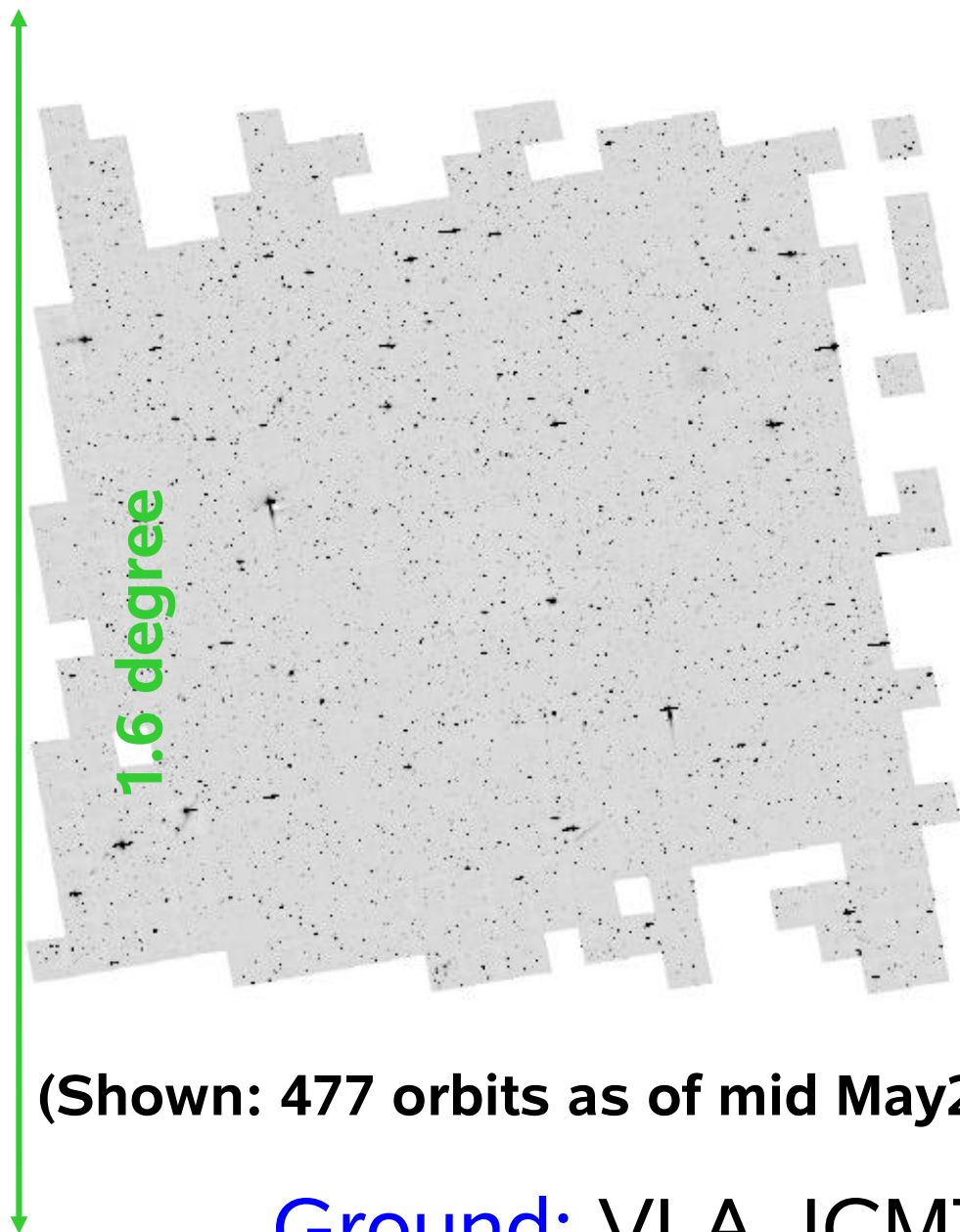
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# 1) The COSMOS survey



## HUBBLE Treasury Data

PI: N. Scoville 590 orbits

ACS/F814W (~ 1.6 sq. deg)

+ NICMOS-3/F160W

## SUPRIME@ Subaru

PI: Y. Taniguchi

1.5x1.5 deg

0.7" ~seeing

BVgriz bands

## MEGACAM@ CFHT

$u^*, i^*$

Seeing < 1.0"

Ground: VLA, JCMT, IRAM

Space: XMM, Chandra, Spitzer

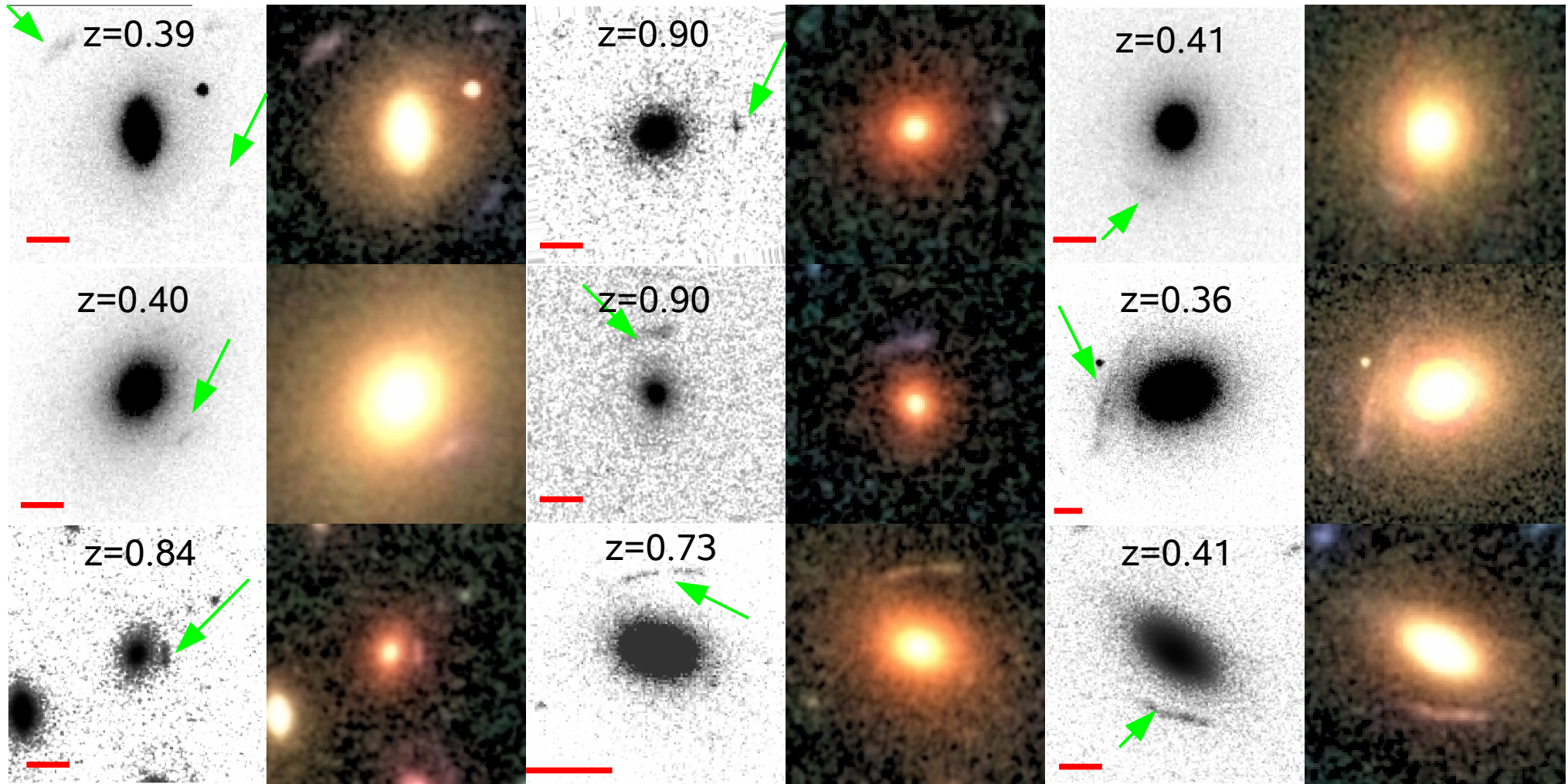
## 2) The search for strong galaxy-galaxy lenses

