

CMB polarisation from clusters and filaments

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Secondary CMB Polarisation

- Lensing effect

Mode mixing

due to E converted to B

e.g.:Zaldarriaga, Seljak, 1998

Benabed et al. 2001

- Faraday rotataion

phase shift

→ mode mixing

e.g.:Takada, Ohno, Sugiyama 2001

- Scattering off ionised gas

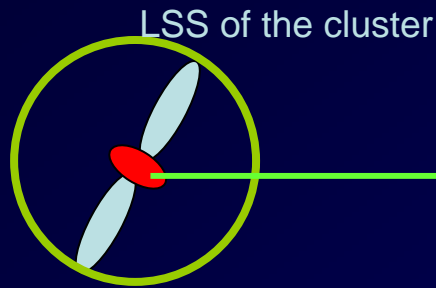
power generated

Sunyaev, Zel'dovich 1980

e.g.:Sazonov, Sunyaev 1999

Sources of Quadrupole

1. Primordial CMB Quadrupole

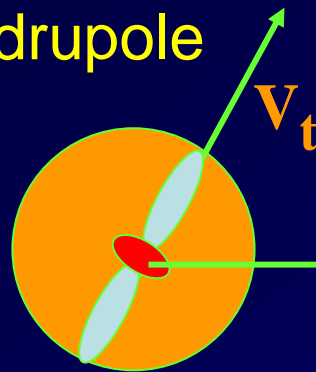


$$P \propto \tau Q_{\text{cmb}}$$

τ : Optical depth

Obs 

2. Kinetic Quadrupole

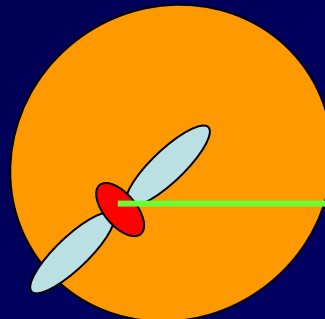


$$P \propto \tau v_t^2$$

v_t : transverse velocity

Obs 

3. Double scattering



$$P \propto \tau^2 v_t$$
$$P \propto \tau^2 T_e$$

T_e : cluster temperature

Obs 

Modulated Quadrupole

Density fluctuation

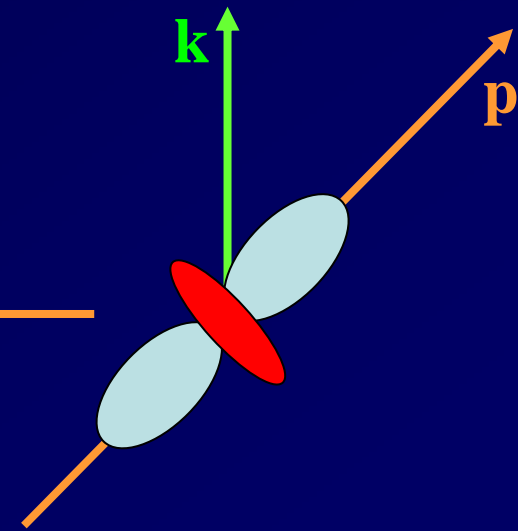
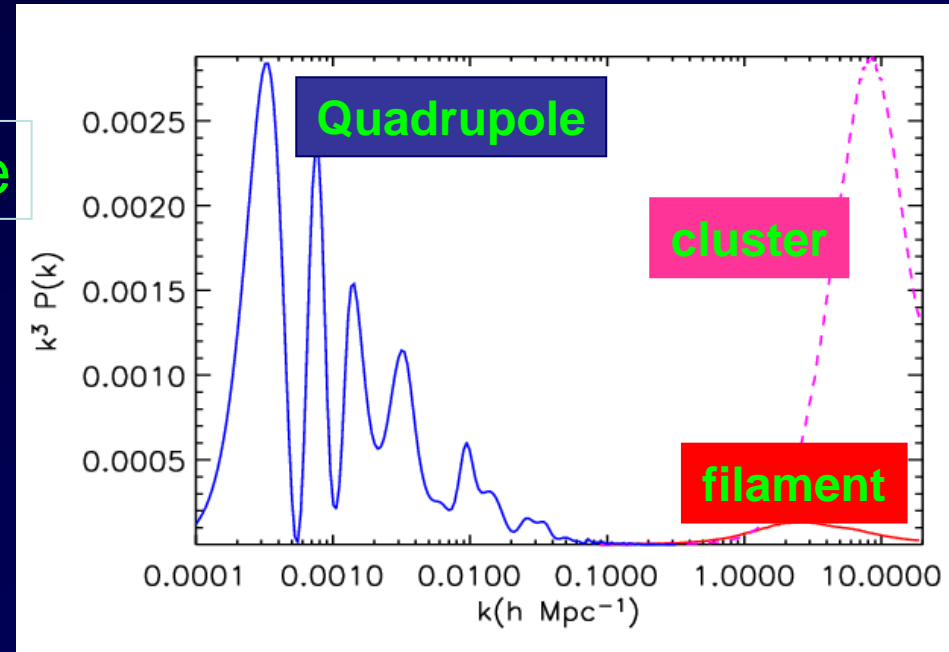
Quadrupole

$S^m(\mathbf{k}, \eta)$:

$$= \int d^3\mathbf{p} \delta_e(\mathbf{k}-\mathbf{p}, \eta) \square_{T2}^m(\mathbf{p}, \eta)$$

$$\cong \delta_e(\mathbf{k}, \eta) \int d^3\mathbf{p} \square_{T2}^m(\mathbf{p}, \eta)$$

$$= \delta_e(\mathbf{k}, \eta) \int d^3\mathbf{p} \square_{T2}^0(\mathbf{p}, \eta) Y_2^m(\mathbf{p})$$



