Evolutionary Picture of Giant Molecular Cloud Mass Functions on Galactic Scales



Masato I.N. Kobayashi (Nagoya U)

3) Successfully reproduce observed variation of GMC mass functions.



2) Formulate **Coagulation equation** of GMC Mass Function

$$\begin{aligned} \frac{\partial n_{\rm cl}}{\partial t} + \frac{\partial}{\partial m} \left(n_{\rm cl} \frac{\mathrm{d}m}{\mathrm{d}t} \right) &= -\frac{n_{\rm cl}}{T_{\rm d}} \\ &+ \frac{1}{2} \int_0^\infty \int_0^\infty K(m_1, m_2) n_{\rm cl,1} n_{\rm cl,2} \\ &\times \delta(m - m_1 - m_2) \mathrm{d}m_1 \mathrm{d}m_2 \\ &- \int_0^\infty K(m, m_2) n_{\rm cl} n_{\rm cl,2} \mathrm{d}m_2 \end{aligned}$$

4) Observations may put unique constraints on GMC formation/dispersal timescales

by observing the mass function slope.

Steady State Solution

$$n_{\rm cl}(m) = \frac{N_0}{M_{\odot}} \left(\frac{m}{M_{\odot}}\right)^{-1 - \frac{T_{\rm f}}{T_{\rm d}}}$$

The CCC effect is limited only in the massive end of GMC mass functions.