→ 278526 galaxies with z-phot (*Mobasher et al. 2007*)

Visual inspection of potential lensing galaxies in ACS I-band stamp images of 10"x10":

- **photometric redshift:**  $0.2 \leq z \leq 1.0$
  - **luminosity:**  $M_V < -20$ . mag
  - **galaxy type:** early-type as fitted from the SED when computing the photo-z
- 9452 galaxies in **parent catalog**
- 337 lens candidates after ACS inspection
- 67 lens candidates after color criterion and galaxy light profile fit and subtraction (47 single arcs and 20 multiple images)

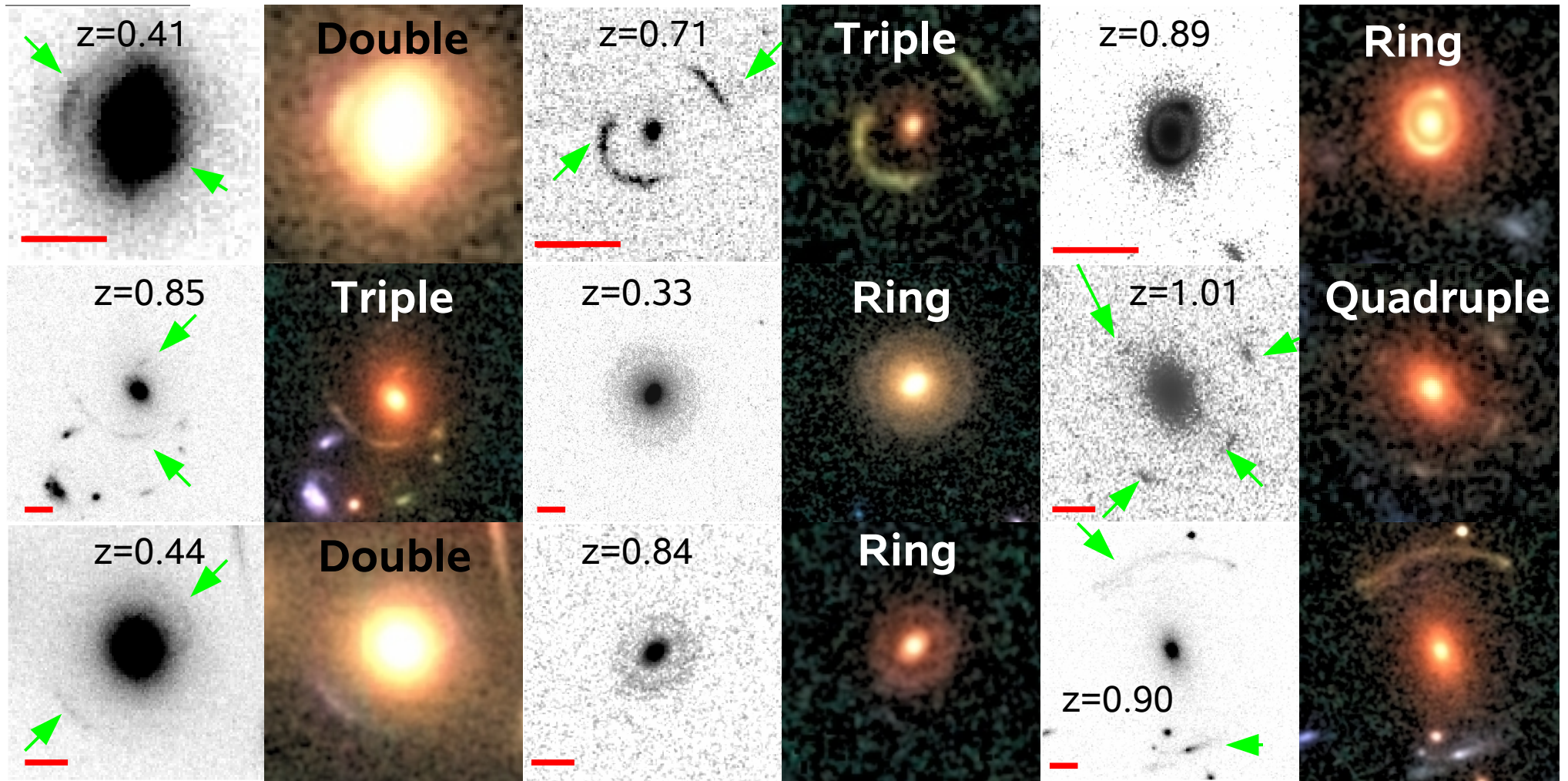
# 3) Some candidates: single arc systems



— 1 arc-second

ACS-F814w / Subaru-B<sub>r</sub><sup>+</sup>,z<sup>+</sup> images sharpened with the ACS images