Formula

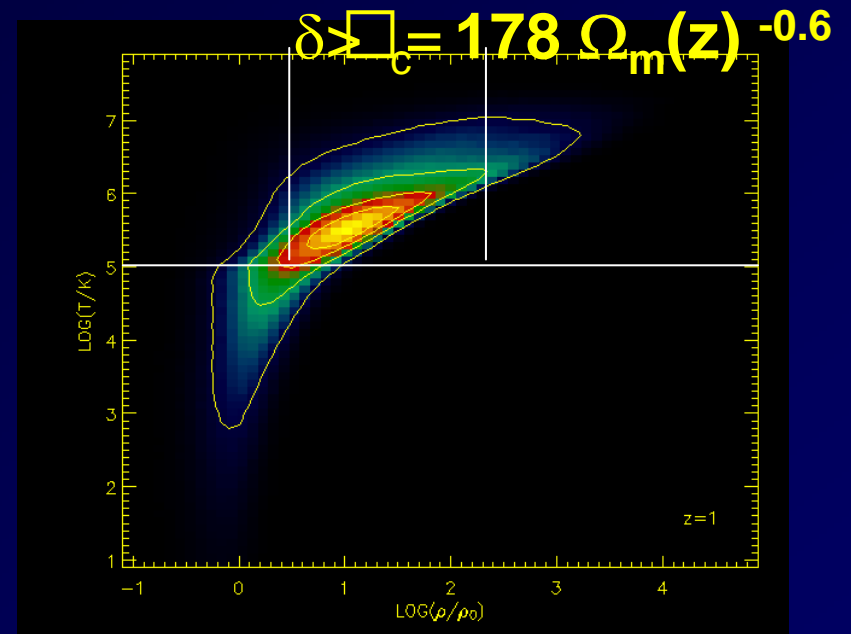
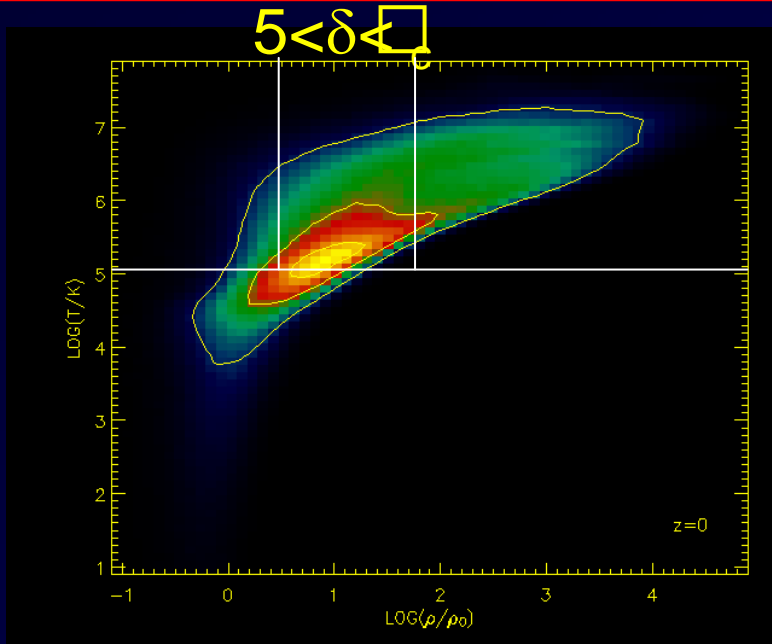
$$g(\eta) \equiv - e^{\tau(\eta_0) - \tau(\eta)} d\tau/d\eta$$

visibility function: possibility of last scattering at epoch η

$$C_{(E,B)\ell} \propto \ell^4 \sum_m \int k^2 dk \langle |\int d\eta g(\eta) S^m(k, \eta) T^m_{(E,B)\ell}(k, r)|^2 \rangle$$

m	$T^m_{E/}$	$T^m_{B/}$
0	$(-i)^l j_l(kr)/(kr)^2$	0
± 1	$\pm (-i)^l [(l+1)j_{l-1}(kr) - l j_l(kr)] / [(2l+1)kr - \sqrt{6}l(l+1)]$	$\pm (-i)^l [\sqrt{3}/2l(l+1)]^* j_l(kr)/(kr)^2$
± 2	$\pm (-i)^l \{ [(l+2)(l+1)/(2l-1) + l(l+1)/(2l+3) - (2l+1)(l-1)(l+2)/(2l-1)(2l+3)] j_l(kr) - (l+2)(l+1) j_{l-1}(kr)/kr - l(l-1) j_{l+1}(kr)/kr \} \sqrt{((l-2)!/6(l+2)!)/(2l+1)}$	$\pm (-i)^l \{ (l+2) j_{l-1}(kr) - (l-1) j_{l+1}(kr) \} \sqrt{((l-2)!/6(l+2)!)/(2l+1)}$