# Some candidates: best systems

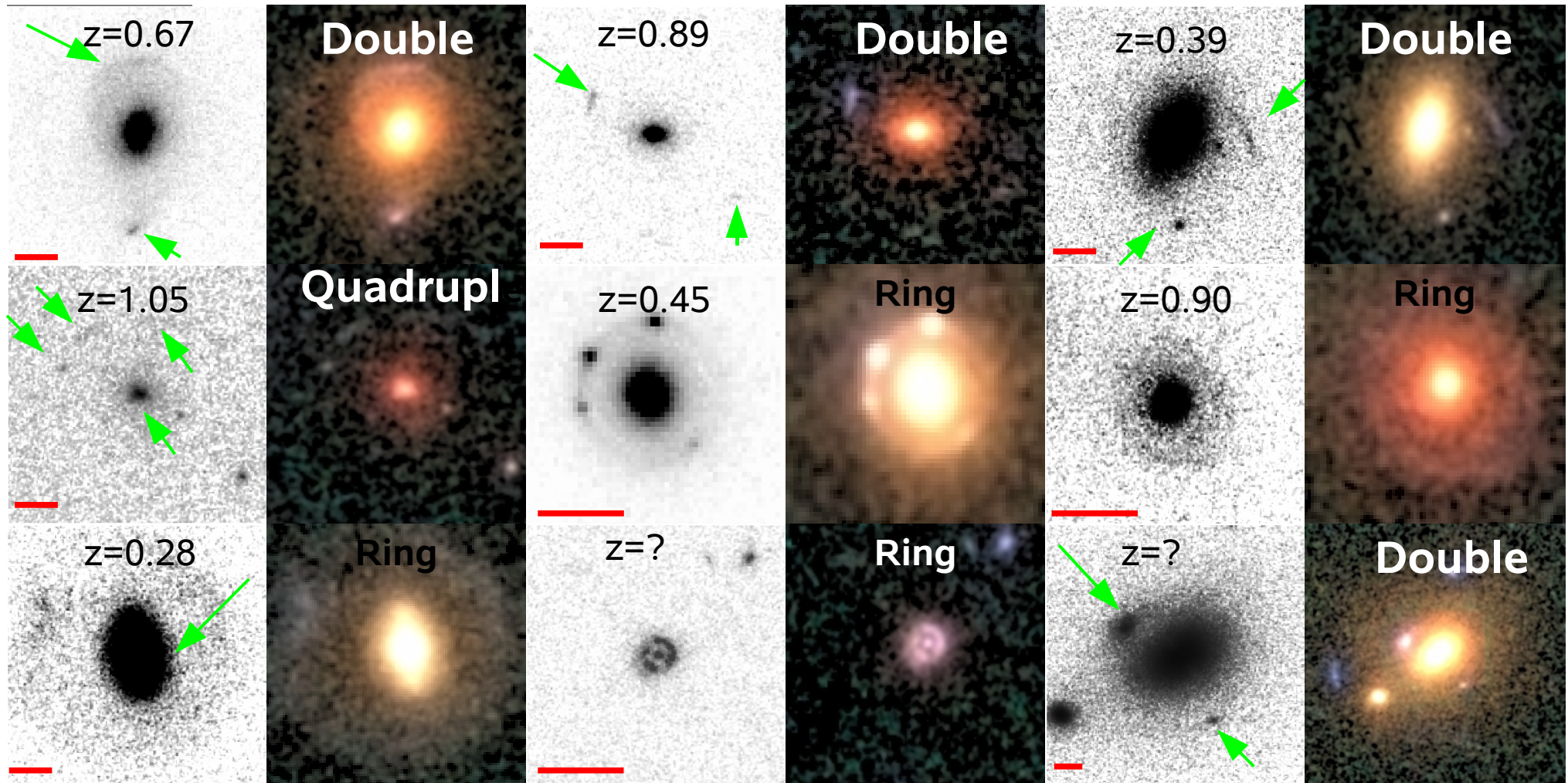


— 1 arc-second

2xDoubles  
or 2xTriples

ACS-F814w / Subaru-B<sub>r</sub><sup>+</sup>,z<sup>+</sup> images sharpened with the ACS images

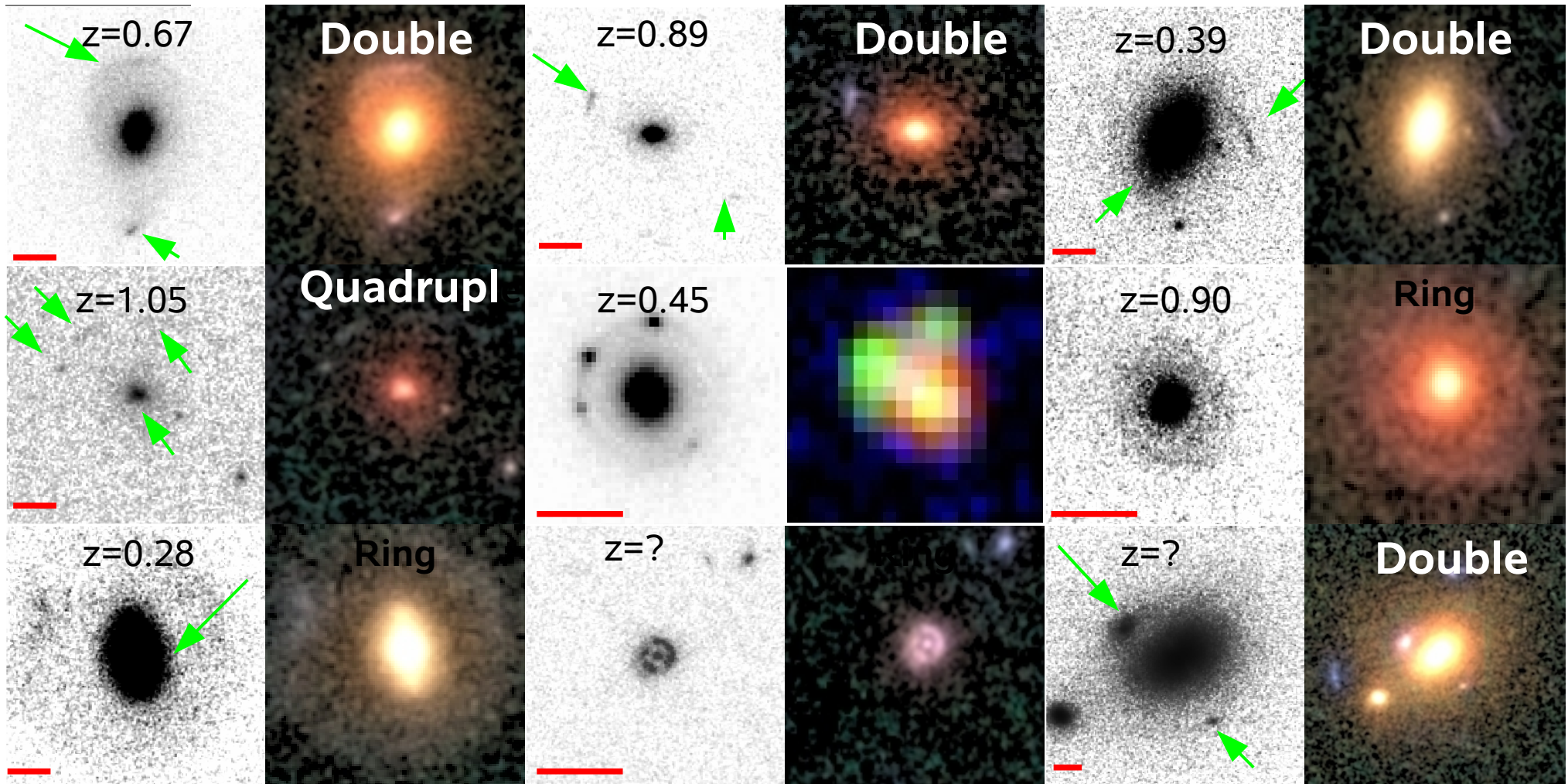
# best candidates 2



— 1 arc-second

ACS-F814w / Subaru-B<sub>j</sub>,r<sub>+</sub>,z<sub>+</sub> images sharpened with the ACS images