Gas Evolution



$$\Omega_m = 0.3$$

$$\Omega_b = 0.044$$

$$\Omega_\Lambda = 0.7$$

$$h = 0.71$$

$$L_{\text{box}} = 100 h^{-1} \text{Mpc}$$

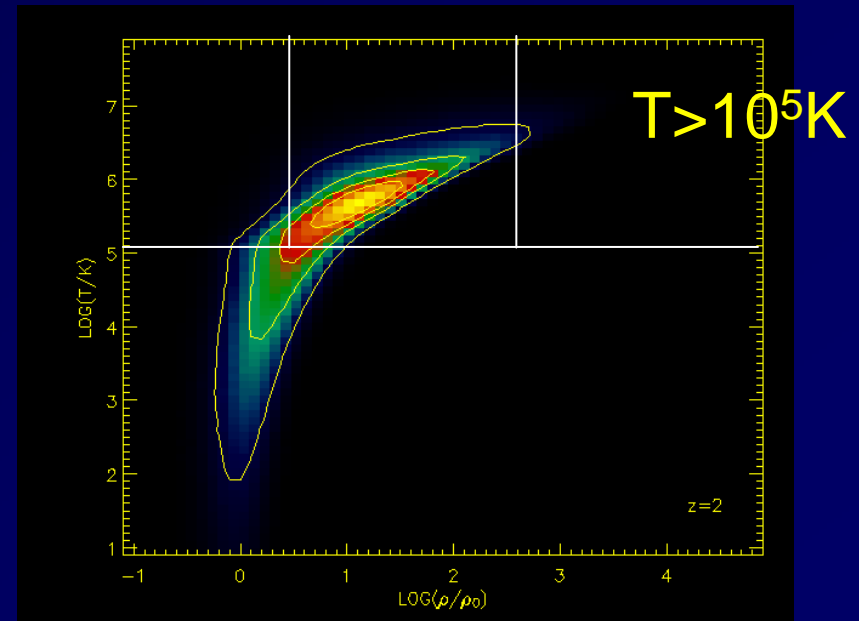
$$\sigma_8 = 0.9$$

Mass of particle: $10^{10} M_\odot/h$

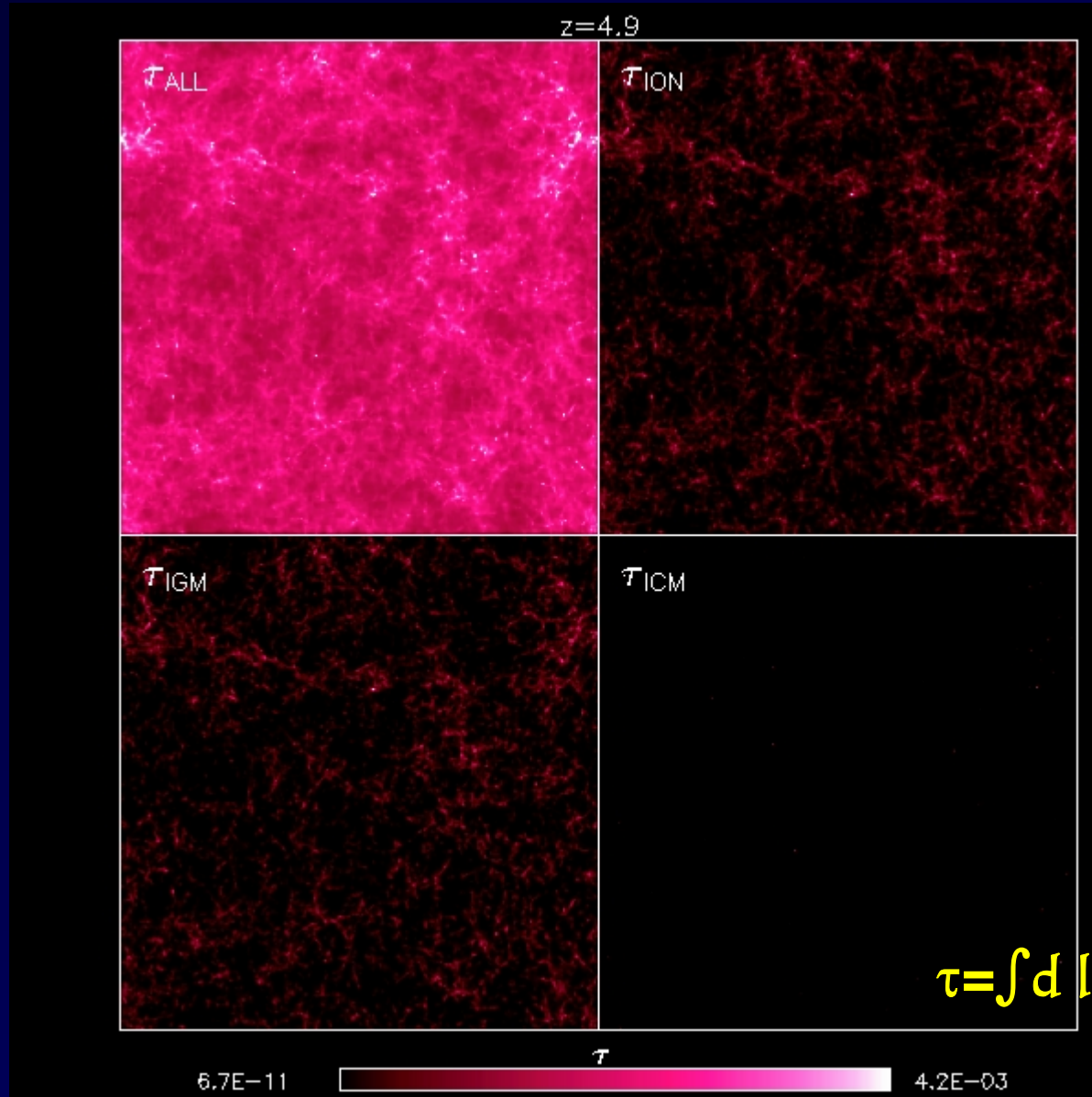