# best candidates 2



— 1 arc-second

ACS-F814w / Subaru-B<sup>j</sup>,r<sup>+</sup>,z<sup>+</sup> images sharpened with the ACS images

# 4) Lens modelling

- **Lens model SIE+ $\gamma$**

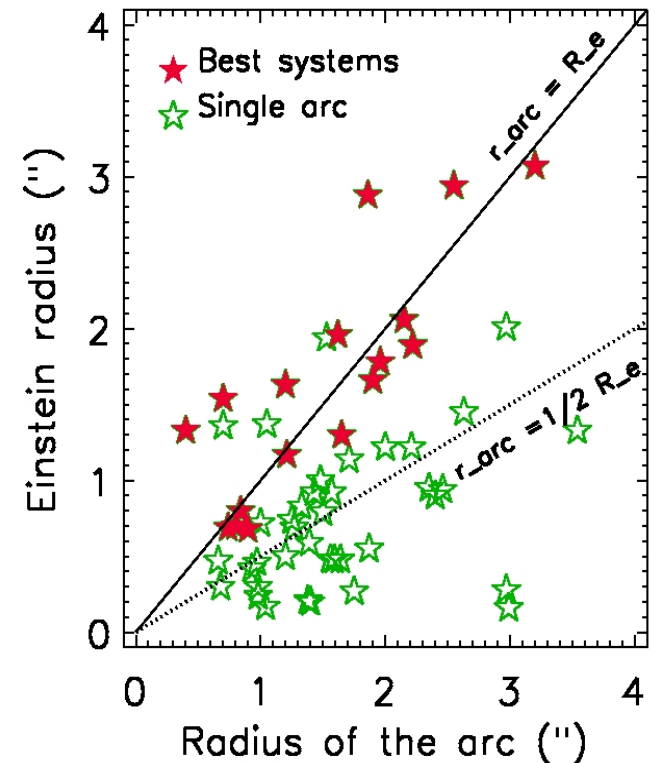
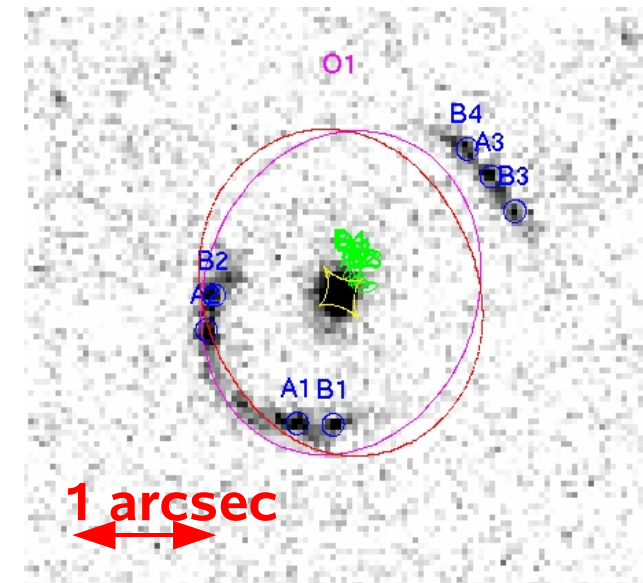
(LensTool, *Kneib et al 1993, Jullo et al 2007*)

assuming:  $z_S = 2 \cdot z_L$

- **Measurement of the Einstein radius**, or equivalent mass within the Einstein radius

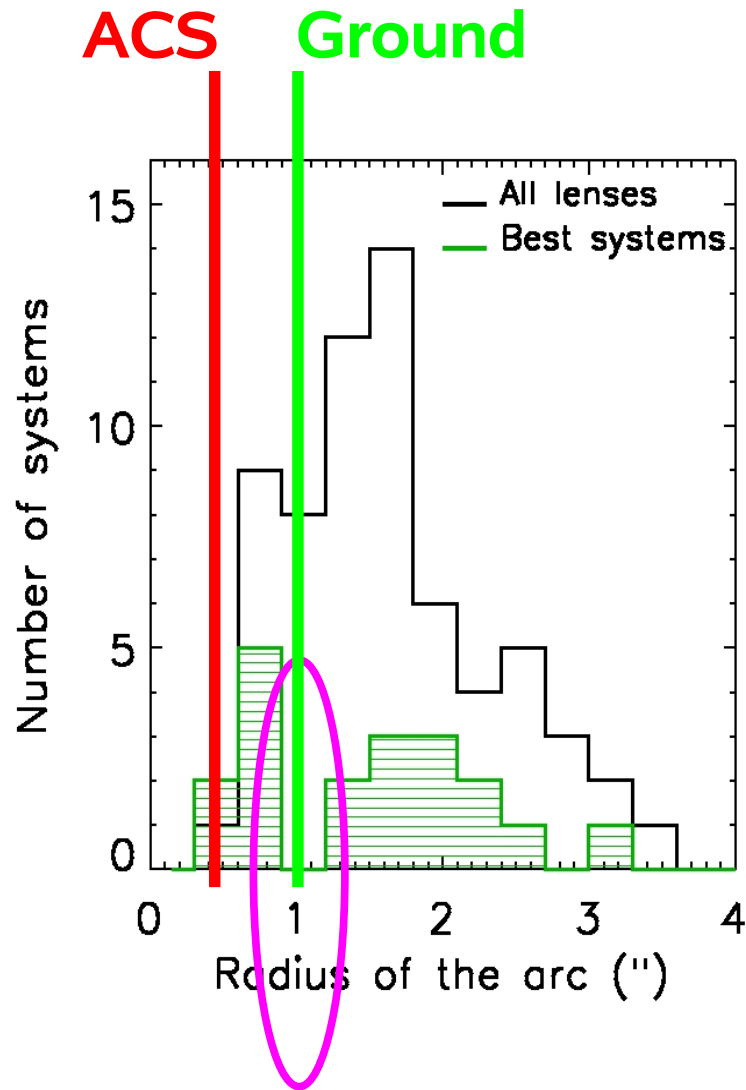
➔ **Confirmation of the single/multiple images classification**