Non-rad. Hydra Code

Couchman et al. 1995

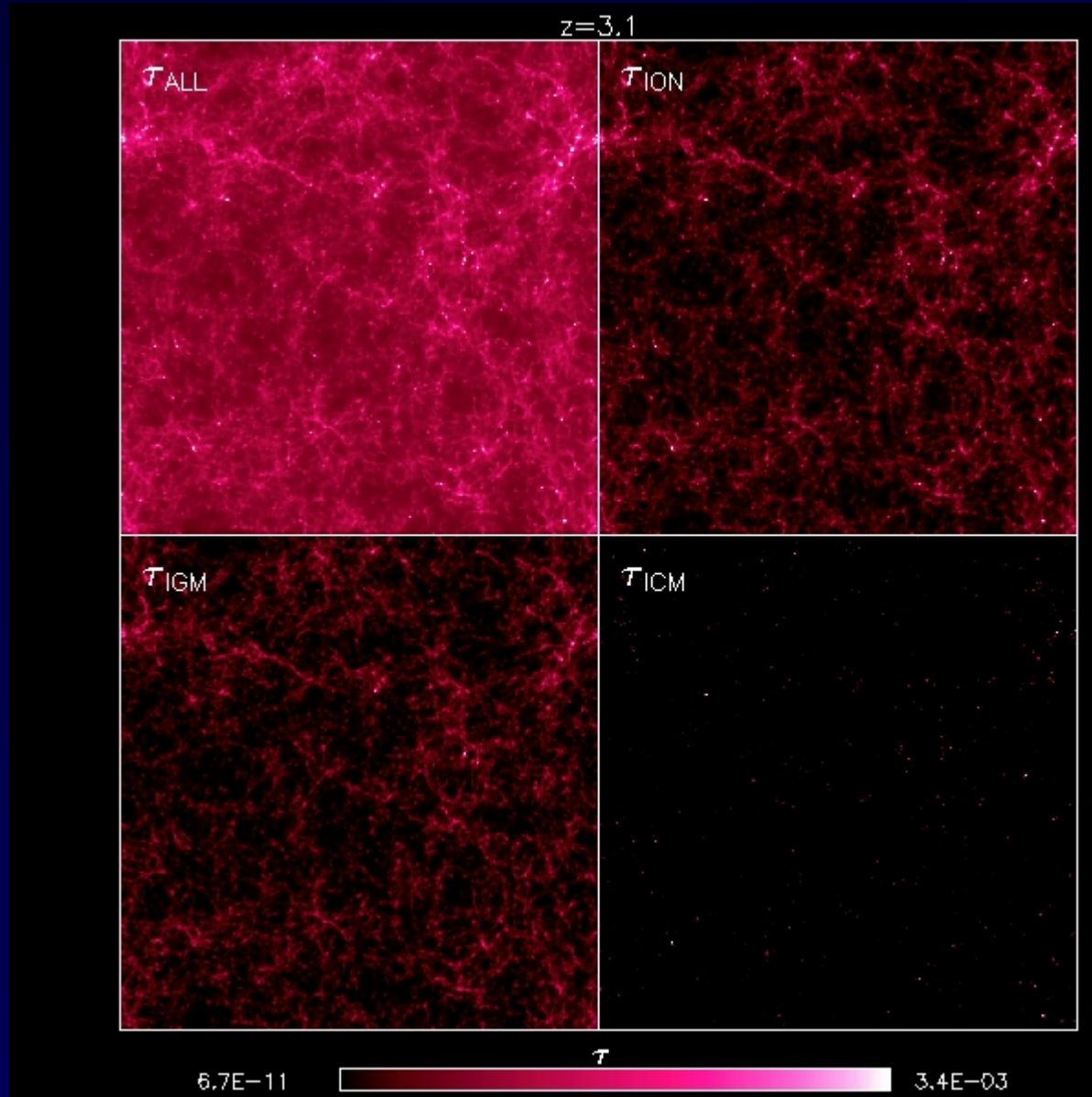
Pearce & Couchman 1997



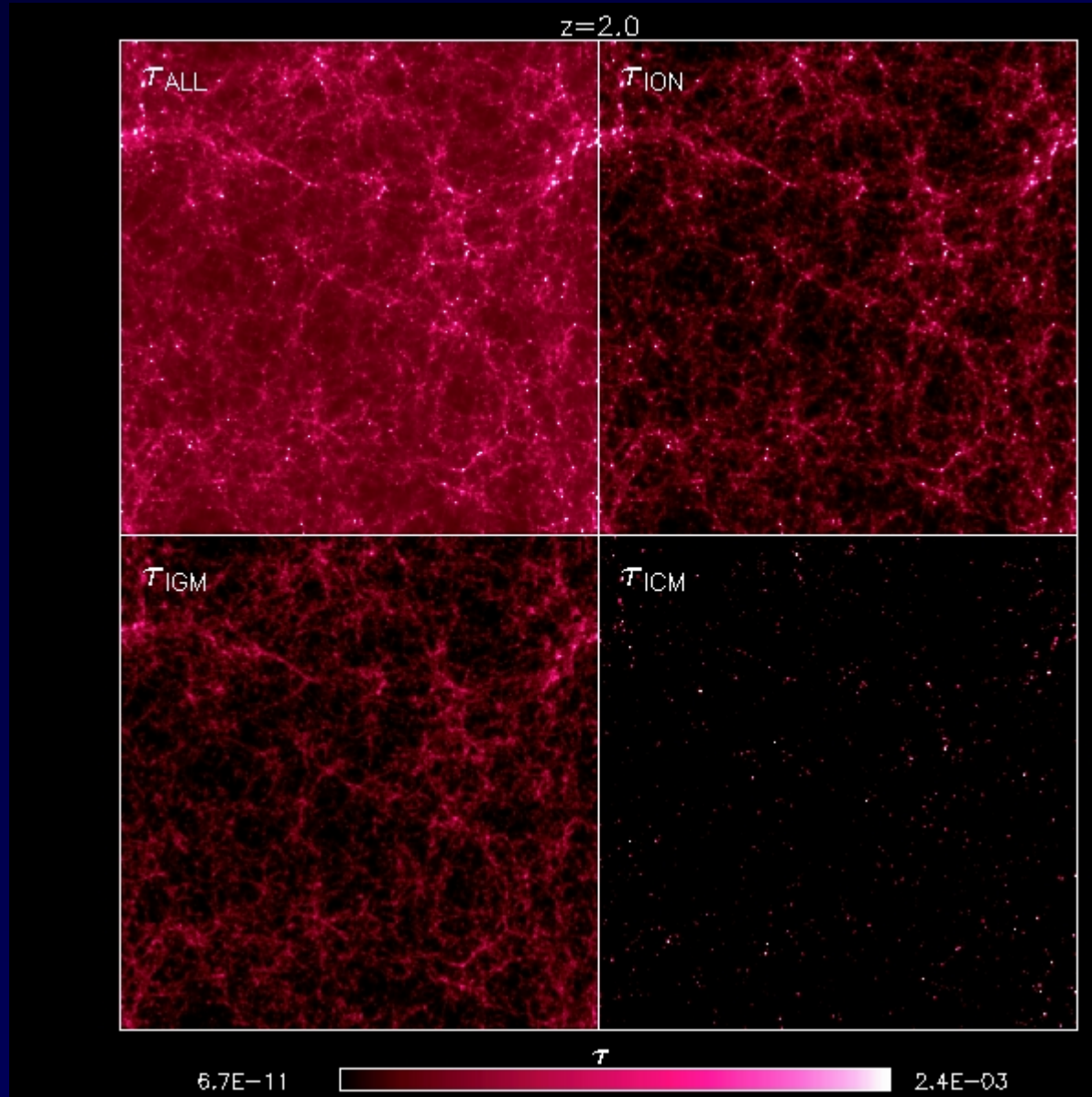
Gas Phase



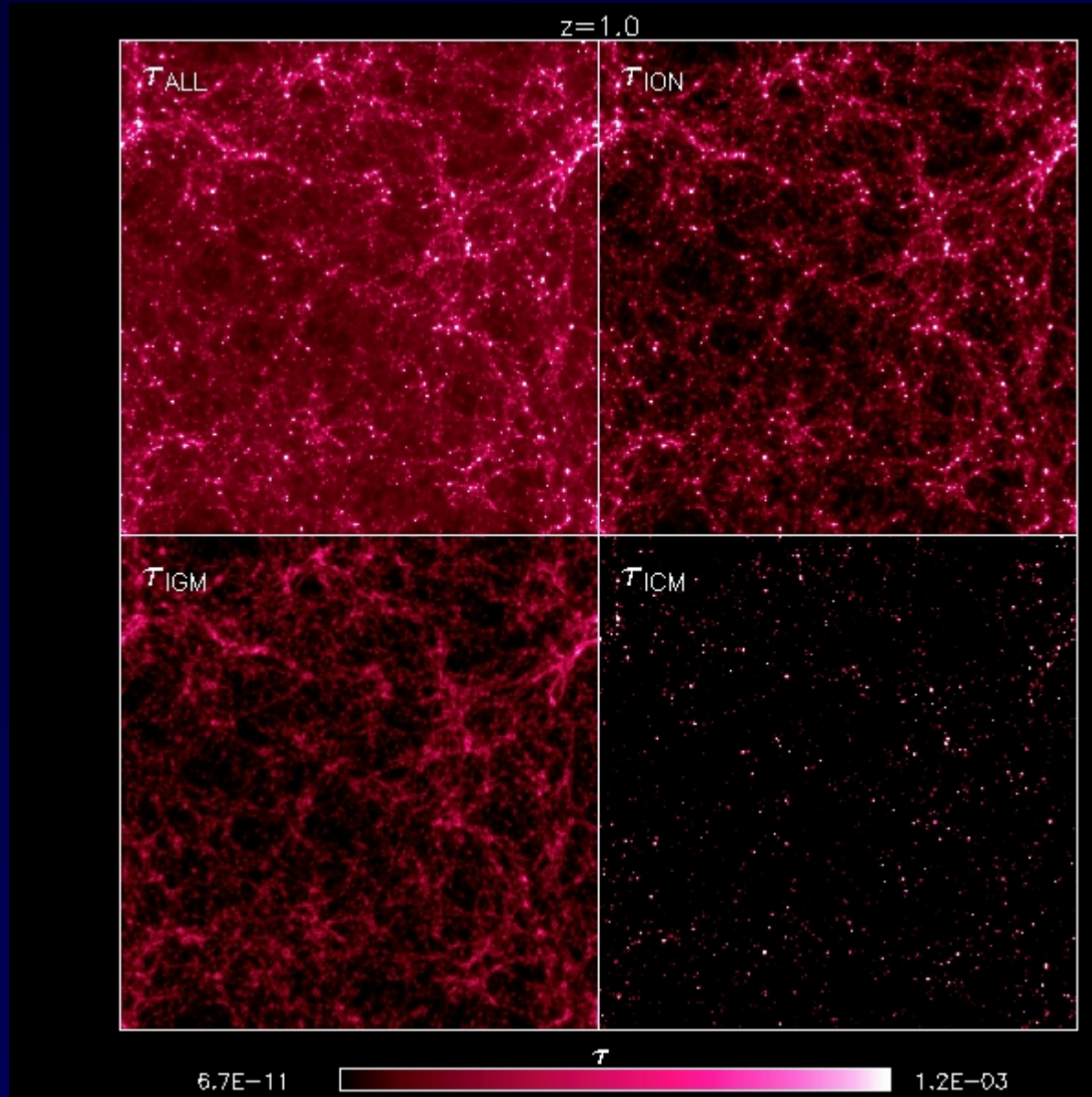
Gas Phase



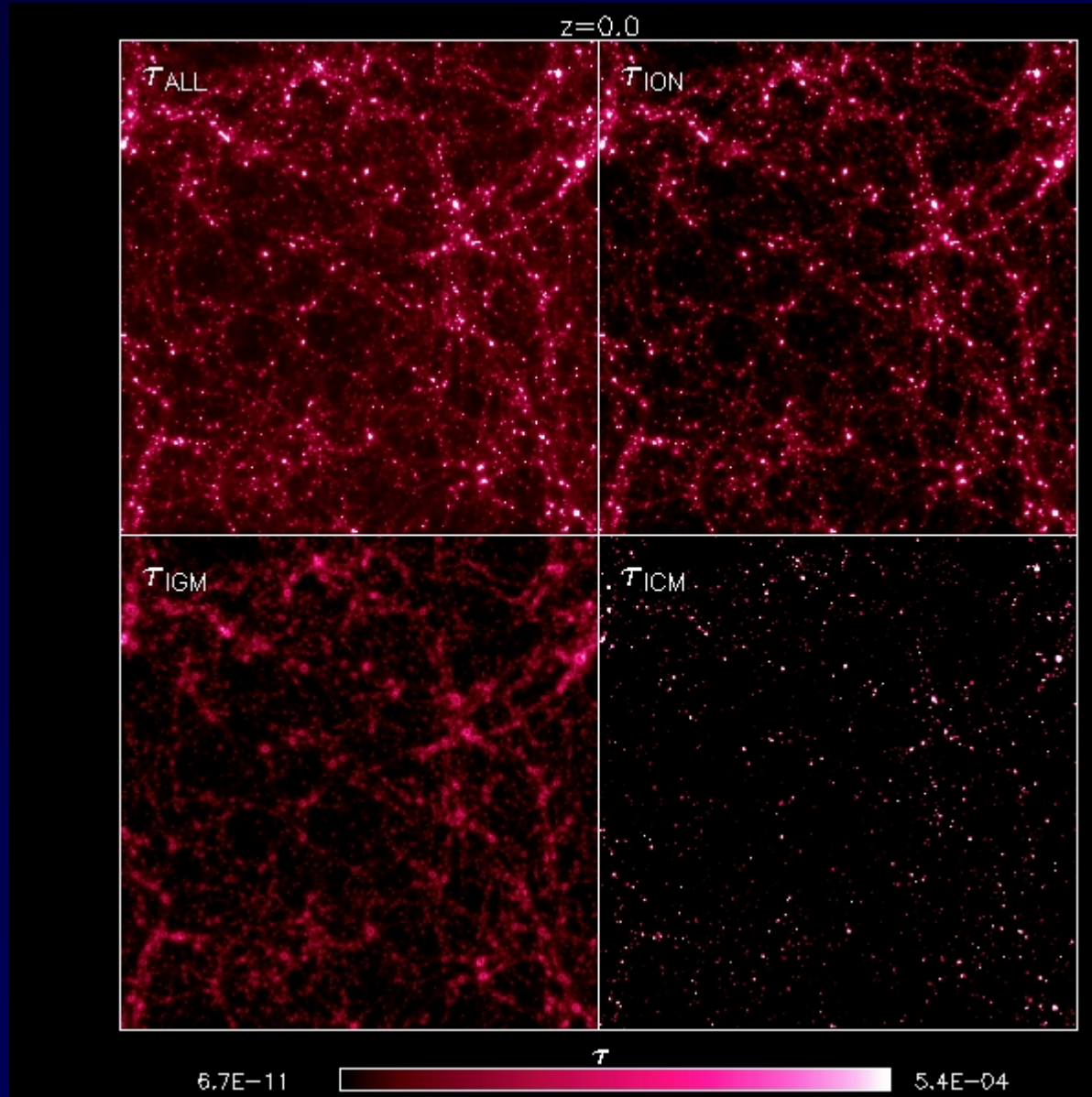
Gas Phase



Gas Phase