➔ Measuring redshifts is key for proper mass models (zCOSMOS and FORS1-MOS data)





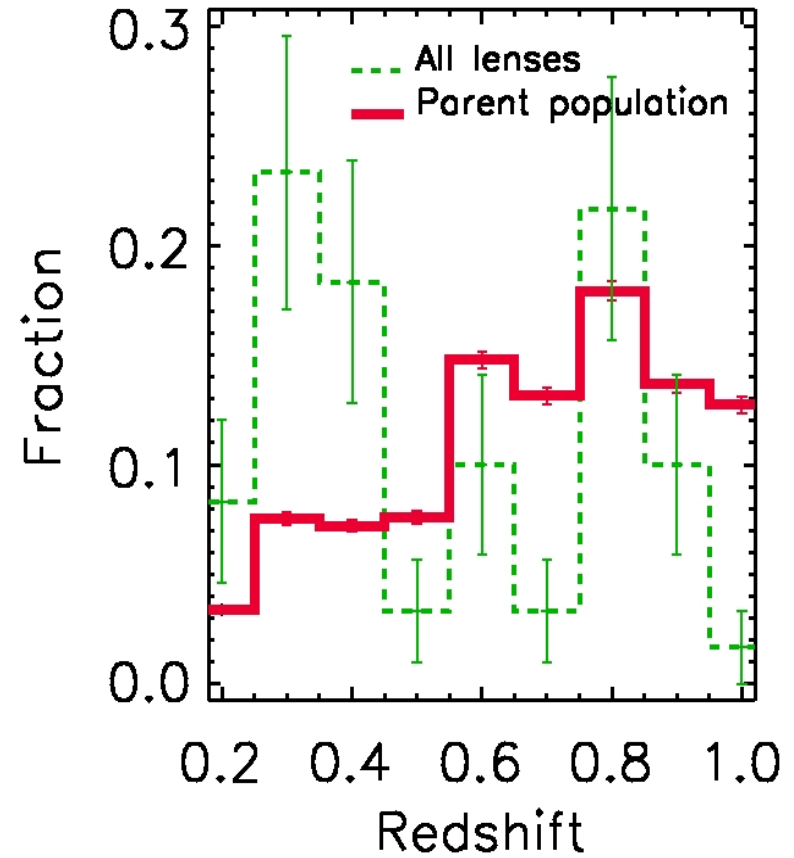
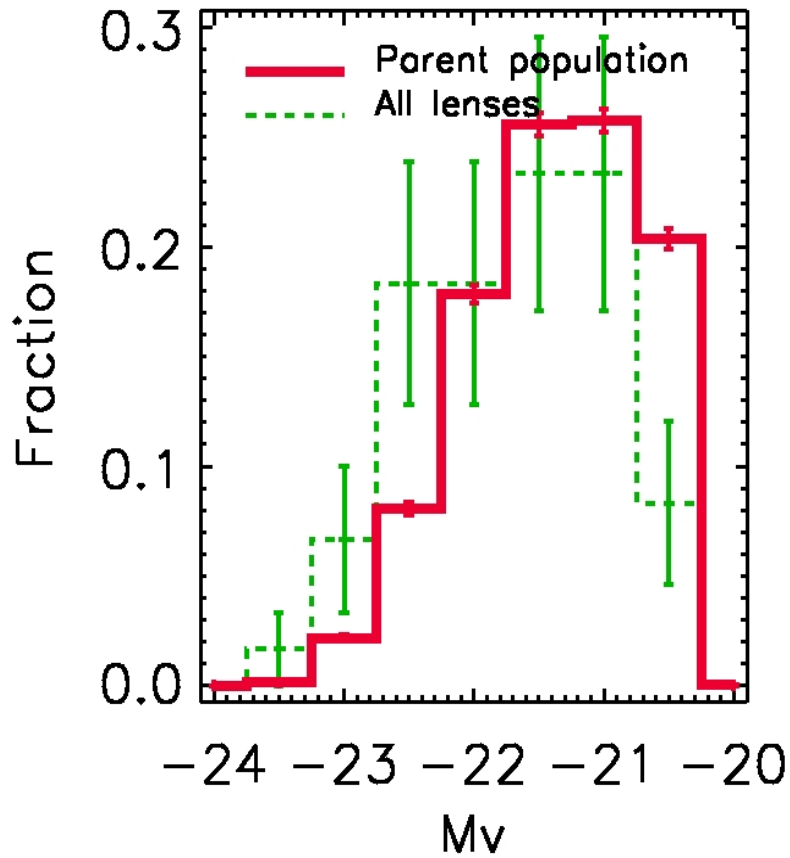
# Some properties of the sample



Lack of multiple image systems with  $r_{\text{arc}} \sim 1''$ !

- transition between ACS high resolution images to ground color images
- some of the single arcs systems are multiple?
- incompleteness of the lens sample

# Some properties of the sample 2



lensing galaxies  
brighter than the parent  
population

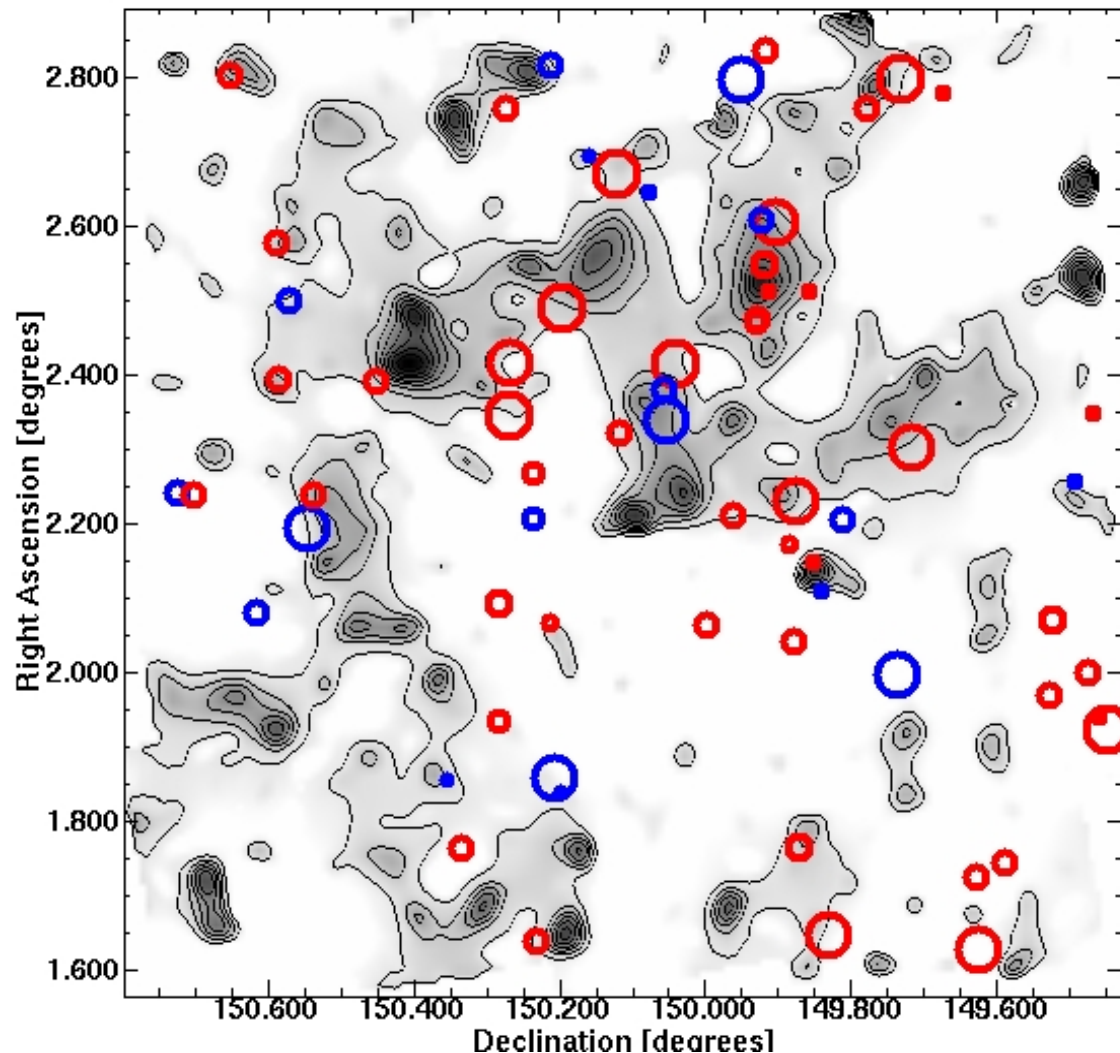
**2 peaks?**

Kosmogorov-Smirnov test:  
92% prob the same  
distribution

(poissonian noise)

# 5) Strong lenses in the mass map

Blue: best lenses Red: single arcs



Size depends on arc radius

small:  $r_{\text{arc}} < 1''$

medium:  $1'' \leq r_{\text{arc}} < 2''$

large:  $2'' \leq r_{\text{arc}} < 5''$

We define three regions:

« empty »:

$\kappa < 0.4\% \Leftrightarrow 70\%$  field

« filament »:

$0.4 \leq \kappa < 2\% \Leftrightarrow 26\%$  field

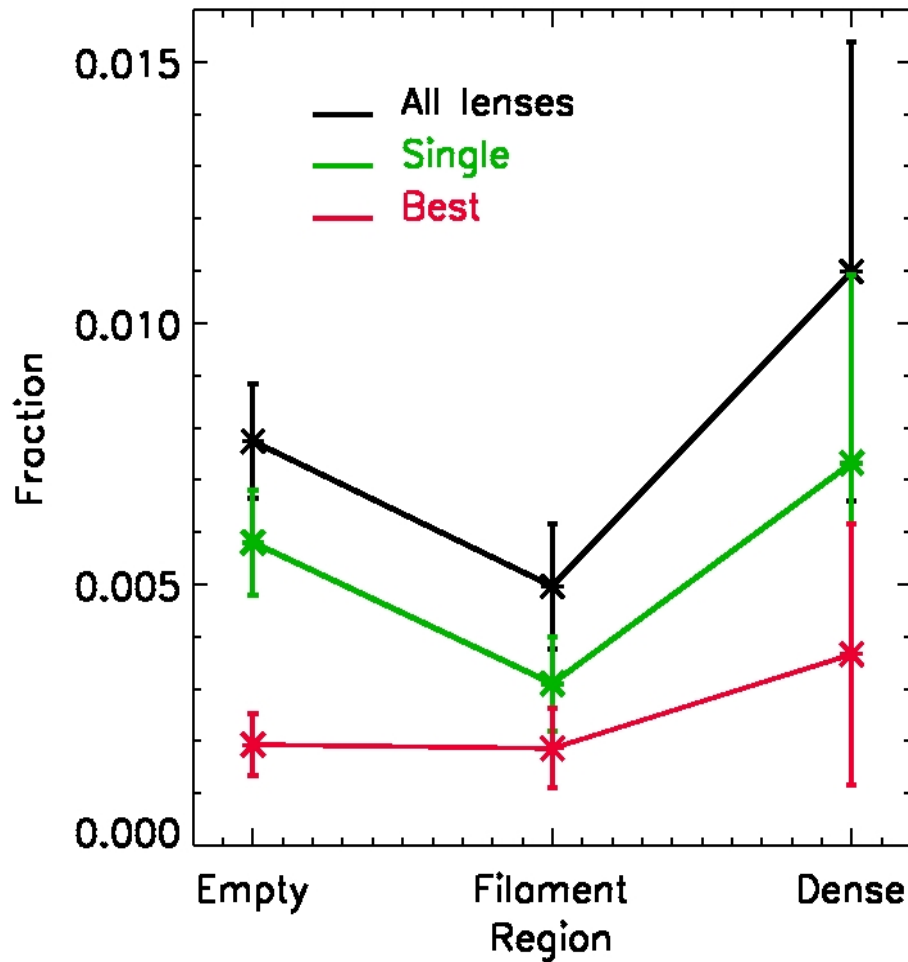
« dense »:

$\kappa > 2\% \Leftrightarrow 4\%$  field

(peak at  $\sim 4.5\%$ )