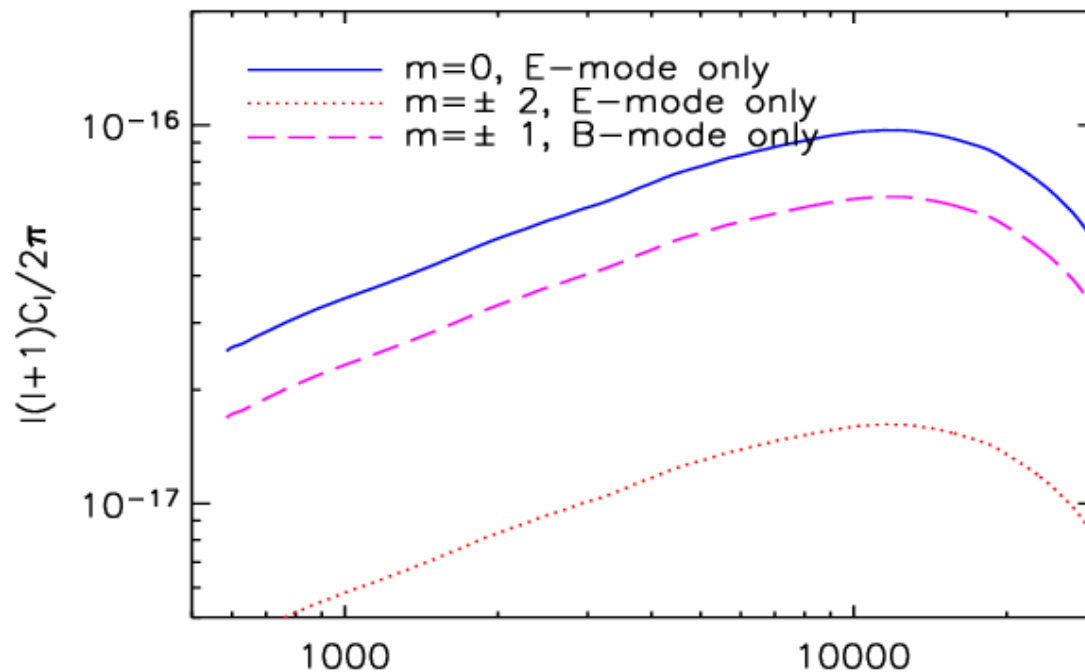


Gas Phase



$$E=B$$

	E-mode	B-mode
$m=0$	6	0
$m=\pm 1$	0	4
$m=\pm 2$	1	0



For all ionised particles

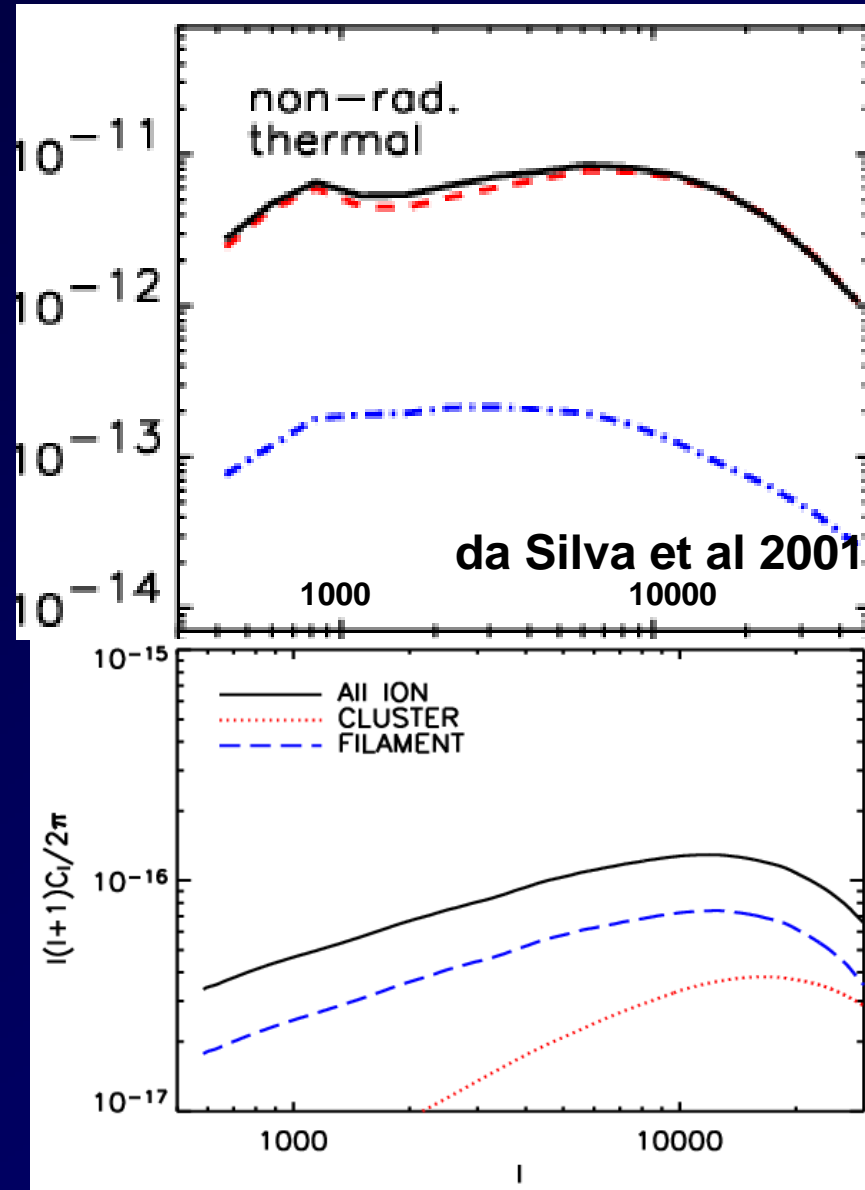
Comparing with SZ temperature

For total ionised particles
5 order of magnitudes