Convergence map: *Massey et al. 2007*

## A) The strong lenses versus the parent population:



### In complex environment:

33 $\pm$ 1 % of all candidates

41 $\pm$ 2 % of the best candidates

30 $\pm$ 1 % of the single arc

### Bright elliptical galaxies that are strong lenses:

• “complex”: 1.5 $\pm$ 0.5 %

• “empty”: 0.7 $\pm$ 0.1 %

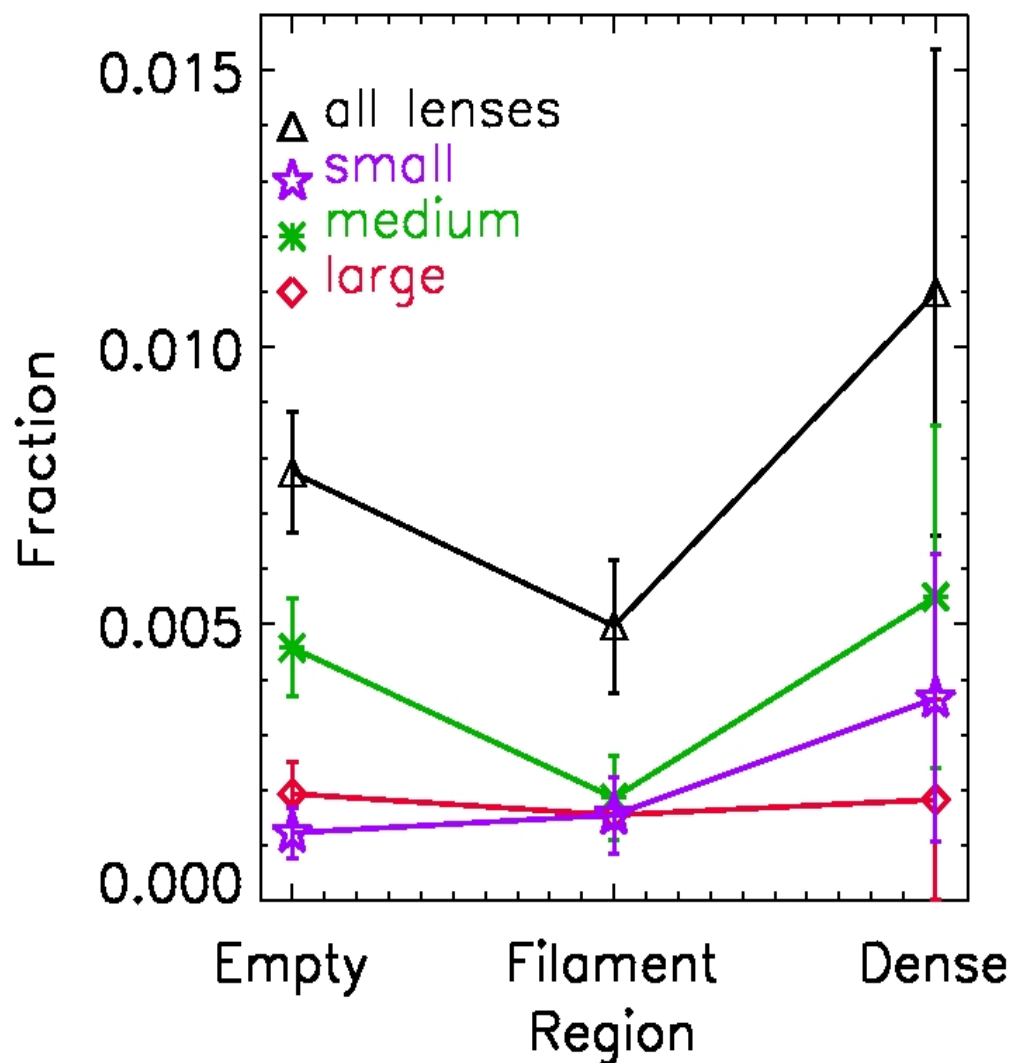
•  $0.2 \leq z < 0.5$ : 1.7 $\pm$ 0.3 %

•  $0.5 \leq z \leq 1.0$ : 0.3 $\pm$ 0.1 %



filament+dense=**complex** environment

## B) Distribution of the strong lenses according to their arc radius:



small:

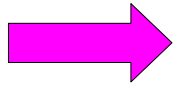
$r_{\text{arc}} < 1''$  (14 lenses)

medium:

$1'' \leq r_{\text{arc}} < 2''$  (35 lenses)

large:

$r_{\text{arc}} \geq 2''$  (17 lenses)



20+/-20 % of our best candidates  
(or 6+/-6 % of all lenses)  
with “large” arc radius ( $2'' \leq r_{\text{arc}} < 5''$ ) are in  
a “dense” environment ( $\kappa > 0.016$ )

Simulations by Oguri, Keeton & Dalal. 2005:

15% to 60% of lenses with  $2 \leq r_{\text{arc}} < 5$  should be in  
environment with  $\kappa \geq 0.1$

us: 0% of the candidates 8-(

...?!?

- Catalog of strong lens candidates incomplete
- Small galaxy groups favorable to this regime of arc radius
- Efficiency of double layer lenses underestimated in simu ?  
=> Numerical simulations by *Wambsganss, Bode & Ostriker 2005*,  
predicts 20 to 40% of lenses to be double layers

# Preliminary results on Galaxy-Galaxy weak lensing in COSMOS by Alexie Léauthaud

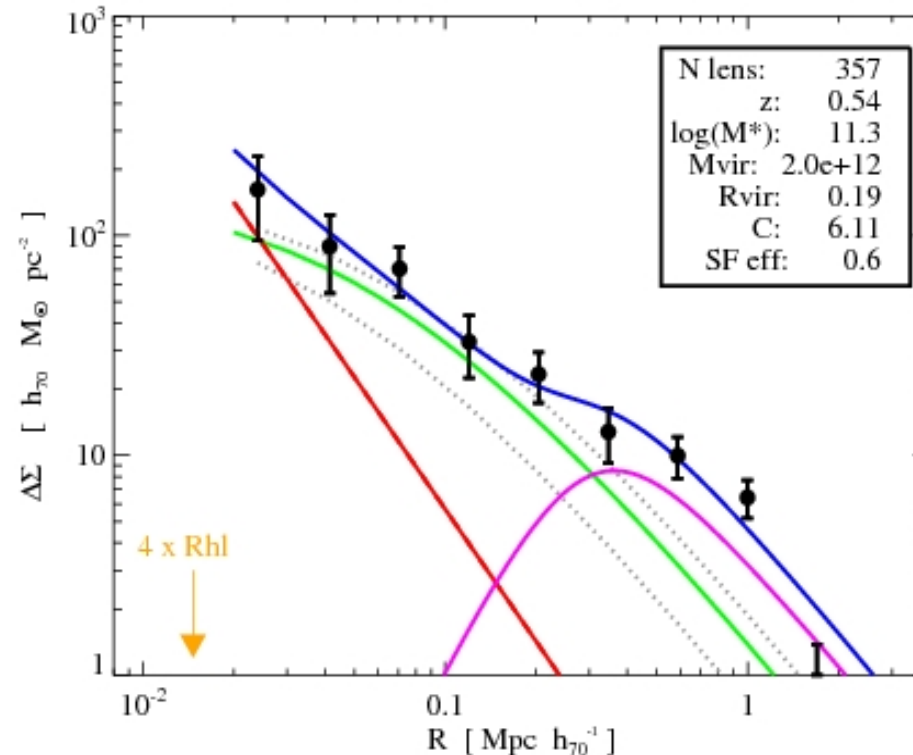
$$\Delta\Sigma = \Delta\Sigma_b + \Delta\Sigma_{\text{NFW}} + \alpha \cdot \Sigma_{\text{NC}}$$

$\Sigma_b$ : Baryonic contribution determined by the stellar mass

$\Sigma_{\text{NFW}}$ : NFW profile assumed for dark matter halos.

$\alpha$ : Fraction of galaxies in sub-halos.

$\Sigma_{\text{NC}}$ : Off centered 'group' contribution.



# Summary

## On the sample

- new sample of strong lens candidates
- lower limit in occurrence of strong lenses in space surveys
- valuable sample to test robustness of semi-automatic detection software (Haggles, *P. Marshall et al. 2007*)

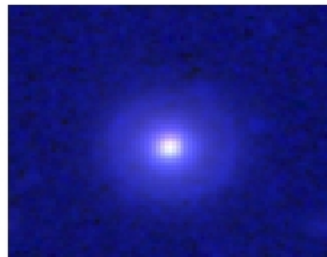
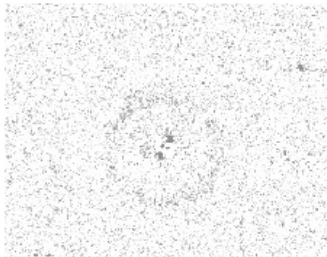
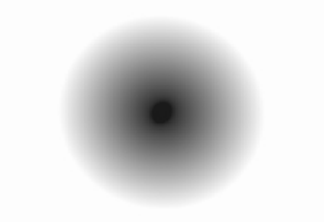
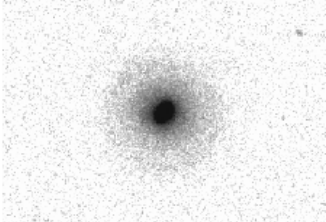
## On the comparison with the mass map and with the parent population

- 30% to 40% of the lensing galaxy candidates are in complex environment
- 1.5% of the bright elliptical galaxies are strong lensing galaxy candidates if they are in a complex environment
- lens candidates are  $\sim 0.5$  mag brighter than the parent population of galaxies
- may be two peaks in the redshift distribution
- the arc radius size is not correlated with the large scale structures in the regime of arc radius probed

(*Faure et al. 2007, ApJ - Faure et al. 2007, Letter to ApJ*)



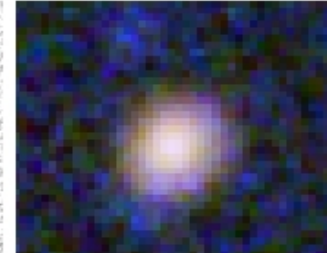
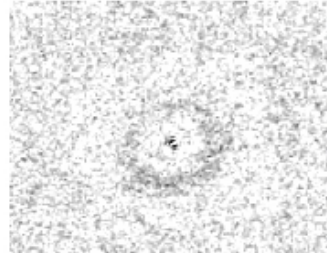
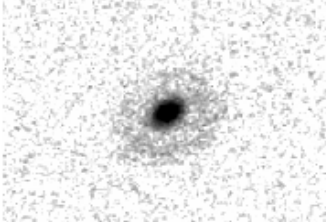
COSMOS0049+5128



$$r_{\text{arc}} = 2.22''$$

6"x6"

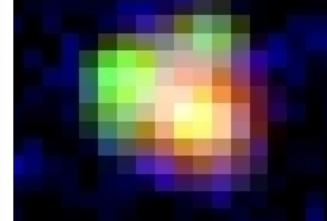
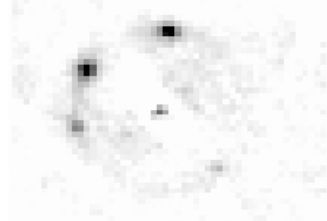
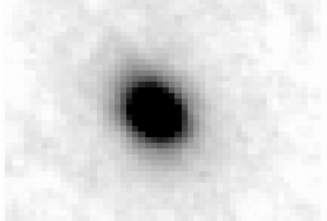
COSMOS0124+5121



$$r_{\text{arc}} = 0.84''$$

3"x3"

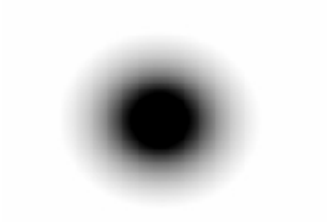
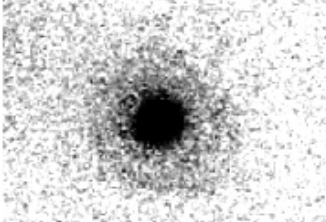
COSMOS5921+0638



$$r_{\text{arc}} = 0.80''$$

3"x3"

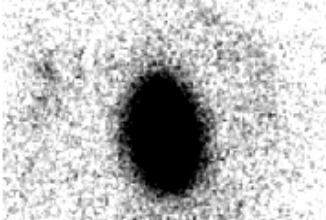
COSMOS5941+3628



$$r_{\text{arc}} = 1.21''$$

3"x3"

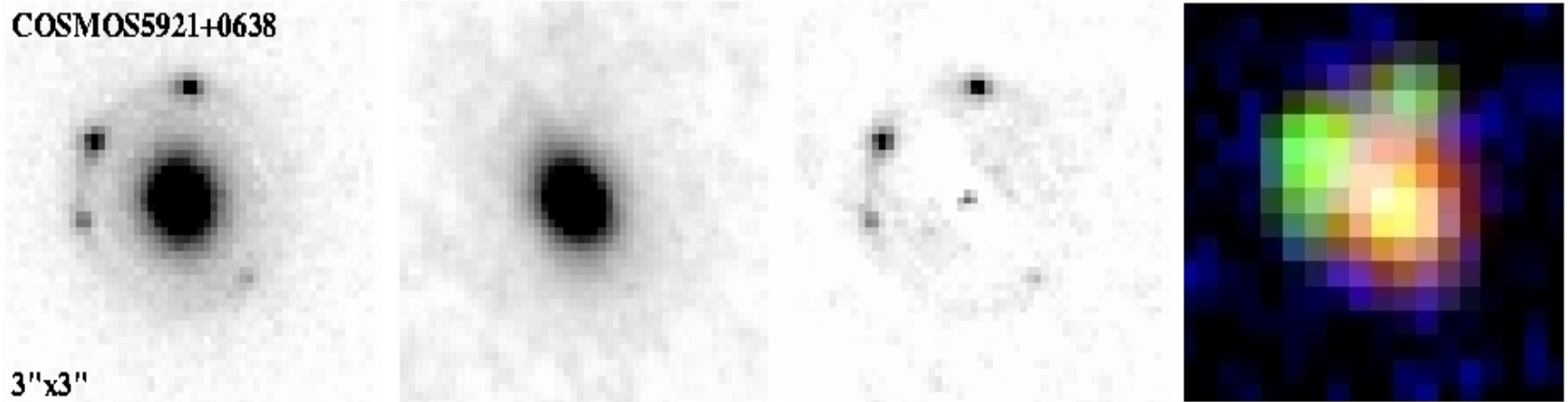
COSMOS5947+4752



$$r_{\text{arc}} = 2.55''$$

3"x3"

COSMOS5921+0638



Gim2D (*Simard 1998, Marleau et Simard 1998*)  
sersic bulge + exponential disc