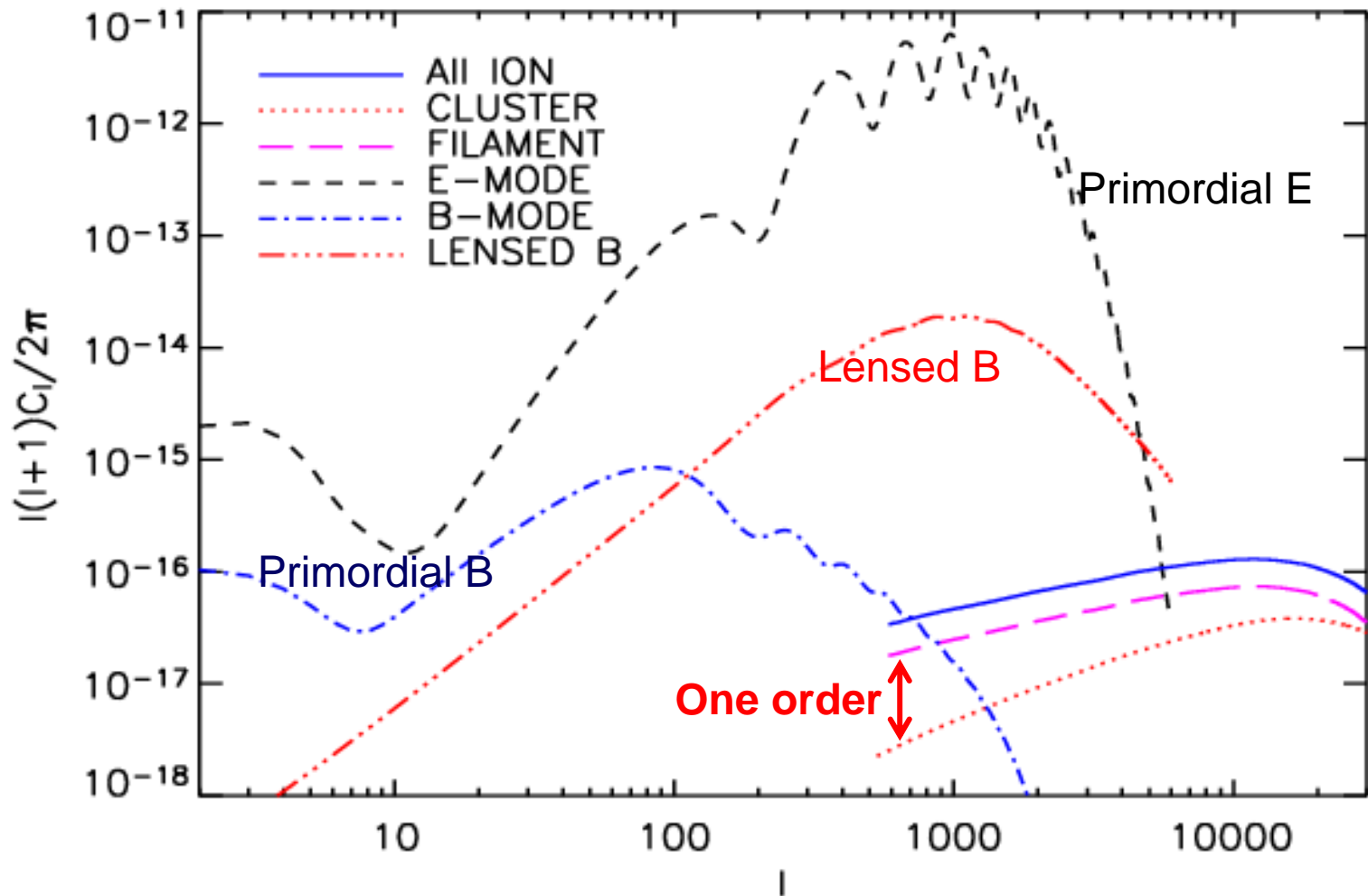
Temp. : Cluster
Polar. : Filament

$$\ell(\ell+1)C_{T_I}/2\pi$$

$$\ell(\ell+1)C_{P_I}/2\pi$$

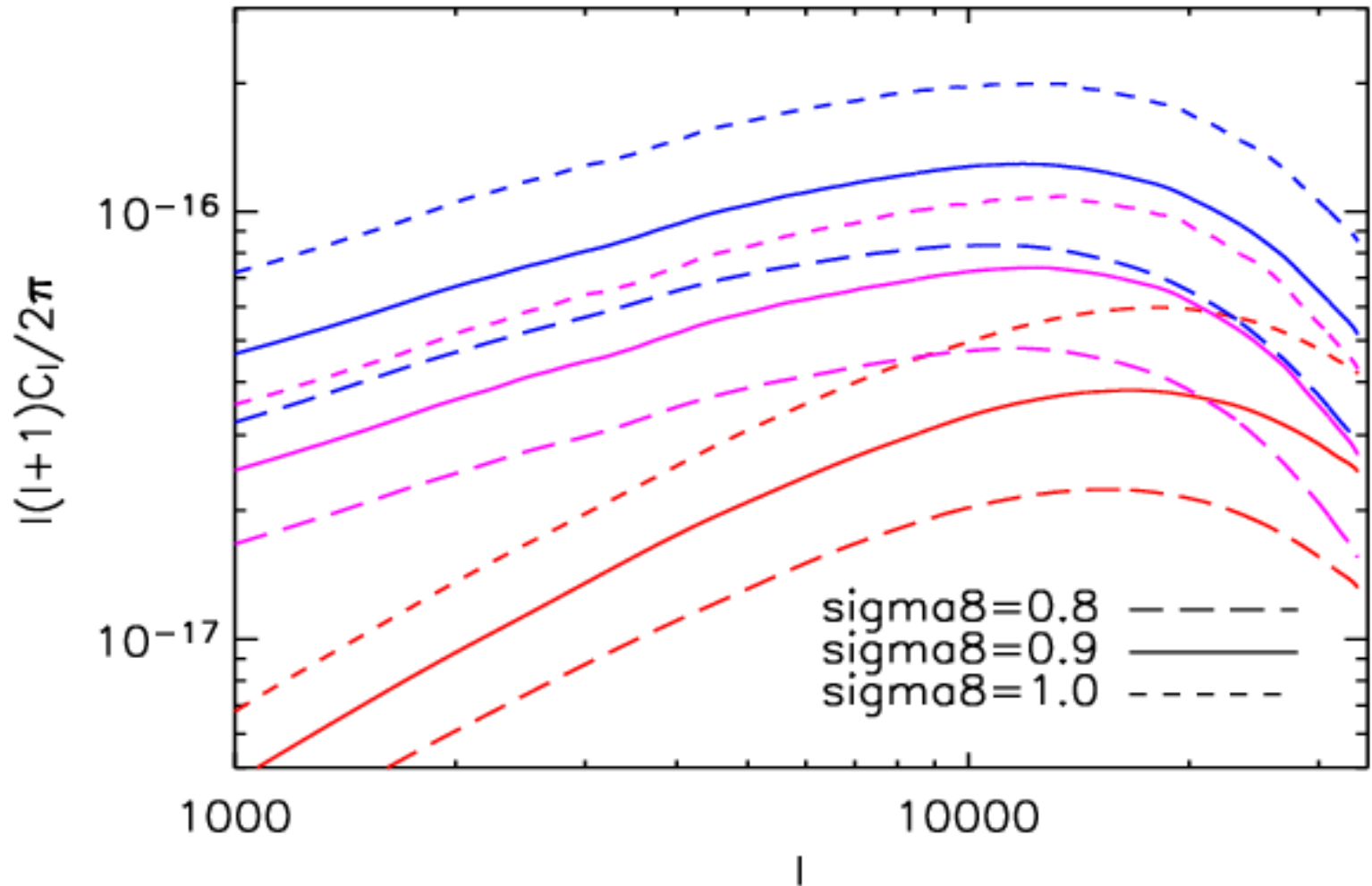


Results



σ_8

Amplitude $\propto \sigma_8^5$



Summary

1. The E-mode and B-mode have same power because the symmetry of the quadrupole in its k-space is broken by coupling with electron field.
2. The power spectrum $l(l+1)Cl/2\pi$ for all the ionising particles $\sim 10^{-15} \sim 10^{-16} K^2$
3. Secondary polarisation for cluster and filament dominate at the small scales.

Summary

4. At the intermediate scales, the B-mode power is dominated by the lensing generated power spectrum.
5. The amplitude of the secondary polarisation $\propto \sigma_8^5$
6. The power contribute from filament is much larger than the ICM different with the case of tSZ.

Why Filament Dominate?

